# Flood Basalts of Vestfjella: Jurassic Magmatism Across an Archaean–Proterozoic Lithospheric Boundary in Dronning Maud Land, Antarctica

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Continental flood basalts (CFBs) of Jurassic age make up the Vestfjella mountains of western Dronning Maud Land and demonstrate an Antarctic extension of the Karoo large igneous province. A detailed geochemical study of the 120-km-long Vestfjella range shows the CFB suite to consist mainly of three intercalated basaltic rock types designated CT1, CT2 and CT3 (chemical types 1, 2 and 3) that exhibit different incompatible trace element ratios. CT1 and CT2 of north Vestfjella record wide ranges of Nd and Sr isotopic compositions with initial  $\varepsilon_{Nd}$  and  $\varepsilon_{Sr}$  ranging from +7.6to -16.0 and -16 to +65, respectively. The southern Vestfjella is dominated by CT3 with near-chondritic  $\varepsilon_{Nd}$  (+2.0 to -4.1) and  $\varepsilon_{Sr}$  (-11 to +19). A volumetrically minor suite of ocean island basalt (OIB-)like CT4 dykes ( $\varepsilon_{Nd}$  + 3.6,  $\varepsilon_{Sr}$  + 1) cuts the lava sequence in north Vestfjella. The pronounced isotopic differences suggest different magmatic plumbing systems for the heterogeneous CT1 and CT2 suites and the relatively homogeneous CT3 lavas. This is further supported by the palaeoflow directions, which point to major source regions to the north (CT1 and CT2) and east (CT3) of Vestfjella. These source regions can be associated with two contemporaneous major lithospheric thinning zones that permitted magma emplacement and controlled the melting of uppermantle sources in the Jurassic Dronning Maud Land. The CT1 and CT2 magmas utilized the northern zone of thinning and were emplaced into the 3 Ga Grunehogna craton, whereas the CT3 magmas were emplaced through thinned Proterozoic Maud Belt lithosphere. Trace element and isotopic studies of the identified magma types reveal a complex history of fractionation and contamination at different lithospheric levels. All extrusive rock types show evidence of crustal contamination but this had rather small

impact on their diagnostic trace element ratios. Much stronger overprint, in the CT1 and CT2 suites, resulted from contamination with veined Archaean lithospheric mantle, which produced wide ranges of isotopic and highly incompatible element ratios. CT3, in turn, does not show evidence of interaction with the Proterozoic lithospheric mantle. The high- $\varepsilon_{Nd}$  endmembers of CT1, CT2 and CT3 probably closely resemble uncontaminated mantle-derived magmas and indicate three different mantle sources. The CT2 primary magmas were derived from light rare earth element (LREE)depleted, slightly large ion lithophile element (LILE)-enriched sources, whereas data on the volumetrically preponderant CT1 and CT3 point to variably LREE-enriched, strongly LILE-enriched sources. The sources of CT1, CT2 and CT3 may record largescale lateral heterogeneity generated by subduction-contamination of the Gondwanan upper mantle. The OIB-like CT4 dykes probably reflect asthenospheric heterogeneities that were unrelated to the proposed subduction-contamination.

KEY WORDS: Karoo magmatism; Vestfjella (Dronning Maud Land, Antarctica); Sr and Nd isotopes; magmatic differentiation; depleted and enriched mantle sources

#### INTRODUCTION

Among the fundamental questions pertaining to the generation of continental flood basalt (CFB) provinces are the importance of crustal contamination and lithospheric

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mantle sources and the relationships between rifting and magmatism. In contrast to the widely supported plume origin for the Mesozoic Karoo magmatism in Africa and Antarctica (Fig. 1) (e.g. Cox, 1989; White & McKenzie, 1989), the great majority of the Karoo basalts exhibit negative initial  $\varepsilon_{Nd}$ , positive initial  $\varepsilon_{Sr}$ , and distinct crustlike trace element characteristics-the so-called lithospheric signature (Erlank, 1984). Whereas a crustal origin has been repeatedly suggested for the lithospheric signatures of these and many other CFBs (e.g. Arndt et al., 1993; Chesley & Ruiz, 1998), the bulk of Karoo magmatism has been generally ascribed to enriched lithospheric mantle sources (e.g. Duncan et al., 1984; Hawkesworth et al., 1984; Sweeney et al., 1994). On the other hand, lithospheric mantle has also been considered as a potential contaminant of the Karoo basalts (Ellam & Cox, 1991; Ellam et al., 1992).

Another problematic issue of Karoo volcanism and the Gondwana break-up magmatism in general is the role of subduction processes, previous and contemporaneous, which could have produced enriched upper-mantle sources or otherwise influenced magma generation (e.g. Cox, 1978; Elliot, 1990; Hergt *et al.*, 1991; Brewer *et al.*, 1992; Storey *et al.*, 1992). Considerable ambiguity also surrounds the nature of plume–lithosphere interactions and the causal relationship between magmatism and continental rifting (Sweeney & Watkeys, 1990; Griffiths & Campbell, 1991; Thompson & Gibson, 1991; Cox, 1992; Saunders *et al.*, 1992; White, 1992; Grantham, 1996; Ebinger & Sleep, 1998; Leitch *et al.*, 1998).

The Jurassic CFBs of western Dronning Maud Land, Antarctica, have been viewed as a part of the Karoo large igneous province (e.g. Faure et al., 1979; Harris et al., 1990; Luttinen & Siivola, 1997) (Fig. 1). This paper is an extension of our previous work on the Vestfjella CFB suite (Furnes & Mitchell, 1978; Furnes et al., 1982, 1987; Luttinen & Siivola, 1997; Luttinen et al., 1998). Geochemical studies by Furnes and his coworkers on southern Vestfjella CFBs during the 1980s revealed a heterogeneous suite of lavas and dykes that differ from other Jurassic basalts of western Dronning Maud Land [see also Harris et al. (1990)]. On the basis of detailed sampling of two sections (Basen and Plogen; Fig. 1), Luttinen & Siivola (1997) were able to recognize three chemically different basaltic magma series in north Vestfield and designated them chemical types 1, 2 and 3 (CT1, CT2 and CT3, respectively). Geochemical comparison established an affinity to the Karoo magmatism of the Lebombo Monocline area, notably to Southern Lebombo low-Ti tholeiites by CT1 and to mid-ocean ridge basalt (MORB)-like Rooi Rand dolerites and their extrusive equivalents by CT2. Subsequently, Nd and Sr isotopic data for these rocks were reported by Luttinen et al. (1998), who also discovered a minor subgroup of dyke rocks (CT4) and suggested that the heterogeneous CFB magmatism recorded in north Vestfjella had been fundamentally linked to lithospheric mantle sources.

Here we report 210 new chemical and 27 Nd and Sr isotopical analyses of CFBs throughout the 120-km-long Vestfjella range. Combining these with the previously published data, we assess the spatial and temporal relationships between the various magma types on a regional scale. The existing data now cover the Vestfjella suite in considerable detail and enable us to better evaluate the processes that were responsible for generating an exceptionally heterogeneous CFB suite. We examine the magmatic history of each chemical type and evaluate the significance of crustal and lithospheric mantle level contamination and heterogeneous upper-mantle sources. Finally, we propose a tectono-magmatic model that links the origin of the Vestfjella CFBs to several magmatic plumbing systems and active zones of lithospheric thinning in western Dronning Maud Land during the Jurassic.

#### **GEOLOGICAL SETTING**

Mesozoic flood lavas crop out at three main localities in Dronning Maud Land: Vestfjella, Kirwanveggen and Heimefrontfjella (Fig. 1). Volumetrically, the most significant lava suites are found in Vestfjella and southern Kirwanveggen. Intrusive equivalents of the flood lavas are abundant in Vestfjella but rare elsewhere. In the Ahlmannryggen–Sverdrupfjella region, however, Mesozoic basaltic lavas are absent but dolerites fairly abundant (Harris *et al.*, 1991).

In Vestfjella, the thickness of the lava pile exceeds 900 m in the north and 400 m in the south. The lavas show an overall W-SW tilt with dips generally <10° (Furnes et al., 1987; Luttinen & Siivola, 1997). The lava pile is cut by dolerite dykes and sills and, at Muren and Utpostane, by gabbro intrusions (Hjelle & Winsnes, 1972; Räisänen et al., 1999). Normal faulting is associated with the dykes and probably was related to dyke emplacement (Spaeth, 1987). The age of the Vestfjella sequence is not accurately constrained. Available K-Ar age data for bulk rocks show large variations, with the majority of the dyke rocks yielding 160-170 Ma, and the lavas 170-230 Ma (Furnes & Mitchell, 1978; Furnes et al., 1987; Peters et al., 1991). The wide range of ages probably reflects subsolidus alteration. Plagioclase K–Ar ages of  $\sim$  180 Ma for north Vestfjella lavas (Peters et al., 1991) provide the best age estimate available and correspond to a recently obtained <sup>39</sup>Ar/<sup>40</sup>Ar age of 183  $\pm$  1 Ma for basalts in Kirwanveggen and southeast Africa (Duncan et al., 1997).

The base of the lava pile is not exposed in Vestfjella. Elsewhere, the Dronning Maud Land lavas were erupted on a Palaeozoic continental sedimentary sequence overlying a Precambrian basement (Aucamp *et al.*, 1972;

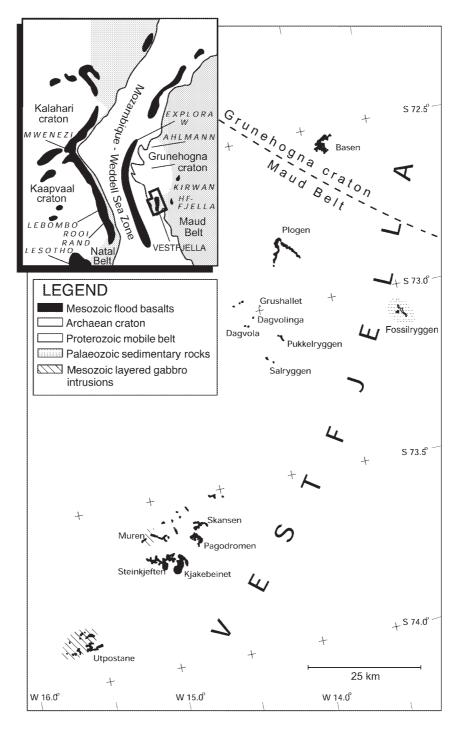


Fig. 1. Distribution of Karoo and related flood basalts and the main lithospheric domains in a schematic Mesozoic Gondwana reconstruction of Africa and Antarctica (inset) and distribution of Mesozoic flood basalts and related gabbro intrusions in Vestfjella, western Dronning Maud Land, Antarctica. Abbreviations for Kirwanveggen, Heimefrontfjella and Ahlmannryggen are Kirwan, HF-Fjella and Ahlmann, respectively. Gondwana reconstruction and Mozambique–Weddell Sea lithospheric thinning zone after Cox (1992).

Hjelle & Winsnes, 1972). The basement is divided into two major domains. Before Mesozoic break-up, the Archaean Grunehogna craton (Halpern, 1970) was probably part of the Kaapvaal craton (Groenewald *et al.*, 1995; Fig. 1). The craton is bounded to the east and southeast by the Mesoproterozoic Maud Belt, the Antarctic extension of the Natal Belt of Africa (e.g. Jacobs *et al.*, 1993) (Fig. 1). The exact position of the Archaean–Proterozoic lithospheric terrane boundary is not firmly established, but, on the basis of gravity and aeromagnetic data, it has been interpreted to be located within  $72^{\circ}$ – $73^{\circ}$ S, i.e. close to north Vestfjella (Corner, 1994; Fig. 1).

Volumetrically, an important part of the Karoo magmatism is represented by the opposing sequences of volcanic rocks of the Lebombo Monocline and the western part of the Explora Wedge (Cox, 1992; Fig. 1). It has been proposed that the generation of these basalt suites was related to development of a major lithospheric thinning zone and the initial opening stage of the Southwest Indian Ocean (e.g. Marsh, 1987; Cox, 1992; Duncan *et al.*, 1997). The southward extension of this so-called Mozambique lithospheric thinning zone (Cox, 1992) followed the present Weddell Sea margin of Dronning Maud Land and could have been petrogenetically linked to the seaward dipping basalt succession of Vestfjella (Fig. 1).

## DATASET AND ANALYTICAL METHODS

The 210 samples reported in this study have been collected by the authors during five Norwegian and Finnish Antarctic expeditions during the period 1985–1998 from the localities indicated in Fig. 1. The analysed basalts include one sill rock sample (P27-AVL). The samples were analysed by X-ray fluorescence spectrometry (XRF) for major and trace elements at the Geoanalytical Laboratory, Washington State University (120 samples) and the University of Bergen (90 samples). The analytical procedures used at the Geoanalytical Laboratory are summarized below. For the XRF procedure used in the University of Bergen, the reader is referred to Furnes *et al.* (1996).

The major and minor oxides and Ni, Cr, Sc, V, Ba, Zr, Y, Cu and Zn were analysed by XRF (Rigaku 3370) using the procedure described by Johnson *et al.* (1999). A set of 26 duplicate samples were first analysed at Bergen and then at the Geoanalytical Laboratory to evaluate interlaboratory bias (Appendix, available on the *Journal of Petrology* web site at http://www.petrology. oupjournals.org). Rock chips were pulverized in a tungsten carbide swing mill. The duplicate samples (labelled VF in Table 1) were ground in an agate chamber. The reported precision of analyses is better than 2% for most oxides and trace elements, whereas that for Ni, Cr, Sc and V is better than 10% and that of Cu is better than 18% (Johnson *et al.*, 1999).

A subset of representative samples for isotopic and high-resolution trace element analyses was selected on the basis of the XRF data. Inductively coupled plasma mass spectrometry (ICP-MS) was used for analysis of the lanthanides, Rb, Sr, Nb, Hf, Ta, U, Pb and Th at the Geoanalytical Laboratory. The analytical procedure has been described in detail by Knaack *et al.* (1994). Chips of crushed samples were ground in an iron bowl using a shatter-box swing mill, apart from samples labelled VF (see Table 1), which were pulverized in an agate chamber. The statistics for a single sample (BCR-P) run over a 4 month period indicate that precision for the lanthanides, Rb, Sr, Nb and Hf is better than 2.5%. Precision for Ta and Pb is better than 3.5%, and for Th and U it is better than 10% (Knaack *et al.*, 1994).

Strontium and neodymium isotopes were analysed mainly on a Finnigan 262 mass spectrometer at the University of Bergen. All chemical processing was carried out in a clean-room environment with high efficiency particulate air (HEPA) filtered air supply and positive pressure, and the reagents were purified in two-bottle Teflon stills. Samples were dissolved in a mixture of HF and HNO<sub>3</sub>. Strontium was separated by specific extraction chromatography using the method described by Pin *et al.* (1994) and was loaded on a double filament and analysed in static mode. The Sr isotopic ratios were corrected for mass fractionation using a <sup>88</sup>Sr/<sup>86</sup>Sr value of 8·375209. Repeated measurements of the NBS 987 Sr standard yielded an average <sup>87</sup>Sr/<sup>86</sup>Sr value of 0·710251 ± 19 (2 $\sigma$ ) (n = 27).

Rare earth elements (REE) were separated by specific extraction chromatography using the method described by Pin *et al.* (1994). Neodymium was subsequently separated using low-pressure ion-exchange chromatography with di(2-ethylhexyl)orthophosphoric acid (HDEHP)-coated Teflon powder as the ion-exchange resin (Richard *et al.*, 1976) and then loaded on a double filament and analysed in static mode. Nd isotopic ratios were corrected for mass fractionation using a <sup>146</sup>Nd/<sup>144</sup>Nd value of 0.7219. Repeated measurements of the JM Nd-standard yielded an average <sup>143</sup>Nd/<sup>144</sup>Nd value of 0.511113  $\pm$  15 (2 $\sigma$ ) (n = 62).

Two of the samples (P27-AVL and P55-AVL) were analysed for Nd, Sm, Rb and Sr concentrations and isotopic ratios at the Geological Survey of Finland. For analytical methods, the reader is referred to Luttinen *et al.* (1998).

### PETROGRAPHY

The chemically analysed samples are typical tholeiitic CFBs. Most of the rocks are porphyritic (83%). The predominant phenocryst assemblages in the order of appearance are: plagioclase (29% of the porphyritic samples), plagioclase + augite (28%), plagioclase + olivine (12%) and plagioclase + augite + olivine (11%).

Mole and more detarements were, were, were and and an analysis of a serie and analysis of a serie and analysis and a serie a	Sample: Type:* Location:	VF57- 85 CT1 B	P11- AVL CT1 P	P2- AVL CT1 P	P23- AVL CT1 P	P40- AVL CT1 P	P42- AVL CT1 P	P59- AVL CT1 P	P69- AVL CT1 P	VF76- V 85 85 8 CT1 0 P F	VF80- U 85 A CT1 C P U	UP206- SK AVL AV CT1 CT UP SK	SK224- P37- AVL AVL CT1 CT1 SK P	r- P44- - AVL 1 CT1 P	P45- AVL CT1 P	VF147- 85 CT1 KB	VF26- 85 CT1 PR	VF30- V 85 85 85 8 CT1 0 PR 8	VF40- V 85 85 CT1 0 SR 8	VF44- V 85 85 8 CT1 0 SR S	VF46- V 85 85 8 CT1 C SR P	VF81- PE 85 A\ CT1 C1 PD P	P55- B33- AVL AVL CT2 CT2 P B	3- B67- L AVL 2 CT2 B	- P18- - AVL 2 CT2 P
65.2         55.1         56.1         55.2         56.4         56.4         56.7         55.7         55.4         56.4         56.4         56.7         55.7         55.1         55.1         55.1         56.1         55.2         55.1         56.1         55.2         56.1         56.3 <th< td=""><td>Major and</td><td>minor ele</td><td>ements (</td><td>'wt %)†</td><td></td><td></td><td></td><td></td><td></td><td></td><td></td><td></td><td></td><td></td><td></td><td></td><td></td><td></td><td></td><td></td><td></td><td></td><td></td><td></td><td></td></th<>	Major and	minor ele	ements (	'wt %)†																					
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142         161         151         163         153         153         153         153         153         153         153         154         164         105         173         154         165         173         153         154         164         107         017 <td><math>TIO_2</math></td> <td>1.36</td> <td>3 1.13</td> <td>0.89</td> <td></td> <td>1.03</td> <td>1.26</td> <td></td> <td></td> <td>1.24</td> <td></td> <td></td> <td></td> <td></td> <td></td> <td></td> <td>1.56</td> <td>1-46</td> <td>1.45</td> <td>1-57</td> <td>1.51</td> <td></td> <td></td> <td></td> <td></td>	$TIO_2$	1.36	3 1.13	0.89		1.03	1.26			1.24							1.56	1-46	1.45	1-57	1.51				
0.5         0.4         0.5 <td><math>AI_2O_3</math></td> <td>14-97</td> <td>14.92</td> <td>12.15</td> <td></td> <td>16.87</td> <td>15.12</td> <td></td> <td></td> <td>14.48</td> <td></td> <td></td> <td></td> <td></td> <td></td> <td></td> <td>15.08</td> <td>15-58</td> <td>14.73</td> <td>14.86</td> <td>14.85</td> <td></td> <td></td> <td></td> <td></td>	$AI_2O_3$	14-97	14.92	12.15		16.87	15.12			14.48							15.08	15-58	14.73	14.86	14.85				
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197         331         246         204         320         216         212         320         216         212         320         216         217         030         55         316         317         320         230         230         230         230         230         230         230         230         230         301         017         019         025         014         016         025         014         016         025         016         056         056         056         056         056         057         054         055         056	CaO	10.50	7.18	10.88		9.11	9-83			8.79							10-83	7.58	10.69	10.68	10.18				
1:0         1:15         0:43         1:37         1:20         0:18         1:31         0:43         0:53         0:54         0:54         0:54         0:54         0:54         0:55         0:54         0:54         0:55	Na <sub>2</sub> O	2.58	3.02	1.50		3.31	2.45			2.22							2.58	3.56	2.16	2.39	2.39				2.49
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3:31         2:62         n.d.         3:23         2:76         0:17         2:16         0:53         0:53         0:53         0:53         0:53         0:53         0:54         0:54         0:54         0:54         0:54         0:54         0:54         0:54         0:55         0:55         0:55         0:55         0:55         0:51         0:55         0:51         0:55         0:51         0:54         0:54         0:54         0:54         0:55         0:55         0:55         0:55         0:55         0:55         0:55         0:55         0:56         0:51         0:57         0:56         0:51         0:57         0:56         0:51         0:55         0:56         0:51         0:55         0:56         0:57         0:56         0:57         0:56         0:57         0:56 <th< td=""><td><math>P_2O_5</math></td><td>0.27</td><td>0.15</td><td>0.14</td><td></td><td></td><td>0.23</td><td></td><td></td><td>0.15</td><td></td><td></td><td></td><td></td><td></td><td></td><td>0.23</td><td>0.24</td><td>0.22</td><td>0.26</td><td>0.25</td><td></td><td></td><td></td><td></td></th<>	$P_2O_5$	0.27	0.15	0.14			0.23			0.15							0.23	0.24	0.22	0.26	0.25				
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selements (ppm)           27         2         1         25         21         23	<i>mg</i> -no.	0-56	9-00	0.73			0-58			0.55							0.54	0-55	0.54	0.54	0.52				
27         29         16         28         31         34         29         31         31         29         28         31         31         29         31         31         29         31         32         31         31         32         31         32         31         32         31         32         31         32         32         32         32         32         32         32         32         32         32         32         32         32         32<	Trace elem	ents (ppn	(11																						
277         246         25         74         235         273         235         288         291         293         271         307         276         286         283         281	Sc	28	27		16				23	29		34 2		1 31		29	34	28				27 2	9 2	8 29	
234         320         627         436         536         137         137         139         137         139         137         139         137         139         137         130         136         137         136         137         136         137         137         139         130         136         137         137         130         130         131         130         137         139         130         130         137         139         130         130         131         306         137         137         139         130         131         306         137         137         139         131         306         131         137         131         306         131         136         131         131         301         311         306         131         311         306         131         311         306         131         311         301         311         301 <td>&gt;</td> <td>277</td> <td>246</td> <td></td> <td>214</td> <td></td> <td></td> <td></td> <td>243</td> <td>262</td> <td></td> <td>25 26</td> <td></td> <td>18 291</td> <td></td> <td>271</td> <td>307</td> <td>276</td> <td></td> <td></td> <td></td> <td>261 37</td> <td>8 34</td> <td>9 43(</td> <td></td>	>	277	246		214				243	262		25 26		18 291		271	307	276				261 37	8 34	9 43(	
65         96         326         143         102         84         111         107         45         92         67         44         36         37         43         34         45         34         22         32         43         34         35         34         55         34         55         31         85         77         73         88         70         73         88         70         73         88         70         73         88         70         73         88         70         73         88         70         73         88         70         73         88         70         73         88         70         73         88         70         73         88         70         73         88         70         73         88         70         73         88         70         73         88         70         73         88         70         73         73         88         70         73         73         73         88         70         88         70         88         703         88         703         88         703         88         70         73         88         70         73       <	Ċ	234	320		435				138	137		58 19		2 184		166	191	195				198 30	15 7	9 119	
95         86         80         95         86         95         86         95         86         95         77         73         90         101         99         87           165         64.6         73         73         73         74         16.1         16.1         16.1         17.1         73         73         73         73         73         73         73         73         73         73         74         16.1         16.3         17.3         75         73         73         73         73         73         73         73         74         15.2         17.1         74         15.3         13.4         15.3         13.4         15.4         15.4         15.4         15.3         13.4         15.4         13.4         15.4         13.3	ïZ	65	96		143				107	45		57 4		6 32		43	34	45				43 11	0 7	7 7,	
95         77         78         74         122         113         92         104         90         103         67         96         113         75         91         85         77         73         89         70         73           165         64.6         23         387         940         673         296         47         371         344         550         65         25         26         23         379         554         371         306         405         57         26         413         306         407         57         55         25         26         24         28         30         55         56         45         313         136         131         306         57         56         313         314         106         573         338         146         15.4         146         15.4         166         146         57         30         165         56         26<	Zn	95	86		90				86	68		95 8		100 100		84	96	94				88	11 60	8 144	
	Cu	95	77		74				92	104		9 60		113		91	85	77				73 10	11 11	2 195	
320         306         81         337         940         673         299         447         377         344         260         450         353         351         364         353         379         264         287         311         306         422         25         20         21         23         23         24         129         24         25         21         23         23         24         129         26         70         604         575         26         270         604         571         311         421         327         320         465         707         604         577         311         421         327         320         146         333         143         120         143         327         320         146         575         520         120         121         121         121         121         121         121         123         123         133         146         1436         1435         1453         1359         146         1357         1329         1316         132         132         132         132         132         132         132         132         132         132         132         132	Rb	16-5	64-6		38.7				37.9	54.4		2.6 2		5.6 13.4		22.6	6.3	14.6				41.9 1	8.6 3	6.1	
29         24         20         21         18         24         19         25         103         18         24         19         25         103         133         137         135         137         135         137         160         78         58         515         664         419         515         610         513         610         614         137         137         131         133         373         311         471         377         160         187         160         142         152         103         337         311         421         373         311         421         373         311         421         373         311         421         373         311         421         161         161         163	Sr.	320	306		337				497	327		50 46		15 361		353	379	254				402 45	51 51	0 269	
137         109         78         98         113         122         80         155         135         133         133         134         120         135         127         165         142         132           6.05         5.93         151         5.53         5.11         5.53         5.26         7.07         6.04         5.75           6.06         5.93         151         5.33         3.31         387         173         812         10.53         5.66         7.42         15.73         316         145           5.06         715         5146         1547         18.61         15.73         13.88         15.73         13.88         15.73         389         916         12.52         10.3         5.65         5.64         2.64         2.64         2.64         2.64         2.64         2.64         2.64         2.64         2.64         2.64         2.64         2.64         2.64         14.8         14.8         14.8         16.6         14.8         14.8         16.6         14.8         14.8         16.4         14.4         16.4         14.4         16.4         14.4         16.4         14.4         16.4         16.4         16.3 <td>~</td> <td>29</td> <td>24</td> <td></td> <td>21</td> <td></td> <td></td> <td></td> <td>22</td> <td>27</td> <td></td> <td>28 2</td> <td></td> <td>5 26</td> <td></td> <td>24</td> <td>28</td> <td>26</td> <td></td> <td></td> <td></td> <td>26 2</td> <td>3</td> <td>35 5(</td> <td>50 35</td>	~	29	24		21				22	27		28 2		5 26		24	28	26				26 2	3	35 5(	50 35
8.05         5.93         4.11         6.56         4.19         6.10         5.51         5.50         5.10         6.42         5.12         6.33         5.85         5.26         7.07         6.04         5.73           508         780         153         153         153         153         153         153         153         153         153         153         163         163         1463           20.43         15-13         15-24         1322         1722         153         11-32         13-13         163         16-13         143         1463           20.43         15-1         15-1         15-2         15-3         15-3         15-3         15-3         15-3         16-3         14-3         143           20.43         15-1         15-1         15-1         15-4         15-2         10-6         14-3         14-3           20.41         22-8         15-3         15-3         15-4         15-4         15-4         15-4         15-4         15-4         15-4         15-4         15-4         15-4         15-4         15-7         15-1         15-1         15-1         15-1         15-1         15-1         15-1         1	Zr	137	109		98				105	135		53 13		6 133		120	131	135				132 13	39 15	5 19,	
508         780         29         782         501         986         79         533         531         132         1102         369         773         772         311         421         327         327         330         1463           20.43         13-94         15-73         13-98         15-73         13-98         15-73         19-64         14-67         10-67         14-20         12-35         16-87         13-63         15-81         13-96         12-65         25-62         25-24         25-31         16-14         16-47         16-167         14-20         12-35         16-14         14-61         16-167         16-167         13-16         16-14         16-147         16-167         16-167         16-167         16-1	Nb	8.05	5.93		6.59				6.19	5.51		5.10		6.02 6.4(		5.12	6.03	5.85				5.75	6.73	6.81	
20-43         15-34 <td< td=""><td>Ba</td><td>508</td><td>780</td><td></td><td>762</td><td></td><td></td><td></td><td>533</td><td>531</td><td></td><td>73 82</td><td>•</td><td>369</td><td></td><td>772</td><td>311</td><td>421</td><td></td><td></td><td>-</td><td>463 17</td><td>4 29</td><td>5 15</td><td></td></td<>	Ba	508	780		762				533	531		73 82	•	369		772	311	421			-	463 17	4 29	5 15	
41/16       27.72       31.41       34.42       18.81       30.33       28.67       27.205       26.46       23.01       26.09       25.40       21.62       26.39       25.66       146       14.91       15.67       13.79       15.63       15.61       14.94       18.76       16.41       16.41       16.41         22.18       14.49       22.83       13.65       3.76       4.77       4.16       4.38       4.65       4.69       4.46       4.92       4.89       4.65       4.64       4.93       4.94       157       1.66       1.67       1.93 <td>La</td> <td>20.43</td> <td>3 13.94</td> <td></td> <td>15.34</td> <td></td> <td></td> <td></td> <td>15.73</td> <td>13.89</td> <td></td> <td>9.16 1</td> <td></td> <td>0.62 12.3</td> <td></td> <td>9.96</td> <td>12.05</td> <td>12.26</td> <td></td> <td></td> <td></td> <td>12.33 1</td> <td>4.91 1</td> <td>7.28 17</td> <td></td>	La	20.43	3 13.94		15.34				15.73	13.89		9.16 1		0.62 12.3		9.96	12.05	12.26				12.33 1	4.91 1	7.28 17	
22.18       14.49       22.58       14.57       18.04       10.46       15.05       14.64       14.45       16.43       14.28       15.28       13.75       14.57       14.94       18.76       16.41       16.41       16.43         5.05       4.11       7.25       3.17       4.06       4.79       4.16       4.38       4.77       4.28       4.88       4.65       4.65       4.42       5.95       1.64       1.64       1.45         1.79       1.25       1.36       3.15       1.11       1.51       1.51       1.53       1.64       1.45       1.65       1.64       1.465       1.65       1.64       1.455       1.64       1.65       1.65       1.64       1.65       1.65       1.64       1.65       1.65       1.64       1.64       1.65       1.65       1.64       1.65       1.65       1.64       1.66	e ce	41./t	21.12		21.94				30.33	/9.97		7 96-07		3-01 20-0		71.02	20.39	26-30				20.24 3	5 0C-12	0.38 4(	
0.00       4.11       1.21       4.00       4.71       4.10       4.20       5.20       5.20       5.20       5.20       5.20       5.20       5.20       5.20       5.20       5.20       5.20       5.20	PN -	22.15	3 14.49		13.25				15.02	16-90		14.45 1		4.28 16.2		13.79	16.35	16.18				16.14 2	1.52 2	0.64 26	
1/3       1/2       1/3       1	0	20.0			00.0				00.4 7	6/. <del>1</del>		• • •		1 50 4:0		0.4 1	107	50.4 10.4				20.4			
0.04       4.34       0.11       0.05       0.11       0.05       0.11       0.05       0.11       0.05       0.11       0.05       0.11       0.05       0.11       0.05       0.11       0.05       0.11       0.05       0.11       0.05       0.11       0.05       0.11       0.15       0.10       0.65       0.11       0.15       0.16       0.11       0.15       0.10       0.65       0.11       2.25       2.11       2.16       2.16       2.16       2.16       2.16       2.14       2.19       2.18       0.87       0.88       0.88       0.88       0.89       0.89       0.89       0.89       0.89       0.89       0.81       0.31	n 77	2/-1	VC V		77 C0 c					70.1		20.1		1.1 20.1		10.1	0/	+ 0.1				1100	00.1	2.03 2.03	
2.24       1.93       3.29       1.71       1.55       1.96       1.67       1.93       2.41       2.16       2.44       2.19       2.18         0.34       0.31       0.47       0.26       0.53       0.34       0.34       0.34       0.34       0.34       0.34       0.34       0.34       0.34       0.34       0.35	D 4	0.91	0.78		0.68				0.71	0.94		+0 		0.83 0.86		0.80	0.89	0.86				0.87	68.0	1.21	
0.34       0.31       0.47       0.26       0.23       0.36       0.31       0.33       0.34       0.35       0.34       0.35       0.34       0.35       0.34       0.35       0.34       0.35       0.34       0.35       0.34       0.35       0.34       0.35       0.34       0.35       0.34       0.35       0.34       0.35       0.34       0.35       0.34       0.35       0.34       0.35       0.34       0.35       0.34       0.44       0.35	Υb	2.24	1.98		1.71				1.93	2.41		2.25		2.07 2.2		1.90	2.26	2.16				2.18	1.87	2.59	
3.53       2.75       2.00       2.38       2.26       2.81       1.97       2.66       3.79       3.80       3.37       3.30       3.52       3.34       3.13       3.50       3.56       3.36       4.27       3.65       3.46       3.78       3.45       3.46       3.73       0.36       0.34       0.35       0.37       0.37       0.36       0.37       0.37       0.37       0.36       0.36       0.36       0.36       0.36       0.36       0.36       0.36       0.36       0.36       0.37       0.34       0.37       0.34       0.36       0.37       0.36       0.47       0.47	Lu	0.34	1 0.31		0.26				0.29	0.36		0.33		0.32 0.3;		0.29	0.34	0.33				0.33	0.26	0.37 (	
0.44 0.35 0.24 0.37 0.25 0.32 0.24 0.34 0.36 0.35 0.26 0.35 0.33 0.36 0.34 0.35 0.37 0.34 0.42 0.37 0.34 0.36 0.37 0.34 0.42 0.37 0.36 0.49 1.60 1.72 1.01 2.66 0.88 1.13 1.01 1.63 1.83 1.72 0.38 0.48 0.47 0.51 0.44 0.54 0.53 0.47 0.73 0.59 0.49 4.27 2.65 3.11 4.00 3.38 3.46 2.72 3.87 4.67 3.15 2.19 3.59 4.31 2.78 3.40 2.61 4.03 3.07 3.36 4.74 4.19 3.26 0.31 0.38 0.41 0.28 0.41 0.29 0.61 0.19 0.25 0.20 0.33 0.47 0.11 0.11 0.11 0.17 0.19 0.16 0.11 0.19 0.16 0.11 0.11 0.11 0.11 0.11 0.11 0.11	Ηf	3-53	3 2.75		2.38				2.66	3.79		3.80		3.30 3.5.		3.13	3.50	3.56				3.48	3.83	4.16	
1-60 1-72 1-01 2-66 0-88 1-13 1-01 1-63 1-83 1-72 0-38 0-48 0-47 0-51 0-44 0-54 0-53 0-47 0-73 0-59 0-49 4-27 2-65 3-11 4-00 3-38 3-46 2-72 3-87 4-67 3-15 2-19 3-59 4-31 2-78 3-40 2-61 4-03 3-07 3-36 4-74 4-19 3-26 0-38 0-41 0-11 0-17 0-11 0-19 0-26 0-20 0-18 0-17 0-18 0-17 0-19 0-16 0-19 0-21 0-18 0-18 0-17 0-19 0-16 0-19 0-21 0-19 0-21 0-18 0-19 0-18 0-19 0-19 0-18 0-18 0-19 0-18 0-19 0-18 0-18 0-19 0-18 0-18 0-18 0-18 0-19 0-21 0-19 0-21 0-21 0-21 0-18 0-18 0-19 0-18 0-19 0-18 0-18 0-19 0-18 0-18 0-19 0-21 0-18 0-18 0-18 0-18 0-18 0-18 0-18 0-1	Та	0.44	1 0.35		0.37				0.34	0.36		0.26		0.33 0.3		0.32	0.35	0.37				0.36	0.44	0.43 (	
4.27 2.65 3.11 4.00 3.38 3.46 2.72 3.87 4.67 3.15 2.19 3.59 4.31 2.78 3.40 2.61 4.03 3.07 3.36 4.74 4.19 3.26 0.38 0.41 0.29 0.61 0.19 0.25 0.20 0.33 0.47 0.41 0.11 0.17 0.17 0.19 0.16 0.15 0.17 0.17 0.16 0.20 0.18 0.17	Th	1-60	1.72		2.66				1.63	1.83		0.38		0.47 0.5		0.44	0.54	0.53				0.49	1.52	2.25	
0.38 0.41 0.29 0.61 0.19 0.25 0.20 0.33 0.47 0.41 0.11 0.17 0.17 0.19 0.16 0.15 0.17 0.17 0.16 0.20 0.18 0.17	Pb	4.27	2.65		4.00				3.87	4.67		2.19		4.31 2.78		2.61	4.03	3.07				3.26	3.06	4.00 4	
	D	0.38	3 0-41		0.61				0.33	0.47		0.11		0.17 0.1		0.15	0.17	0.17				0.17	0.27	0.51 (	

Move not interement in Fig. 1     The field of the	Sample: Type:* Location:	P27- AVL CT2 P	VF51 85 CT2 B	VF72- 85 CT2 B	P24- AVL CT3 P	P29- AVL CT3 P	P31- AVL CT3 P	VF24- 85 CT3 PR	VF34- V 85 85 85 85 CT3 0 0 PR	VF50- V 85 85 8 CT3 0 SR B	VF63- L 85 A CT3 C B L	UP203- U AVL A CT3 C UP U	UP205- UF AVL AV CT3 CT UP UF	UP209- UF AVL AV CT3 CT UP UF	UP212- SK AVL A CT3 C1 UP SK	SK214- SK: AVL AVI CT3 CT: SK SK	SK215- SK216- AVL AVL CT3 CT3 SK SK	116- SK219- AVL CT3 SK	19- VF110- 85 CT3 S	10- VF111- 85 CT3 S	VF145- 85 CT3 KB	- VF146- 85 CT3 KB	VF152- 85 CT3 KB
00.2         00.3         10.4 <th< td=""><td>Major and</td><td>minor eler</td><td>nents (w</td><td></td><td></td><td></td><td></td><td></td><td></td><td></td><td></td><td></td><td></td><td></td><td></td><td></td><td></td><td></td><td></td><td></td><td></td><td></td><td></td></th<>	Major and	minor eler	nents (w																				
17.         2.3         3.4         1.4 <th1.4< th=""> <th1.4< th=""> <th1.4< th=""></th1.4<></th1.4<></th1.4<>	SiO <sub>2</sub>	50.42	55-09		51.92	52.26	52.33	51.06	51.44	51.24		51.61											51-59
10.0         3.1.1         3.0.11         3.0.1         3.0.1 <th< td=""><td>TIO<sub>2</sub></td><td></td><td></td><td>3.08</td><td>1-46</td><td>1-64</td><td>1.66</td><td>1.87</td><td>1.50</td><td>1-44</td><td></td><td>1.49</td><td></td><td></td><td></td><td></td><td></td><td></td><td></td><td></td><td></td><td></td><td>1-5(</td></th<>	TIO <sub>2</sub>			3.08	1-46	1-64	1.66	1.87	1.50	1-44		1.49											1-5(
100         114         111         110         113         110         113         110         113         110         113         110         113         110         113         110         113         110         113         110         113         110         113         110         113         110         113         110         113         110         113         110         113         110         113 <td><math>AI_2O_3</math></td> <td></td> <td></td> <td>13.06</td> <td>14.52</td> <td>14-51</td> <td>14.86</td> <td>14.42</td> <td>14.23</td> <td>15-07</td> <td></td> <td>14.68</td> <td></td> <td>14.0</td>	$AI_2O_3$			13.06	14.52	14-51	14.86	14.42	14.23	15-07		14.68											14.0
01         017         018	FeO <sub>tot</sub>			14.77	11-01	11.23	11-06	12.06	10.95	11.39		10.19											11-58
708         648         571         649         771         644         709         641         700         644         700         641         700         641         700         641         700         641         700         641         700         641         700         741         700         641         700         741         700         741         760         741         760         741         750         751         753 <td>MnO</td> <td></td> <td></td> <td>0.17</td> <td>0.18</td> <td>0.18</td> <td>0.18</td> <td>0.18</td> <td>0.19</td> <td>0.18</td> <td></td> <td>0.17</td> <td></td> <td>0.19</td>	MnO			0.17	0.18	0.18	0.18	0.18	0.19	0.18		0.17											0.19
124         080         041         010         034         103         113         112         230         035         017         036         037         036         037         036         037         036         037         036 <td>NgO</td> <td></td> <td></td> <td>5.29</td> <td>7.17</td> <td>6.91</td> <td>6.18</td> <td>6.18</td> <td>7.20</td> <td>6.47</td> <td></td> <td>7.53</td> <td></td> <td>7-46</td>	NgO			5.29	7.17	6.91	6.18	6.18	7.20	6.47		7.53											7-46
10         20<	CaO			9.61	10.00	8.44	10.09	11.02	11.36	11.52		12.06											10-68
023         106         038         124         048         0.43         0.24         0.43         0.24         0.41         0.44	Va <sub>2</sub> O			2.39	2.21	2.78	2.41	2.63	2.47	2.30		1.97											2.07
015         0.23         0.34         0.14         0.24         0.14         0.35         0.34         0.14         0.35         0.34         0.34         0.35         0.34         0.35         0.34         0.35         0.34         0.35         0.34         0.35         0.34         0.35         0.34         0.35         0.34         0.35         0.34         0.35         0.34         0.35         0.34         0.35         0.35         0.34         0.35         0.34         0.35         0.34         0.35         0.34         0.35         0.34         0.35         0.35         0.34         0.35         0.34         0.35         0.34         0.35         0.34         0.35         0.34         0.35         0.35         0.34         0.35         0.34         0.35         0.34         0.35         0.34         0.35         0.34         0.35         0.34         0.35         0.34         0.35         0.34         0.35         0.35         0.34         0.35         0.34         0.35         0.34         0.35         0.35         0.34         0.35         0.35         0.35         0.35         0.35         0.35         0.35         0.35         0.35         0.35         0.35	<20 20			0-89	1.23	1.69	0.93	0.28	0.48	0.20		0.10											0.73
223         123         300         244         100         176         246         140         170         140 <td>205</td> <td></td> <td></td> <td>0.26</td> <td>0.18</td> <td>0.25</td> <td>0.21</td> <td>0.30</td> <td>0.20</td> <td>0.19</td> <td></td> <td>0.18</td> <td></td> <td>0.17</td>	205			0.26	0.18	0.25	0.21	0.30	0.20	0.19		0.18											0.17
10         0.44         0.43         0.56         0.54         0.56         0.54         0.56         0.54         0.72         0.43         0.54         0.72         0.43         0.54         0.72         0.43         0.54         0.54         0.54         0.54         0.56         0.54         0.72         0.43         0.54         0	10			3.00	2.04	2.10	n.d.	1.10	1.76	2.68		1.36											1-88
24         26         44         19         33         32         33         32         33         32         33         31<	ng-no.			0.43	0.58	0.56	0.54	0.52	0.58	0.54		0.61											0.51
242644193334323033323234333433343734373456115266373837317317319310310310561152663071332112891913846611725674446659390300300561162754616165610775744667659300300661982910313310312617737378381274667571310310743733612553173492578713637314369751031031047437336125630131427777039366373171047437336126668777770392663171044061047437336136475747675747675747476743733617676757476767574767675717074373361747676767676767676 <td>race elem</td> <td>ents (ppm,</td> <td>(</td> <td></td>	race elem	ents (ppm,	(																				
5713833073883253323353131931931311311319310621102254516673869766172574456106757562110225451667766598627744561086957511414092991018889928610715781738687738675144456553013043252751661797392697107160133861553013043252751846617172814719929029016013386155201304325216147157331411622929291601338615521730432521817417720144145243625174133264267663714117221411726147175361741433562728241711712614717636147146174144146242521714714	ç	33	24	26	26	44	19	33						29 2					33				
51152963071532112831951913546611172267444262561065592802811410726451611716617166171678687738882731712619829499101888912610715787398687738882711041331691331091261071578126184821171081061551571443738612522181331091261071578120241194811448133611343042512542161147157114715711471561571041561441338616571587433644574738823114104145174159174196751441477514414614715714414514414614713211441471571441461461471561441451441461471321144147157144146147146147145144146147<		415	371	383	307	338	325	352						75 31					411				
10         72         64         61         65         67         68         73         78         78         73         74         74         74         74         74         75         73           133         146         32         195         117         136         133         109         101         88         25         15         17         136         137         136         137         136         137         136         137         136         137         136         137         136         137         136         137         136         137         136         137         136         137         136         137         136         137         136         137         136         137         136         137         136         137         136         137         137         136         137         137         131         134         136         137         136         137         137         131         136         137         136         137         136         137         136         137         136         137         136         137         136         137         136         137         136         137	<u>ہ</u>	414	55	115	296	307	153	211						17 22					5 59				
114         140         92         96         101         88         89         92         86         107         157         87         38         87         38         87         38         87         38         87         38         173         186         173         166         173         166         173         38         133         109         126         177         166         137         103         133         104         137         133         153         144         373         361         152         137         103         103         103         103         103         103         103         103         103         104         133         314         217         147         153         144         147         153         144         147         153         144         147         145         147         133         334         335         335         345         346         347         140         147         146         147         146         147         146         147         146         147         146         147         146         147         146         147         146         147         146         147	i i	154	62	110	72	54	51	65						:12 6					3 16				
103         146         33         125         117         136         133         109         126         107         157         35         37         117         136         133         109         126         117         136         133         103         145         35         117         136         133         103         104         133         31         104         117         23         32         21         25         25         13         34         353         34         35         34         35         31         104         103         20         24         37         31         143         35         34         35         34         35         34         35         34         35         34         35         34         35         34         35         34         35         34         35         34         35         34         35         34         35         34         35         34	L,	80	114	140	92	06	66	101						86 8					2 119				
266         198         29.4         39.6         17.3         4.9         8.2         2.5         168         0.7         2.7         7.0         3.9         2.6         9.1         16.2         2.9         6.3         3.1         10.4           1.10         1.33         361         5.5         301         304         32         25         31         31         37         39         20         24         317         39         26         47           2.0         1.3         393         1.09         100         96         78         71         90         75         144         17         56         345         344         371         303         356         456         346         356         143         373         256         356         356         157         140         153         371         271         174         152         141         176         174         158         856         556         556         556         556         556         556         556         556         556         556         556         556         556         556         556         556         556         556         556	'n	185	103	146	93	125	117	136						87 3					3 160				
474         373         361         525         301         304         325         217         369         259         290         260         230         260         230         361         325         301         304         325         21         22         21         17         22         19         24         17         36         23         25         23         25         23         25         23         25         23         25         23         25         23         25         23         25         25         23         25         25         23         26         34         45         468         27         30         20         250         23         25         23         24         35         36         3	d5	2.4	26.6	19.8	29.4	39.6	17.3	4.9						7.0					5.9				
	ŗ	226	474	373	361	525	301	304						18 35					9 259				
140         133         393         109         100         96         18         11         96         12         96         142         96         142         96         142         96         142         96         451         453         545         546         326         143         336         451         545         517         94         753         336         453         546         517         643         586         153         536         446         12         386         551         143         153         351 <t< td=""><td></td><td>22</td><td>30</td><td>43</td><td>25</td><td>25</td><td>27</td><td>25</td><td></td><td></td><td></td><td></td><td></td><td>17 2</td><td></td><td></td><td></td><td></td><td>7 36</td><td></td><td></td><td></td><td></td></t<>		22	30	43	25	25	27	25						17 2					7 36				
6.20         6.31         4.37         3.36         4.37         3.36         4.37         3.36         4.37         3.46         4.37         4.20         3.51           14.43         14.3         1169         4.24         2.35         5.17         6.43         5.36         11.43         3.36         4.35         5.46         17.50         5.26         6.83         6.73         6.63         4.36         2.93         5.93         13.16         7.63         8.33           14.43         14.30         17.96         17.49         17.40         15.26         6.83         6.73         6.93         13.46         15.73           17.37         2.83         17.96         17.40         17.57         15.78         15.13         14.13         18.98         8.53         29.06         16.99         13.47           17.37         2.88         3.95         4.01         4.25         3.42         3.27         3.73         17.19         17.23         18.98         8.53         29.06         16.99         13.47           190         2.52         4.34         1.06         1.55         1.21         1.21         1.21         1.21         1.21         1.21         1.21 <td>5</td> <td>98</td> <td>140</td> <td>1/3</td> <td>63</td> <td>601</td> <td>100</td> <td>96</td> <td></td> <td></td> <td></td> <td></td> <td></td> <td>64 1L</td> <td></td> <td></td> <td></td> <td></td> <td>142</td> <td></td> <td></td> <td></td> <td></td>	5	98	140	1/3	63	601	100	96						64 1L					142				
$ \begin{array}{ ccccccccccccccccccccccccccccccccccc$	db S	3.13	6-20	6.74	3.97	3.94	4.21	4-50						3.36					2.18 6				
14.49         14.39         17.96         7.63         8.44         5.45         5.45         5.46         17.10         17.57         6.538         6.538         8.56         5.55         3.15         17.65         17.96         17.63         3.838         3.55         3.56         17.65         17.61         17.67         15.78         17.11         15.78         15.11         14.13         18.98         8.55         3.56         19.47         15.66         13.24         3.312         3.313         3.317         3.411         16.71         15.78         15.11         14.16         16.33         16.34         13.47           5.26         7.30         3.85         1.36         1.36         1.32         1.49         1.30         2.40         8.47         1.90         12.7         1.44         1.66         3.83         4.41         1.66         1.63         13.47         1.69         13.47           1.90         2.52         1.37         1.49         1.30         2.41         1.66         1.55         1.21         1.41         1.66         3.83         4.41         1.66           2.55         1.30         0.31         0.71         1.32         1.42         3.86	29	N -	20.7	233	100		424	0.52						10 21					1/0 1				
$ \begin{array}{ ccccccccccccccccccccccccccccccccccc$	e ,	12.22	30.93	33.83	8.11 17.95	17.49	17.07	8.49						5.26 11.77 1					3-59 13. 1.53 29.				15.45
$ \begin{array}{ ccccccccccccccccccccccccccccccccccc$		10.53	17.37	22.83	12.07	12.45	12.02	13.56						8.47 1					3.90 19.				
	E.	3.85	5.26	7.30	3.85	3.95	4.01	4.25						2.82					38 6.				
634         8-66         4-35         4-83         4-72         5-3         4-32         4-14         4-66         4-07         7-73         3-38         4-42         3-86         3-71         4-77         3-04         7-35         4-47         4-97           1-06         1-50         0-80         0-83         0-77         0-83         0-71         1-32         0-58         0-73         0-56         1-26         0-72         0-86           2-55         3-30         0-32         0-33         0-32         0-33         0-32         0-47         1-64         1-64         1-26         0-73         0-86           0-37         0-31         0-32         0-32         0-33         0-32         0-32         0-26         0-27         0-37         1-39         2-19         2-13           0-37         0-31         0-32         0-32         0-32         0-49         1-44         1-48         1-26         0-74         0-39         0-32         0-26         0-74         0-30         0-29         0-32         0-26         0-24         0-14         1-49         1-26         1-14         1-48         1-26         2-17         0-32         0-24         0-14	'n	1.43	1.90	2.52	1.37	1.47	1.53	1.58						1.06					1.04 2.				
$ \begin{array}{ ccccccccccccccccccccccccccccccccccc$	be	4.34	6.34	8.66	4.35	4.58	4.72	5.23						3.38					3-04 7.				
255         330         2.10         2.08         2.44         2.11         1.96         1.90         2.12         1.82         3.33         1.56         1.81         1.64         1.64         2.07         1.37         3.18         1.92         2.12           0.37         0.47         0.32         0.33         0.32         0.30         0.32         0.29         0.32         0.26         0.48         0.22         0.26         0.24         0.24         0.30         0.20         0.48         0.23         0.32         0.32         0.32         0.32         0.26         0.48         0.21         1.71         2.56         1.44         2.07         1.48         2.19         2.31         2.03           3.74         0.48         0.21         0.49         1.71         2.56         1.44         4.08         2.19         2.62         0.32 <td>م</td> <td>0.79</td> <td>1.06</td> <td>1.50</td> <td>0.80</td> <td>0.83</td> <td>0.87</td> <td>0.85</td> <td></td> <td></td> <td></td> <td></td> <td></td> <td>0.58</td> <td></td> <td></td> <td></td> <td></td> <td>)-54 1.</td> <td></td> <td></td> <td></td> <td></td>	م	0.79	1.06	1.50	0.80	0.83	0.87	0.85						0.58					)-54 1.				
0.37         0.47         0.32         0.33         0.39         0.30         0.29         0.32         0.48         0.22         0.26         0.24         0.24         0.30         0.29         0.32           3.73         5.13         2.31         2.55         2.64         2.55         2.07         1.98         2.31         2.00         4.97         1.71         2.56         1.92         1.94         4.65         1.44         4.08         2.19         2.62           0.44         0.48         0.21         0.19         0.19         0.19         0.17         2.56         1.92         1.94         4.08         2.19         2.62           0.44         0.48         0.21         0.19         0.19         0.19         0.19         0.20         0.20         0.20         0.20         0.20         0.23 <td>d)</td> <td>1.73</td> <td>2.55</td> <td>3.30</td> <td>2.10</td> <td>2.08</td> <td>2.24</td> <td>2.11</td> <td></td> <td></td> <td></td> <td></td> <td></td> <td>1.58</td> <td></td> <td></td> <td></td> <td></td> <td>1.37 3.</td> <td></td> <td></td> <td></td> <td></td>	d)	1.73	2.55	3.30	2.10	2.08	2.24	2.11						1.58					1.37 3.				
3.73       5.13       2.31       2.55       2.64       2.55       2.07       1.98       2.31       2.00       4.97       1.71       2.56       1.94       2.55       1.44       4.08       2.19       2.62         0.44       0.48       0.21       0.22       0.29       0.19       0.18       0.21       0.19       0.63       0.20       0.20       0.26       0.13       0.39       0.23       0.28         1.92       1.53       0.52       0.29       0.19       0.18       0.21       0.19       0.63       0.20       0.20       0.26       0.13       0.39       0.23       0.28         1.92       1.53       0.52       0.29       0.21       0.20       0.20       0.20       0.26       0.13       0.39       0.23       0.28         1.92       1.53       0.52       0.24       0.22       0.26       0.29       0.21       0.10       0.40       0.23       0.28       0.28       0.24       0.70       0.20       0.72       0.70       0.20       0.71       0.70       0.20       0.21       0.40       0.33       0.38       3.40       0.33       0.32       0.40       0.70       0.23 <td< td=""><td>ŋ</td><td>0.26</td><td>0.37</td><td>0.47</td><td>0.32</td><td>0.33</td><td>0.33</td><td>0.32</td><td></td><td></td><td></td><td></td><td></td><td>0.22</td><td></td><td></td><td></td><td></td><td>0.20 0.</td><td></td><td></td><td></td><td></td></td<>	ŋ	0.26	0.37	0.47	0.32	0.33	0.33	0.32						0.22					0.20 0.				
0-44 0-48 0-21 0-22 0-22 0-29 0-19 0-18 0-21 0-19 0-63 0-20 0-27 0-20 0-26 0-13 0-39 0-23 0-28 1-32 1-32 1-32 1-32 1-32 1-32 1-32 1-32	₽	2.46	3.73	5.13	2.31	2.55	2.64	2.55						1.71					1-44 4-				
1.92       1.53       0.52       0.38       0.37       0.22       0.36       0.22       0.36       0.22       0.44       0.41       0.26       0.29       0.27       0.42       0.11       0.78       0.40       0.32         3.85       3.46       2.57       2.19       2.22       1.21       1.26       3.74       1.54       1.84       1.17       1.63       2.24       0.71       3.33       1.60       2.38         3.85       3.46       2.57       2.19       2.22       1.21       1.26       1.47       1.56       3.74       1.54       1.84       1.17       1.63       2.24       0.71       3.23       1.60       2.38         0.49       0.73       0.73       0.11       0.07       0.11       0.07       0.11       0.07       0.12       0.04       0.08       0.08       0.11       0.03       0.10	Ta	0.17	0.44	0.48	0.21	0.22	0.22	0.29						0.20					0.13 0.				
3.85 3.46 2.57 2.19 2.25 2.22 1.21 1.26 1.47 1.26 3.74 1.54 1.84 1.17 1.63 2.24 0.71 3.23 1.60 2.38 0.49 0.73 0.13 0.10 0.11 0.10 0.08 0.01 0.11 0.01 0.10	Ъ	0.23	1.92	1.53	0.52	0.28	0.36	0.37						0.17					0.11 0.				
0.49 0.73 0.13 0.10 0.11 0.10 0.08 0.07 0.11 0.07 0.23 0.04 0.09 0.08 0.08 0.11 0.03 0.24 0.11 0.10	<sup>b</sup>	0.98	3.85	3.46	2.57	2.19	2.25	2.22						1.54					0.71 3.				
	_	0.06	0.49	0.73	0.13	0.10	0.11	0.10						0.04					0.03				

Table 1: continued

Table 1: continued

Sample:	VF88-	VF89-	VF91-	VF126-	VF129-
	85	85	85	85	85
Type:*	CT3	CT3	CT3	CT3	CT3
Location:	PD	PD	PD	Μ	Μ
Major and r	ninor elem	ents (wt %)	ht.		
SiO <sub>2</sub>	50.89	53.32	50.86	49.51	49.00
TiO <sub>2</sub>	1.43	2.19	1.17	1.50	1.97
Al <sub>2</sub> O <sub>3</sub>	14.89	14.01	14.34	15.14	14.46
FeO <sub>tot</sub>	10.31	13.31	10.75	10.72	12.44
MnO	0.18	0.20	0.19	0.17	0.17
MgO	8.09	4.22	7.40	7.92	6.21
CaO	11.34	9.60	12.99	12.57	12.46
Na₂O	2.02	2.37	2.00	2.12	2.27
K <sub>2</sub> O	0.62	0.47	0.18	0.17	0.77
$P_2O_5$	0.18	0.32	0.12	0.19	0.26
LOI	2.22	1.67	1.46	2.29	1.53
<i>mg</i> -no.	0.62	0.40	0.59	0.61	0.51
Trace eleme	ents (ppm)				
Sc	30	32	31	34	31
V	301	378	290	295	391
Cr	408	67	284	546	407
Ni	141	19	65	170	108
Zn	86	113	81	78	87
Cu	56	156	142	95	129
Rb	7.8	5.8	2.5	1.0	4.6
Sr	282	296	225	317	341
Υ	24	33	19	17	25
Zr	76	136	64	80	104
Nb	3.23	5.97	2.48	4.93	6.76
Ba	249	289	149	140	377
La	6.49	11.92	4.77	5.70	8.28
Ce	14.94	26.31	10.40	13.79	19.46
Nd	10.26	17.68	7.87	10.18	14.44
Sm	3.03	5.56	2.83	3.20	4.63
Eu	1.31	1.95	1.14	1.31	1.78
Gd	4.09	6.60	3.59	3.86	5.35
Tb	0.71	1.10	0.63	0.63	0.86
Yb	1.90	2.95	1.62	1.40	1.86
Lu	0.29	0.44	0.24	0.20	0.29
Hf	2.12	3.71	1.73	2.02	2.76
Та	0.19	0.35	0.15	0.38	0.53
Th	0.25	0.72	0.30	0.21	0.31
Pb	1.12	3.57	1.03	0.93	0.90
U	0.07	0.21	0.08	0.06	0.08

n.d., not determined; LOI, loss on ignition; mg-number = Mg/(Mg + 0.85Fe), atomic.

\*CT1–CT3 refer to geochemically different basalt types recognized in Vestfjella (see text). Abbreviations for sample locations: B, Basen; P, Plogen; PR, Pukkelryggen; SR, Salryggen; S, Skansen; PD, Pagodromen; KB, Kjakebeinet; SK, Steinkjeften; M, Muren; UP, Utpostane.

†The compositions are normalized to 100% volatile free.

Plagioclase is by far the most abundant phenocryst phase and typically occurs as glomerocrysts. Augite occurs as separate euhedral phenocrysts, subhedral granular aggregates, or glomerocrysts with plagioclase and olivine. Orthopyroxene phenocrysts have been identified in few lava flows only. Unaltered olivine is rare in the lavas, but many samples contain pseudomorphs after euhedral or slightly resorbed olivine phenocrysts. Chromian spinel inclusions are present in augite, orthopyroxene and olivine phenocrysts. The groundmass is aphanitic or fine grained, typically intersertal and consists of plagioclase laths, subhedral augite  $\pm$  pigeonite, quenched Fe–Ti oxides, minor sulphides and cryptocrystalline mesostasis.

Most of the Vestfjella lavas are altered. Olivine and orthopyroxene are commonly replaced by green or brown phyllosilicates and plagioclase is variably sericitized or saussuritized. In the most altered units, fresh plagioclase is rare and clinopyroxene is partially replaced by chlorite. The samples from Utpostane (Fig. 1) record low-grade metamorphism, most probably as a result of intrusion of the Utpostane gabbro; in many specimens, augite is replaced by tremolite–actinolite.

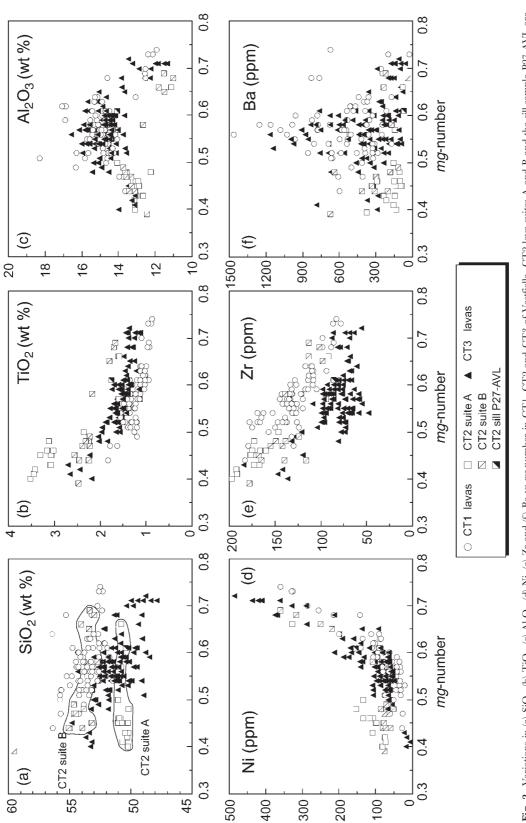
The bulk of the exposed strata consist of compound lava flows that show the characteristic features of inflated pahoehoe flows (Self *et al.*, 1996) with pipe amygdules at the bases and highly amygdaloidal upper parts. The lava units range from thin flow lobes to sheet flows exceeding 20 m in thickness. The samples for this study were collected from the massive interiors of the flow units to avoid amygdules. Nevertheless, many of the analysed samples contain small (<1 mm) dictytaxitic voids filled with secondary minerals, mainly chlorite. In the Utpostane suite, the predominant void-filling minerals are chlorite and tremolite–actinolite.

#### GEOCHEMISTRY

#### Major and trace elements

Chemical compositions of Vestfjella CFBs are listed in Table 1 and the Appendix. The data exhibit notably variable SiO<sub>2</sub> (47·9–56·5 wt %), TiO<sub>2</sub> (0·88–3·53 wt %), K<sub>2</sub>O (0·02–2·39 wt %), P<sub>2</sub>O<sub>5</sub> (0·11–0·47 wt %) and mg-number [atomic ratio of Mg/(Mg + 0·85Fe<sub>tot</sub>)] (0·74–0·40) (Fig. 2; Table 1). The rocks are subalkaline and range from basalt to basaltic andesite. On the basis of the CIPW norms (Yoder & Tilley, 1962) they may be classified as olivine and quartz tholeiites.

The major and trace elements exhibit compositional trends typical of tholeiites with, for example, negative correlation of  $FeO_{tot}$  and  $TiO_2$  with *mg*-number. Al<sub>2</sub>O<sub>3</sub> first increases with decreasing *mg*-number and then decreases when *mg*-number is <0.55 (Fig. 2). Trace elements with high ionic potential, such as Ni and Zr, typically show considerable ranges (Ni from 486 to 8 ppm; Zr





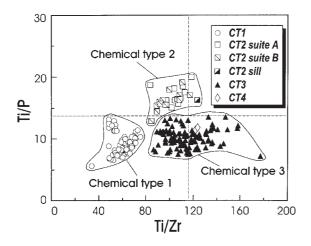


Fig. 3. Variations in Ti/Zr vs Ti/P in Vestfjella CFBs. Primitive mantle values (Sun & McDonough, 1989) are shown by stippled lines. Data for CT4 dyke rocks are from Luttinen *et al.* (1998).

from 48 to 238 ppm) and good correlations with *mg*number, whereas alkali and alkaline earth metals in general show large scatter and poor correlation with *mg*number (Fig. 2).

An important feature of the Vestfiella CFBs is that they fall naturally into three subgroups based on incompatible high field strength element (HFSE) and REE ratios. Such ratios are not considered to be susceptible to subsolidus alteration, fractional crystallization or small variations in degree of melting of peridotitic mantle sources. Basalts with notably different ratios thus represent different magmatic lineages with different evolution or mantle sources, or both. The plot of Ti/Zr vs Ti/P (Fig. 3) shows grouping for the new analyses demonstrating that the previously proposed subdivision for the CFBs of Basen and Plogen (Luttinen & Siivola, 1997) is applicable throughout Vestfjella. Following the nomenclature of Luttinen & Siivola (1997) the groups have been designated chemical types 1, 2 and 3 (CT1, CT2 and CT3) (Fig. 3). The CT4 dyke rocks of Luttinen et al. (1998) have Ti/Zr and Ti/P similar to those of CT3 (Fig. 3) but otherwise have a different overall geochemical composition.

The most diagnostic features of CT1, CT2, CT3 and CT4 can be summarized as follows. CT1 exhibits low Ti/Zr (<80) and TiO<sub>2</sub> compared with CT2 and CT3. It is also characterized by relatively high SiO<sub>2</sub> and (La/Sm)<sub>n</sub> and low Ni (Figs 2–4). CT2 shows considerable within-group variations. All CT2 rocks have near primitive mantle Ti/Zr and are characterized by high FeO<sub>tot</sub>, TiO<sub>2</sub> and Ti/P, and low Al<sub>2</sub>O<sub>3</sub> in comparison with CT1 and CT3 (Figs 2 and 3). The (La/Sm)<sub>n</sub> values typically are lower than those of CT1 (Fig. 4). A sill rock sample P27-AVL exhibits the most light REE (LREE)-depleted pattern among the Vestfjella CFBs with (La/Sm)<sub>n</sub> of

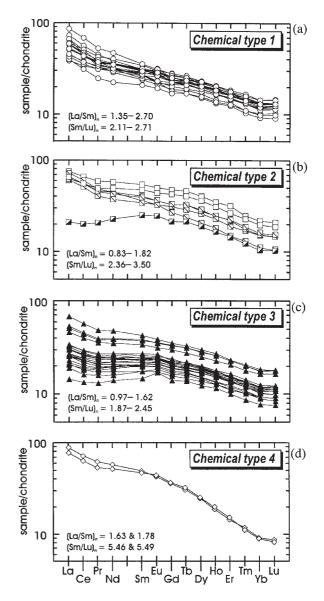


Fig. 4. Chondrite-normalized REE diagrams for (a) CT1, (b) CT2, (c) CT3 and (d) CT4 of Vestfjella. Normalizing values after Sun & McDonough (1989). Symbols for CT2 subtypes as in Fig. 2. Data for CT4 are from Luttinen *et al.* (1998).

0.83. A unique feature of CT2 is the bimodal distribution of SiO<sub>2</sub> contents leading to recognition of CT2 suites A and B (Fig. 2). The CT3 lavas exhibit near primitive mantle Ti/Zr and Ti/P and low HFSE (such as Zr) (Fig. 2) and REE concentrations relative to CT1 and CT2. Their REE patterns resemble those of CT2 with relatively low (La/Sm)<sub>n</sub>. In comparison, the CT4 dykes are notably rich in TiO<sub>2</sub>, Zr and other HFSE regarding their relatively high *mg*-number of 0.64 (Luttinen *et al.*, 1998) and show strong enrichment in LREE with (La/Lu)<sub>n</sub> of 8.9 and 9.8 (Fig. 4).

The incompatible element signatures of the Vestfjella CFB types are illustrated in mantle-normalized diagrams in Fig. 5. CT1, CT2 and CT3 typically exhibit negative Nb and Ta anomalies and positive Ba, K and Sr anomalies but have otherwise different patterns. All CT1 rocks show relative enrichment of LREE, positive Ba and K anomalies, and a distinctive negative Ti anomaly; this is designated the common signature of CT1. Most of the CT1 lavas further record a positive Sr anomaly and a negative P anomaly (Fig. 5). A subgroup of CT1 lavas, however, is distinguished by relatively high (P/  $Nd_n$  and low  $(Th/Ta)_n$  (Fig. 5). The CT2 rocks typically exhibit enrichment from HREE to Rb. The sill rock P27-AVL, however, shows affinity to depleted MORB in terms of  $(La/Nd)_n$ ,  $(Th/Ta)_n$  and  $(Rb/Th)_n <1$  but exhibits positive Ba, K and Sr anomalies. The CT2 tholeiites do not record a negative Ti anomaly but the lavas exhibit a distinctive negative P anomaly. The incompatible element signature of CT3 is rather similar to that of the low (Th/Ta)<sub>n</sub> subgroup of CT1 in terms of low Th and U and the pronounced positive Ba, K and Sr anomalies (Fig. 5). The CT3 rocks lack the negative Ti anomaly of CT1 and are less enriched in LREE. Besides Ba, K and Sr their pattern is relatively smooth and akin to that of MORB. The CT4 dykes have smooth patterns with an affinity to ocean island basalts (OIBs) (Fig. 5; Luttinen et al., 1998).

In summary, the four basaltic rock types of Vestfjella exhibit many differences in their chemical composition. The most abundant types, CT1 and CT3, have high  $Al_2O_3$  and low FeO<sub>tot</sub> and TiO<sub>2</sub> compared with CT2, show pronounced positive Ba and K anomalies in mantle-normalized diagrams, and also exhibit a broad negative correlation between SiO<sub>2</sub> and *mg*-number (Figs 2–5).

#### Geochemical stratigraphy

Stratigraphic variations in Ti/Zr, Ti/Y and *mg*-number for >200 lava samples from Vestfjella are illustrated in Fig. 6. CT1 is readily distinguished from CT2 and CT3 on the basis of Ti/Zr apart from a few intermediate cases. TiO<sub>2</sub> content and Ti/P have been used to divide the data into the CT2 and CT3 magma types (see Luttinen & Siivola, 1997).

The key features of the geochemical stratigraphy include: (1) preponderance of CT1 in the northern area (Basen, Plogen, Dagvola, Dagvolinga, Grushallet, Pukkelryggen and Salryggen) and CT3 in the southern area (Skansen, Pagodromen, Muren, Kjakebeinet, E and W Steinkjeften, and Utpostane); (2) repeated alternation of the three magma types (most notably in Basen and Plogen); (3) abundance of CT2 close to the base of the Basen and Plogen sections; (4) different Ti/Y of the lower and upper CT1 lavas in the Plogen section. Otherwise, the geochemical stratigraphy shows little evidence of systematic variations and, similar to many CFB suites, the basalt types show rather uniform compositions with intermediate *mg*-number values from the bottom to the top (cf. Arndt *et al.*, 1998).

The CT1 lavas may be subdivided into a lower and an upper suite based on their Ti/Y values. Throughout Vestfjella, the CT1 lavas that crop out over 600 m above sea level exhibit relatively high Ti/Y ratios (typically >300 (Fig. 6). These lavas are underlain by a succession of CT3-dominated flows in many localities. In the south, the few CT1 flows below the 600 m level at E and W Steinkjeften and Utpostane exhibit high Ti/Y ratios. At Pagodromen, however, the lower CT1 flows record low Ti/Y similar to those of the lower CT1 lavas in the north. It is possible that the CT3–CT1 transition within 600-800 m above sea level represents a regional stratigraphic horizon. If this is correct, then the predominance of CT3 in the southern area may, at least partially, be an artefact of exposure, and the CT3 suite could have erupted on low-Ti/Y CT1 throughout Vestfjella (Fig. 6).

Vestfjella region is characterized by major faulting structures and many sections show evidence of postvolcanic vertical movements (Furnes et al., 1987; Luttinen & Siivola, 1997). It is difficult to estimate the magnitude of fault-related displacements over the large distances between the sampled sections. To reconstruct the regional lava stratigraphy, the sections in Fig. 6 are arranged so that the CT3–CT1 transition coincides vertically. This requires fairly small height adjustments within the two major outcrop areas in the north (Basen, Plogen, Dagvola, Dagvolinga, Grushallet, Pukkelryggen and Salryggen) and south (Skansen, Pagodromen, Muren, Kjakebeinet, E and W Steinkjeften, and Utpostane) (Fig. 1). The sections in the south, however, have been generally moved up  $\sim 100$  m relative to the sections in the north. The two sections at Steinkjeften have been elevated more, up to  $\sim 200$  m (W Steinkjeften). The correlation between W Steinkjeften and Kjakebeinet is based on CT2 interbeds. Considering the distance of  $\sim 120 \,\mathrm{km}$  between Utpostane and Basen (Fig. 1), the deviations from the present positions are rather small in the reconstruction and may be explained by the regional W-SW dip and the post-volcanic faulting. The reconstruction in Fig. 6 shows the Vestfjella CFB suite to consist of three major stratigraphic units: the lower suite of CT1-dominated flows (Ti/Y <300), the middle suite of CT3-dominated flows and the upper suite of CT1-dominated flows (Ti/ Y >300). The CT2 suite A and B lavas (Fig. 2) are confined to levels below and above 200 m, respectively (Fig. 6).

Published data on 34 dyke and sill rocks that cut the lava sequence expand this discussion to later magmatic stages. On the basis of Ti/Zr and Ti/P ratios and TiO<sub>2</sub>

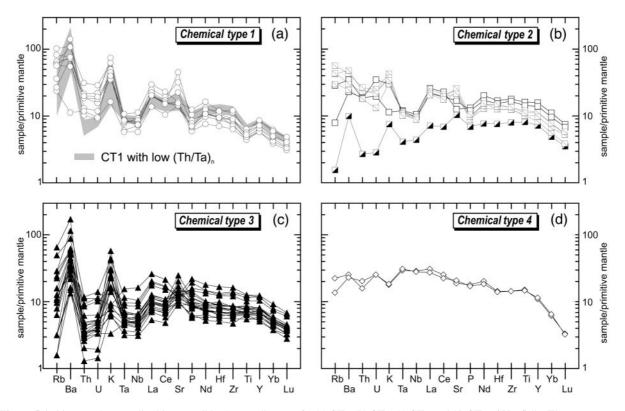


Fig. 5. Primitive mantle-normalized incompatible element diagrams for (a) CT1, (b) CT2, (c) CT3 and (d) CT4 of Vestfjella. The grey pattern in (a) includes nine CT1 lavas with low (Th/Ta)<sub>n</sub> relative to those of other CT1 lavas. Symbols for CT2 subtypes as in Fig. 2. Normalizing values after Sun & McDonough (1989). Data for CT4 are from Luttinen *et al.* (1998).

contents at given mg-number, it is possible to recognize four distinct dyke populations that show affinities to CT1, CT2, CT3 and CT4. Data of Furnes and Mitchell (1978) suggest that the dykes of south Vestfjella (Muren, Kjakebeinet, E and W Steinkjeften, Pagodromen and Skansen) belong almost exclusively to CT3 and only one out of 24 dykes shows affinity to CT1. In contrast to the volcanic record, CT2 is more abundant than CT1 among the studied dyke and sill rocks of north Vestfiella, with five out of 10 analysed samples featuring a CT2 affinity (Luttinen et al., 1998; this study). CT3 dykes have not been reported from the north. Two of the studied dykes record an OIB-like CT4 signature (Luttinen et al., 1998) suggesting this magma type occurs only as a minor dyke phase and is confined to north Vestfiella. In summary, the dyke rock data imply an increase in CT2 magmatism and a decrease in CT1 and CT3 magmatism associated with the later magmatic stages in the north Vestfjella. CT3 seems to have dominated magmatism in the south Vestfjella area throughout the volcanic episode recorded by the exposed part of the CFB succession. Emplacement of CT4 magmas was confined to relatively late magmatic stages in north Vestfjella.

#### Nd and Sr isotopes

New neodymium and strontium isotope data on 26 lavas and a sill rock from Vestfjella are listed in Table 2. The CT1 samples record highly variable Nd isotopic compositions. The  $\varepsilon_{Nd}$  (at 180 Ma) values range from -2.5 to -15.9 and show a bimodal distribution with a gap between a group with  $\varepsilon_{Nd}$  from -2.5 to -4.1 and two samples with  $\varepsilon_{Nd}$  of -11.1 and -15.9 (henceforth referred to as the high- $\varepsilon_{Nd}$  and low- $\varepsilon_{Nd}$  types, respectively) (Fig. 7). The CT2 samples also exhibit a wide range of  $\varepsilon_{Nd}$  values (+7.6 to -7.5). The  $\varepsilon_{Nd}$  of CT3 cluster around chondritic values (+2.0 to -2.2), with the majority being slightly positive.

The high- $\epsilon_{Nd}$  and low- $\epsilon_{Nd}$  CT1 lavas are characterized by different initial  $\epsilon_{Sr}$  values: those with distinctly low  $\epsilon_{Nd}$ exhibit  $\epsilon_{Sr}$  (180 Ma) of +33.4 and +64.7, whereas the others cluster between +8.6 and +20.7. The CT2 lavas record a wide range of  $\epsilon_{Sr}$  values from -15.7 to +49.6. The CT3 lavas show narrower range than CT1 and CT2 and have  $\epsilon_{Sr}$  from -11.0 to +5.5.

Generalizing, the new isotopic results from eight sections throughout the Vestfjella range accord with the previously published values (Luttinen *et al.*, 1998) for

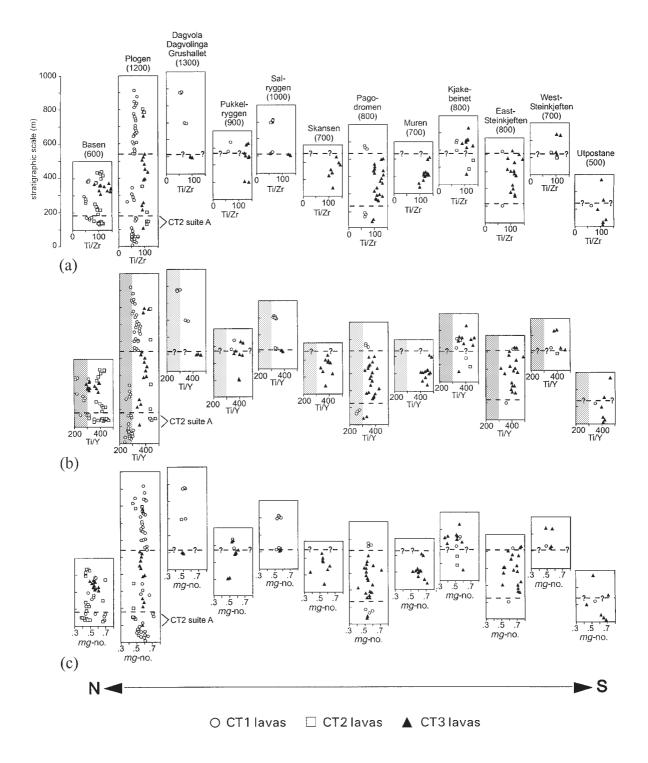


Fig. 6. Variations in (a) Ti/Zr, (b) Ti/Y and (c) mg-number of Vestfjella lavas in 12 sections. Section localities are indicated in Fig. 1. The upper dashed line indicates a potential correlation horizon based on a widespread CT3–CT1 transition. Question marks indicate sections with dubious CT3–CT1 transition. The lower dashed line shows a possible correlation horizon related to CT1–CT3 transition and the grey area in (b) points to Ti/Y values <300, which characterize the lower CT1 suite. The CT2 suite A is indicated below the lower dashed line. Numbers below the section names indicate the height of the top of the section in metres above sea level. Data sources: Furnes & Mitchell (1978); Furnes *et al.* (1987); Luttinen *et al.* (1998); this work.

Sample	Location*	Sm† (ppm)	Nd† (ppm)	<sup>147</sup> Sm/ <sup>144</sup> Nd	<sup>143</sup> Nd/ <sup>144</sup> Nd‡	ε <sub>Nd</sub> §	Rb† (ppm)	Sr† (ppm)	<sup>87</sup> Rb/ <sup>86</sup> Sr	<sup>87</sup> Sr/ <sup>86</sup> Sr‡	ε <sub>Sr</sub> §
Chemical ty	pe CT1										
VF-57-85	Basen	5.05	22.18	0.1377	$0{\cdot}512000{\pm}4$	-11.1	16.5	320	0.1470	$0{\cdot}707029\pm22$	+33.4
VF-80-85	Plogen	4.16	14.64	0.1718	$0.511795 \pm 6$	-15.9	36.4	344	0.3016	$0{\cdot}709623\pm11$	+64.7
VF-26-85	Pukkelryggen	4.62	16.35	0.1708	$0.512476 \pm 6$	-2.6	6.3	379	0.0474	$0{\cdot}705019{\pm}8$	+8.6
VF-30-85	Pukkelryggen	4.69	16.18	0.1752	$0{\cdot}512460{\pm}6$	-3.0	14.6	254	0.1638	$0{\cdot}705330\pm11$	+8.8
VF-40-85	Salryggen	4.42	14.94	0.1789	$0.512492\pm4$	-2.5	15.4	287	0.1529	$0.705858 \pm 30$	+ 16.7
VF-44-85	Salryggen	5.20	18.76	0.1676	$0.512450\pm4$	-3.0	16.1	311	0.1475	$0.706123 \pm 26$	+20.7
VF-46-85	Salryggen	4.81	16.41	0.1772	$0.512466 \pm 4$	-2.9	20.6	306	0.1919	$0.706104 \pm 26$	+ 18.8
VF-147-85	Kjakebeinet	4.05	13.79	0.1776	$0.512408 \pm 8$	-4.1	22.6	353	0.1825	$0.705466 \pm 14$	+ 10-1
VF-81-85	Pagodromen	4.52	16.14	0.1693	$0.512441 \pm 9$	-3.3	41.9	402	0.2971	$0.705805 \pm 28$	+10.7
Chemical ty	pe CT2										
VF-51-85	Basen	5.26	17.37	0.1831	$0{\cdot}512240\pm4$	-7.5	26.6	474	0.1599	$0.708197 \pm 11$	+ 49.6
VF-72-85	Basen	7.30	22.83	0.1933	$0.512522\pm7$	-2·2	19.8	373	0.1513	$0.706573 \pm 13$	+26.9
P55-AVL¶	Plogen	5.16	21.52	0.1450	$0.512569 \pm 10$	-0·2	18.6	457	0.1160	$0{\cdot}705820\pm14$	+ 17.5
P27-AVL¶	Plogen	3.57	11.25	0.1919	$0.513024 \pm 5$	+7.6	2.4	226	0.0303	$0{\cdot}703262{\pm}13$	- 15.7
Chemical ty	pe CT3										
VF-63-85	Basen	3.79	10.91	0.2100	$0{\cdot}512709{\pm}5$	+ 1.0	16.8	451	0.1062	$0{\cdot}704737\pm1$	+2.5
VF-24-85	Pukkelryggen	4.25	13.56	0.1895	$0{\cdot}512622\pm4$	-0·2	4.9	304	0.0459	$0{\cdot}704705 \pm 46$	+4.2
VF-34-85	Pukkelryggen	3.42	9.79	0.2112	$0.512744 \pm 7$	+ 1.7	8.2	325	0.0719	$0{\cdot}704036{\pm}12$	-6.2
VF-50-85	Salryggen	3.27	9.30	0.2126	$0.512743 \pm 7$	+ 1.6	2.5	278	0.0256	$0{\cdot}703872\pm11$	-6.9
VF-110-85	Skansen	6.16	19.48	0.1921	$0{\cdot}512602{\pm}7$	-0.6	6.3	259	0.0693	$0{\cdot}704090\pm12$	-5.4
VF-111-85	Skansen	3.82	11.69	0.1976	$0.512528\pm7$	-2·2	3.1	290	0.0305	$0{\cdot}704406{\pm}8$	+0.5
VF-88-85	Pagodromen	3.03	10.26	0.1785	$0.512693 \pm 8$	+ 1.5	7.8	282	0.0788	$0.703720 \pm 22$	- 11.0
VF-89-85	Pagodromen	5.56	17.68	0.1901	$0.512624\pm4$	-0.2	5.8	296	0.0558	$0.704061 \pm 18$	-5.3
VF-91-85	Pagodromen	2.83	7.87	0.2174	$0.512718 \pm 7$	+ 1.0	2.5	225	0.0317	$0.704254 \pm 12$	-1.7
VF-145-85	Kjakebeinet	4.41	13.47	0.1979	$0{\cdot}512579{\pm}6$	-1.2	10-4	260	0.1140	$0.704567\pm32$	-0.2
VF-146-85	Kjakebeinet	6.33	19.79	0.1934	0.512571±8	-1.3	22.3	245	0.2594	$0.705342\pm9$	+5.5
VF-152-85	Kjakebeinet	3.60	10.91	0.1995	$0.512677\pm 6$	+0.7	9.3	272	0.0974	$0.703971 \pm 15$	-8.1
VF-126-85	Muren	3.20	10.18	0.1900	$0.512733 \pm 6$	+2.0	1.0	317	0.0090	$0.703658 \pm 14$	-9.3
VF-129-85	Muren	4.63	14.44	0.1938	0.512648+40	+0.2	4.6	341	0.0384	0.703679+18	<i>—</i> 10·1

Table 2: Isotopic data of Mesozoic flood basalts in Vestfjella, Dronning Maud Land, Antarctica

\*See Fig. 1.

†Analysed with ICP-MS at the Washington State University apart from samples P55-AVL and P27-AVL (see below).

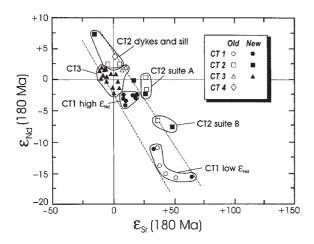
 $^{+143}$ Nd/<sup>144</sup>Nd normalized to <sup>146</sup>Nd/<sup>144</sup>Nd = 0.7219, <sup>87</sup>Sr/<sup>86</sup>Sr to <sup>86</sup>Sr/<sup>84</sup>Sr = 0.1194. Within-run error expressed as  $2\delta_m$  in the last significant digits. Sinitial  $\epsilon_{Nd}$  and  $\epsilon_{Sr}$  values, calculated at 180 Ma using <sup>143</sup>Nd/<sup>144</sup>Nd = 0.51264 and <sup>147</sup>Sm/<sup>144</sup>Nd = 0.1966, and <sup>87</sup>Sr/<sup>86</sup>Sr = 0.7045

similar  $\mathcal{E}_{Nd}$  and  $\mathcal{E}_{Sr}$  values, calculated at 100 Ma using Nu/ Nu = 0.51204 and Sin/ Nu = 0.1960, and Si/ Si = 0.7045 and  $^{87}Rb/^{66}Sr$  = 0.0816, respectively.

**(**Sm, Nd, Rb, and Sr concentrations and isotopic ratios analysed at the Geological Survey of Finland [for analytical methods, see Luttinen *et al.* (1998)].

eight lavas and four dykes from the Basen and Plogen sections. The available data show a broad negative correlation between initial  $\epsilon_{Nd}$  and  $\epsilon_{Sr}$  values (Fig. 7). The CT1 rocks plot in two groups along a low- $\epsilon_{Nd}$ -high- $\epsilon_{Sr}$  continuation of the oceanic mantle array (see Menzies & Murthy, 1980). The high- $\epsilon_{Nd}$  group of CT1 comprises the group with low (Th/Ta)<sub>n</sub> (Fig. 5) delineating a previously unknown subtype of CT1. The CT2 dykes

and sill plot within the depleted part of the mantle array as do most of the CT3 lavas and the CT4 dyke. The CT2 lavas, on the other hand, plot in two groups aside from the mantle array as a result of relatively high  $\epsilon_{Sr}$  at given  $\epsilon_{Nd}$ . The two CT2 lavas with the lowest  $\epsilon_{Nd}$  belong to the high-SiO<sub>2</sub> suite B (Fig. 2), whereas those with near-chondritic  $\epsilon_{Nd}$  have low SiO<sub>2</sub> and belong to suite A apart from sample P55-AVL. The  $\epsilon_{Nd}$  value of the CT2



**Fig. 7.** Initial Sr and Nd isotopic compositions of Vestfjella CFBs presented in  $\varepsilon_{Sr}$  vs  $\varepsilon_{Nd}$  diagram. Previously published analyses (Luttinen *et al.*, 1998) are are shown for comparison. The dashed lines feature the oceanic mantle array and its low- $\varepsilon_{Nd}$ -high- $\varepsilon_{Sr}$  continuation (see Menzies & Murthy, 1980).

sill rock sample P27-AVL (+7.6) is the highest so far reported for Karoo CFBs in Africa and Antarctica, and approaches the isotopic composition of depleted mantle at 180 Ma (DePaolo, 1981*a*).

#### Isotopic stratigraphy

Isotopic data of Vestfjella CFBs are combined in Fig. 8, where  $\varepsilon_{Nd}$  and  $\varepsilon_{Sr}$  are plotted using the proposed stratigraphic reconstruction (Fig. 6). The figure demonstrates dominance of CT1 with low  $\varepsilon_{Nd}$  and high  $\varepsilon_{Sr}$  at low elevations and CT1 with high  $\varepsilon_{Nd}$  and low  $\varepsilon_{Sr}$  at high elevations. It should be noted, however, that CT1 lavas with high and low (Th/Ta)<sub>n</sub> values (Fig. 5) are interbedded in the upper part of the Plogen section. Judging from this and the fact that all the high- $\varepsilon_{Nd}$  CT1 lavas exhibit low (Th/Ta)<sub>n</sub>, it is probable that low- $\varepsilon_{Nd}$  lavas also occur above the 600 m level.

The CT2 samples close to the base of the succession show near-chondritic  $\varepsilon_{Nd}$  and  $\varepsilon_{Sr}$  (Fig. 8) and belong to the CT2 suite A (Figs 2 and 6). The overlying set of CT2 flows record distinctly negative  $\varepsilon_{Nd}$  and positive  $\varepsilon_{Sr}$ and belong to the CT2 suite B. The single CT2 flow at 800 m level (P55-AVL) has a near-chondritic  $\varepsilon_{Nd}$  value and a slightly positive  $\varepsilon_{Sr}$  (Fig. 8) different from the other suite B rocks. The CT3 lavas show a relatively uniform isotopic composition throughout the lava pile. The overall isotopic stratigraphy and the positive  $\varepsilon_{Nd}$  of the crosscutting CT2 dykes and the sill rock (P27-AVL) and the CT4 dyke (Fig. 7) suggest a general increase in  $\varepsilon_{Nd}$  values during the Vestfjella CFB magmatism.

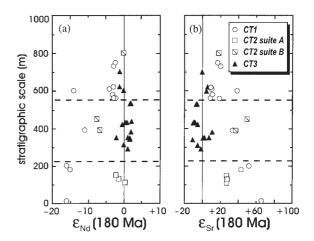


Fig. 8. Stratigraphic variations in  $\epsilon_{Nd}$  (180 Ma) and  $\epsilon_{Sr}$  (180 Ma) of Vestfjella lavas assuming reconstruction of Fig. 6. The dashed lines at  $\sim 550$  m and 220 m indicate divisions between the upper, middle and lower suites. Data sources: Luttinen *et al.* (1998); this work.

#### **Alteration effects**

The scatter of major and trace element data (Fig. 2) probably reflects the net result of variable melt compositions, phenocryst abundances and subsolidus alteration. Detailed assessment of alteration effects is beyond the scope of this study. We deal here with the effects of alteration on incompatible elements and isotopic ratios. Studies on altered basalts have demonstrated that elements with low ion potential, such as Rb, K, Ba and Sr, are often mobile whereas HFSE and REE retain their abundances fairly well (e.g. Wood *et al.*, 1976).

#### Alteration effects on incompatible trace elements

A geochemical traverse through a CT3 lava flow of 16 m thickness (Appendix) illustrates compositional differences between the amygdaloidal, strongly altered lava crust and the massive flow core (Fig. 9). High LOI values reflect abundant secondary phyllosilicates in voids and alteration products of magmatic minerals in the lava crust. Concentrations of Ni, Nb, Zr and Y are nearly constant from the base to the top of the flow, demonstrating within-flow homogeneity and immobility of HFSE. In the flow core, Rb, Ba and Sr contents show relatively small variations (4-7, 199-273 and 258-269 ppm, respectively). In the upper crust, Rb and Ba are clearly lower ( $\leq 2$  and  $\leq 30$  ppm, respectively) and Sr contents generally higher (up to 442 ppm) than in the core. The variation of K is similar to that of Rb. Variations of the trace element contents within the core of lava flows are further illustrated by three analyses from an extensive CT2 lava flow of 10 m thickness (Appendix), which show wide ranges of Rb (1-22 ppm), K<sub>2</sub>O (0·22-1.12 wt %) and Ba (137-217 ppm), and a narrow range of Sr (264–295 ppm) in comparison with relatively

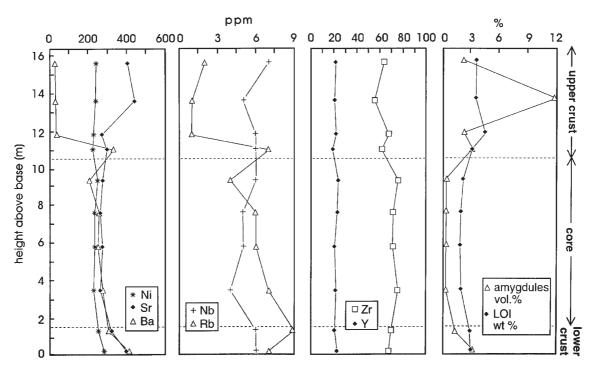


Fig. 9. Variations of selected trace elements, LOI and the content of amygdules in a CT3 lava flow from Skansen, south Vestfjella.

constant Zr (191–197 ppm). These examples lead to the following (qualitative) conclusions: (1) the relative mobility seems to have been Rb  $\approx$  K > Ba  $\gg$  Sr; (2) alteration has decreased concentrations of Rb, K and Ba rather than increased them.

The behaviour of mobile elements is of special interest in the case of CT3 lavas, which, with the exception of high LILE, exhibit incompatible element patterns akin to those of depleted MORB (Fig. 5). It is essential to determine whether the selective enrichment of Ba, K and Sr results from alteration or reflects original magmatic compositions. The examples presented above imply that strong enrichment of LILE is an unlikely result of alteration (see Harris et al., 1990). Assuming that the LILE concentrations have been significantly modified by secondary processes, we would expect some of the rocks to show depletion of Ba relative to other highly incompatible elements. In contrast, the CT3 samples representing a stratigraphic section of >600 m thickness show similar incompatible element signatures with (Ba/Nb)<sub>n</sub> consistently higher than two (Fig. 5). The identical REE patterns of the metamorphosed CT3 lavas from the Utpostane section (Fig. 1; Table 1) and the unmetamorphosed CT3 lavas form other sections in Fig. 4 lend support to immobility of REE.

We conclude that although concentrations of Rb, K, Ba and Sr of individual samples should be viewed with caution it is unlikely that the high overall LILE contents are solely an alteration effect. They more probably represent a regional feature related to the magmatic evolution of the Vestfjella CFBs. It is possible that, on average, the whole-rock analyses of flow cores approximate magmatic compositions of CT1, CT2 and CT3 also in the case of LILE.

#### Alteration effects on Nd and Sr isotopes

Recent studies on Ferrar tholeiites have demonstrated pronounced changes in the Sr isotopic signatures even in rocks that appear relatively unaltered (Fleming *et al.*, 1995). On the other hand, Nd isotopic systematics in extensively altered rocks may still reflect magmatic compositions (Landoll *et al.*, 1994; Foland *et al.*, 1997). It is important to evaluate possible effects of subsolidus alteration on the Sr isotopic compositions of the Vestfjella CFBs.

Whereas CT1, CT3 and the intrusive CT2 define a broad linear array ( $r^2 = 0.87$ ) in the  $\varepsilon_{Nd}$  vs  $\varepsilon_{Sr}$  diagram, the CT2 lavas are distinct owing to high  $\varepsilon_{Sr}$  at given  $\varepsilon_{Nd}$  (Fig. 7). It is rather unlikely that the samples that plot within the 'mantle array' have been significantly affected with regard to their Sr isotopic compositions. The high  $\varepsilon_{Sr}$  values of the CT2 lavas, on the other hand, are a typical feature of many hydrothermally altered basic rocks (Fleming *et al.*, 1995; Foland *et al.*, 1997).

It is probable that the degree of alteration has been largely controlled by the access of hydrothermal fluids

to the lava flows (see Bevins et al., 1991). The CT2 lavas are no different from CT1 and CT3 in terms of vesicularity and jointing or LOI values and petrographic evidence of degree of alteration. Similar to CT1 and CT3, they further record low <sup>87</sup>Rb/<sup>86</sup>Sr (Table 2). The measured <sup>87</sup>Sr/<sup>86</sup>Sr ratios thus closely correspond to the Sr isotopic compositions immediately after the closure of the system with regard to Rb and Sr, presumably in the Jurassic. If the Sr isotopic ratios of the CT2 lavas have been significantly increased as a result of subsolidus alteration, a specific, Sr isotopically different fluid is required that circulated in the sporadic CT2 interbeds but avoided the intercalated CT1 and CT3 flows. Such an explanation seems unlikely and further could not explain why the samples with high  $\varepsilon_{Sr}$  values also have different immobile incompatible element characteristics from the rocks with lower  $\varepsilon_{Sr}$  values: negative correlation between  $\varepsilon_{Nd}$  and Th/Ta (Fig. 10a) suggests the latter has not been affected by alteration. Accordingly, the positive correlation between  $\varepsilon_{sr}$  and Th/Ta (Fig. 10b) and the fact that only the high- $\varepsilon_{Sr}$  rocks exhibit a marked negative P anomaly (Fig. 5) indicate that the high  $\varepsilon_{sr}$  values of the CT2 lavas reflect magmatic compositions.

#### Cause of regional alteration

Pronounced subsolidus alteration of lavas is a regional feature of Vestfjella. By comparison, the basalts in Kirwanveggen and Heimefrontfjella are relatively unaltered. One possible explanation for this difference is the abundance of crosscutting intrusive equivalents of CFBs in Vestfjella.

The replacement of augite by amphibole in the Utpostane basalts is a likely consequence of contact metamorphism as a result of the emplacement of the Utpostane gabbro. Geophysical surveys have indicated sources of magnetic anomalies beneath the Högisen ice dome, N–NW of the Muren gabbro and beneath the Plogbreen ice stream between Basen and Plogen (Fig. 1); these have been interpreted as Jurassic intrusions (Corner, 1994; Ruotoistenmäki & Lehtimäki, 1997). It is thus possible that gabbroic intrusions of variable size characterize the entire Vestfjella region. Furthermore, the CFBs are cut by abundant dolerite dykes and less abundant sills. Emplacement of the various intrusive rock types could have promoted hydrothermal circulation of fluids and the alteration of the lavas.

#### **Relationships to other Karoo CFBs**

It has been generally accepted that western Dronning Maud Land was juxtaposed to southeast Africa before break-up of Gondwana, although the details of Jurassic configuration remain an issue of discussion (e.g. du Toit, 1937; Martin & Hartnady, 1986; Lawver & Scotese,

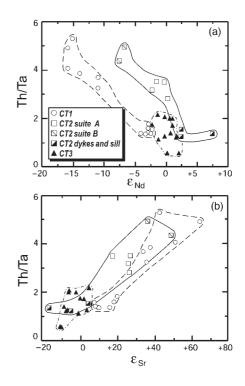


Fig. 10. Variations of (a)  $\varepsilon_{Nd}$  and (b)  $\varepsilon_{Sr}$  vs Th/Ta in Vestfjella CFBs. Data sources: Luttinen *et al.* (1998); this work.

1987). The reconstruction of Martin & Hartnady (1986) has been widely used and it places western Dronning Maud Land adjacent to the Lebombo Monocline of Africa (Fig. 1). Accordingly, the Mesozoic CFBs of these areas have been reported to show isotopic and geochemical similarities (Faure *et al.*, 1979; Harris *et al.*, 1990, 1991; Luttinen & Siivola, 1997; Luttinen *et al.*, 1998).

Cox (1992) has proposed a tight Jurassic fit of southeast Africa and East Antarctica, which positions the Vestfjella and south Lebombo areas notably close to each other. We have taken this to be the initial configuration and examine here the incompatible element and Nd and Sr isotopic signatures of the CFB rock types of the Lebombo and western Dronning Maud Land to evaluate possible relationships between them. In Vestfjella, the main magma types can be recognized based on Ti/Zr and Ti/P alone (Fig. 3). More informative, for a regional comparison of Karoo CFBs in Africa and Antarctica, are depleted MORB-normalized incompatible element patterns and specifically the diagnostic Nb, P and Ti anomalies shown by the various rock types. For the comparison, we have divided the various Karoo CFBs into low-Ti, high-Ti and transitional types (see Cox et al., 1967; Erlank et al., 1988).

### Low-Ti rock types

Low-Ti tholeiites dominate the CFB successions of Vestfjella, Kirwanveggen and southern to central Lebombo Monocline (Erlank, 1984; Harris et al., 1990; Sweeney et al., 1994). All of the low-Ti basalts exhibit a pronounced negative Nb anomaly (Fig. 11). Geochemical and Nd and Sr isotopic similarities led Luttinen & Siivola (1997) and Luttinen et al. (1998) to propose that the low- $\varepsilon_{Nd}$  CT1 suite of Vestfjella represents a counterpart of the low- $\varepsilon_{Nd}$  type of Sabie River basalts of southernmost Lebombo. Few REE analyses have been reported for the Sabie River basalts in this area (Erlank, 1984). Nevertheless, averages of the two magma types show high Ba<sub>n</sub> and nearly indistinguishable trace element patterns from La to Yb with low  $(P/Nd)_n$  (0.56 and 0.69) and  $(Ti/Gd)_n$ (0.71 and 0.75) and high  $(\text{La/Yb})_n$  (9.35 and 9.76) (Fig. 11). Both rock types further exhibit lower  $(Nb/La)_n$  (0.17) and 0.42) and  $\varepsilon_{\rm Nd}$  (180 Ma) values (< -10) and higher  $\epsilon_{Sr}$  (180 Ma) values (> +25) than the other low-Ti basalts (Figs 11 and 12). The average  $Nb_n$  value (1.1) of the representative southern Lebombo basalts is notably lower than that of the low- $\epsilon_{\rm Nd}$  CT1 basalts (2·8). It should be noted, however, that data for 34 Sabie River basalts, southern Lebombo (Erlank, 1984) show Nb<sub>n</sub> of 0.9-6.8with an average of 2.2, which is rather similar to the average value for the low- $\varepsilon_{Nd}$  CT1.

Our new data for the high- $\varepsilon_{Nd}$  CT1 type and CT3 bring detail to the correlation between the low-Ti CFBs in Africa and Antarctica. Relative to the low- $\varepsilon_{Nd}$  CT1 and Sabie River basalts of south Lebombo the other low-Ti rocks show slightly higher  $(Nb/La)_n$  (0.49–0.61), (P/ Nd)<sub>n</sub> (0.69–1.16), (Ti/Gd)<sub>n</sub> (0.85–1.09) and  $\epsilon_{Nd}$  (> –6), and lower (La/Yb)\_n (3.92–7.15) and  $\epsilon_{Sr}$  (< +25) (Figs 11 and 12). Overall, these rock types show many similarities but clear correlatives are not apparent. The high- $\varepsilon_{Nd}$  subtype of CT1 exhibits similar isotopic compositions to those of some Sabie River basalts of the central Lebombo (Fig. 12), but the latter have higher  $TiO_2$  and lower SiO<sub>2</sub> at the same *mg*-number (Sweeney *et al.*, 1994) as well as relatively higher  $(Rb/Ba)_n$  and  $(Th/K)_n$  (Fig. 11). In terms of Sr isotopic compositions the high- $\varepsilon_{Nd}$ CT1 also shows affinity to Kirwan basalts (Fig. 12). Furthermore, both types record a small negative P anomaly and rather similar patterns from K to Yb in general. The Kirwan basalts have notably lower Ba<sub>n</sub> contents and higher  $\varepsilon_{Nd}$  values, however (Figs 11 and 12).

Similarities between the CT3 lavas and the Kirwan basalts include relatively high  $(Ti/Gd)_n$  (1.03 and 0.97) and low  $(La/Yb)_n$  (4.64 and 3.92) and near-chondritic  $\varepsilon_{Nd}$  (+3 to -3) (Figs 11 and 12). The high Ba<sub>n</sub> and low  $\varepsilon_{Sr}$  of CT3 stand in contrast to the lower Ba<sub>n</sub> and higher  $\varepsilon_{Sr}$  of Kirwan basalts. Harris *et al.* (1990) have pointed out chemical similarities between the Kirwan basalts and

the Sabie River basalts of south Lebombo. Judging from the Nd isotopic compositions and relatively low (Ba/Nb)<sub>n</sub>, the Kirwan basalts resemble the extrusive equivalents of the Rooi Rand dykes rather than the low-Ti Sabie River basalts of southern Lebombo (Fig. 12).

#### Transitional rock types

Similar to Rooi Rand dolerites and their extrusive equivalents (e.g. Hawkesworth *et al.*, 1984; Duncan *et al.*, 1990), CT2 shows affinity to high-Ti flood basalt types based on high Ti/Y (mostly from 300 to 600) but their Zr/Y are typically low (<6) and within the range of low-Ti CFBs (Erlank *et al.*, 1988). Compared with the low-Ti and high-Ti rock types, these rock types have intermediate TiO<sub>2</sub> at given *mg*-number. Hence, they can be regarded as transitional types between low-Ti and high-Ti CFBs.

Both transitional rock types are spatially associated with low-Ti basalts and occur as crosscutting dykes and interbedded lavas. Luttinen & Siivola (1997) and Luttinen et al. (1998) have previously pointed out geochemical and Nd and Sr isotopic similarities between the CT2 and Rooi Rand magma types. These are characterized by variable incompatible element patterns and, for this reason, representative CT2 and Rooi Rand dykes are compared with the average of the CT2 lavas in Fig. 11. Some rocks exhibit a strong MORB affinity with rather low LREE relative to heavy REE (HREE), MORB-like HFSE ratios and positive initial  $\varepsilon_{\rm Nd}$  (up to +8). Others, such as the CT2 lavas, record LREE enrichment, marked negative Nb and P anomalies and negative  $\varepsilon_{Nd}$  (Figs 11 and 12). Our new data on CT2 demonstrate similarity between the intrusive CT2 and the Rooi Rand magma type on one hand and difference between these rock types and the extrusive CT2 on the other. These differences imply that the CT2 lavas may correspond to contaminated differentiates of MORB-like CT2 parental magmas.

#### High-Ti rock types

Rocks belonging to the high-Ti category are abundant in Mwenezi area and the northern half of the Lebombo Monocline and characterize the Ahlmannryggen dyke swarm of Dronning Maud Land (e.g. Erlank *et al.*, 1988; Harris *et al.*, 1991; Sweeney *et al.*, 1994) (Fig. 1). In Vestfjella, the CT4 dykes show affinities to the high-Ti group based on notably high HFSE and LREE contents as well as Ti/Y (>400) and Zr/Y (5·7) at *mg*-number of 0·64.

We have subdivided the high-Ti rock types on the basis of  $(Nb/La)_n$  and  $(P/Nd)_n$  values (Fig. 11). The group with relatively high  $(Nb/La)_n$  (0.69–1.13) and  $(P/Nd)_n$  (0.66–0.91) includes the CT4 dykes, the high-Fe subtype of high-Ti rocks, central Lebombo (Sweeney *et al.*, 1994), and some of the Ahlmannryggen dykes (Harris *et al.*,

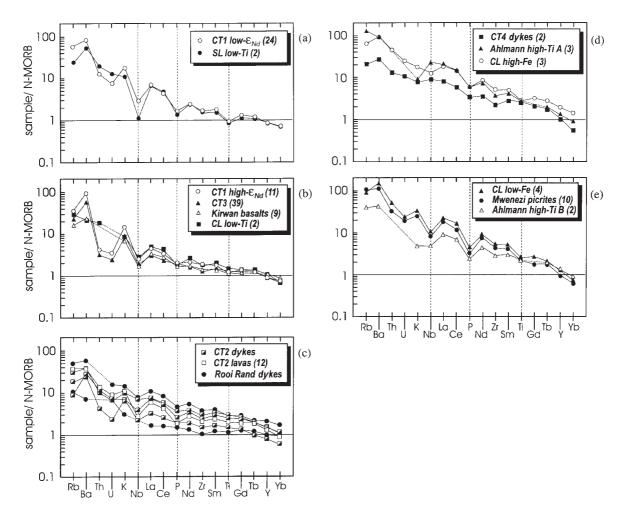
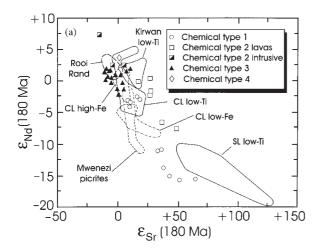


Fig. 11. N-MORB-normalized incompatible element patterns of Jurassic CFBs from southern Africa and Dronning Maud Land, Antarctica. SL low-Ti and CL low-Ti in (a) and (b) are averages of Sabie River low-Ti basalts from south Lebombo and central Lebombo, respectively. CL high-Fe and CL low-Fe in (d) and (e) are Sabie River high-Ti subtypes from central Lebombo. Ahlmann high-Ti A and B in (d) and (e) indicate two subtypes of high-Ti dykes from Ahlmannryggen. Normalizing values after Sun & McDonough (1989). Data sources: Erlank (1984); Harris et al. (1990, 1991); Ellam & Cox (1991); Sweeney et al. (1994); Luttinen et al. (1998); this work.

1991). These rocks have rather similar and smooth patterns with high  $(La/Yb)_n$  (13·13–23·40) (Fig. 11). The CT4 dykes and the Ahlmannryggen dykes are in fact among the few rocks in the Karoo CFB province that do not exhibit a negative Nb anomaly (see Fig. 11). Some of the high-Fe rocks from central Lebombo resemble CT4 also in terms of Nd and Sr isotopic compositions (Fig. 12). Nd isotopic data have not been published for the Ahlmannryggen dykes but their  $\varepsilon_{sr}$  (180 Ma) (-4 to +37; Harris *et al.*, 1991) extend to values that are comparable with those of the CT4 dykes and the high-Fe rocks of central Lebombo. The incompatible element pattern of the low-Fe subtype of the high-Ti rocks of central Lebombo (Sweeney et al., 1994) is nearly identical with that of the Mwenezi picrites: both have distinctive negative Nb and P anomalies (shared also by some Ahlmannryggen dykes) and comparable ranges of  $\varepsilon_{Sr}$ , and their  $\varepsilon_{Nd}$  values extend to notably low values (Figs 11 and 12).

#### Summary of possible relationships

Judging from the Nd and Sr isotope data and the MORBnormalized incompatible element diagrams, unequivocal correlatives for Vestfjella magma types are not found among the low-Ti CFBs of the Lebombo Monocline and Kirwanveggen. The low- $\varepsilon_{Nd}$  CT1 and the Sabie River basalts of southern Lebombo, however, exhibit many similarities and differ from the other low-Ti types. Another feature that links the southern Lebombo and northern Vestfjella is the similarity between the CT2 and Rooi Rand magma types. The high- $\varepsilon_{Nd}$  CT1 and CT3 lavas,



**Fig. 12.**  $\epsilon_{Nd}$  (180 Ma) and  $\epsilon_{Sr}$  (180 Ma) of Jurassic CFBs from southeast Africa and western Dronning Maud Land. Abbreviations as in Fig. 11. Data sources: Hawkesworth *et al.* (1984); Harris *et al.* (1990, 1991); Ellam & Cox (1991); Sweeney *et al.* (1994); Luttinen *et al.* (1998); this work.

which dominate in the southern part of Vestfjella, have notably high Ba contents and do not have obvious counterparts elsewhere in the Karoo volcanic province. The Kirwan basalts show intermediate features between the high- $\varepsilon_{Nd}$  CT1 and the CT3 basalts with respect to HFSE and REE. In summary, the low-Ti succession of Vestfjella may correspond to a southward extension of the laterally heterogeneous Sabie River Basalt Formation of the Lebombo Monocline (see Duncan *et al.*, 1984; Hawkesworth *et al.*, 1984; Sweeney *et al.*, 1994).

The CT4 dykes show similarities to some high-Ti CFBs of Ahlmannryggen and central Lebombo. The other high-Ti rocks are characterized by spiked patterns with negative Nb and P anomalies resembling those of low-Ti and transitional CFBs (Fig. 11). More geochronological, geochemical and isotopic data are required for the Ahlmannryggen and CT4 dykes to address the genetic relationships between the high-Ti CFB types in detail.

### DISCUSSION

In the following discussion emphasis will be placed on immobile incompatible trace elements and the Nd and Sr isotopic systematics. First we examine the compositional variations within each magma type and their implications of low-pressure differentiation. After that we will evaluate the petrogenetic relationships between the rock types and possible parental magma compositions and address processes that could have led to generation of the magma series. The importance of depleted and enriched mantle sources, magmatic plumbing systems, lithospheric terrane boundary and lithospheric thinning zones are also discussed and, eventually, a tectono-magmatic model for the Vestfjella CFBs is presented.

## Low-pressure differentiation of the magma types

The CT1, CT2 and CT3 suites record wide geochemical ranges indicating effective magmatic differentiation. First we discuss the processes responsible for the variations observed in the volumetrically predominant magma types CT1 and CT3. Differentiation of the CT2 tholeites is examined separately. The high- $\varepsilon_{Nd}$  subtype of CT1 shows narrow compositional ranges and its liquid line of descent is thus poorly constrained. Therefore we focus our treatment here on the low- $\varepsilon_{Nd}$  type. The high-*mg*-number CT4 dykes have nearly identical compositions (Luttinen *et al.*, 1998) and are excluded from the following section.

#### Fractional crystallization of CT1 and CT3

The phenocryst assemblages and geochemical variations suggest that early olivine-controlled fractionation and subsequent gabbroic fractionation, when mg-number was <0.55, were largely responsible for differentiation of the CT1 and CT3 suites. Least-squares mixing modelling (Morris, 1984) is hampered by the large scatter in major oxides, specifically  $SiO_2$  at given *mg*-number (Fig. 3). Calculations with several parent-daughter pairs yield feasible results using fractionates consisting of olivine, augite and plagioclase, but the relative proportions of the minerals vary substantially (listings of least-squares models are available from the first author on request). Variations in isotopic and incompatible trace element ratios, however, are not highly susceptible to the exact composition of the fractionating assemblage. A fractionation sequence of olivine (70%) + augite (30%), when *mg*-number is >0.55, and olivine (20%) + plagioclase (50%) + augite (30%), when mg-number is <0.55, accords with the overall major oxide variations of the Vestfjella magma types. Olivine and plagioclase fractionation has little influence on most of the incompatible element ratios. Hence, the bulk D values for isotopic and trace element modelling are largely dependent on the augite abundance, which is 30% in both of the model fractionates.

Differentiation of CT1 and CT3 is modelled in Fig. 13a. The results demonstrate that the decrease of Sm/Nd in CT1 and CT3 cannot be explained by gabbroic fractionation (FC; Fig. 13a) alone and requires some sort of an open magmatic system, such as simultaneous assimilation of crust and fractionation (AFC; DePaolo, 1981*b*). An AFC process could potentially also account for the increase in SiO<sub>2</sub> as recorded by CT1 and CT3 (Fig.

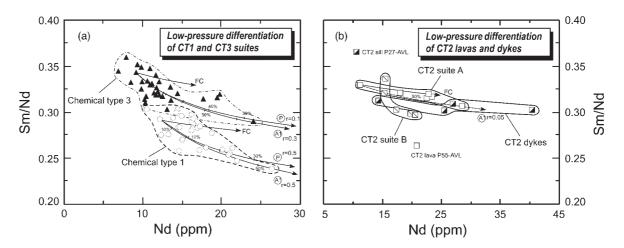


Fig. 13. Modelling of REE variations as a result of low-pressure differentiation of (a) CT1 and CT3 and (b) CT2 magma types. The arrow (FC) indicates 50% fractionation of olivine (20%), clinopyroxene (30%) and plagioclase (50%) with bulk D values for Nd and Sm of 0·12 and 0·18, respectively. AFC models with the same fractionate but different crustal contaminants and r values are indicated. Tick marks show 10% intervals of degree of fractionation. Mixing endmembers for CT1, CT2 and CT3 are P57-AVL, B16-AVL and B403-AVL, respectively. Contaminant A1 is xenolith X4-AVL and P is hypothetical Proterozoic crustal contaminant (Table 4). Data sources: Luttinen *et al.* (1998); this work.

2). We shall next consider possible crustal contaminants of the Vestfjella CFB magmas.

#### Possible crustal contaminants

Little is known about the basement that the Vestfjella CFBs intruded. In southern Vestfjella, however, some lavas and dyke rocks contain felsic crust-derived fragments. Two petrographically similar granitoid xenoliths record  $\varepsilon_{Nd}$  (180 Ma) values of -52 and -53,  $T_{DM}$  model ages (DePaolo, 1981*a*) of 3.27 and 3.24 Ga, and  $\varepsilon_{sr}$ (180 Ma) values of +64 and +59 (Table 3). Geochemically, they resemble the 3.27 Ga (Farrow et al., 1990) Anhalt Granitoid Suite trondhjemites of the Kaapvaal craton (Fig. 1) (see Hunter et al., 1992) with relatively high SiO<sub>2</sub> and Na<sub>2</sub>O/K<sub>2</sub>O, and low FeO, MgO, TiO<sub>2</sub>, Nb, Zr and Y relative to the average uppercrustal composition of Rudnick & Fountain (1995) (Table Similar to Archaean tonalitic-trondhjemitic-4). granodioritic rocks, the HREE concentrations of these xenoliths are notably low (see Martin, 1987). An arkosic sandstone xenolith has  $\varepsilon_{Nd}$  (180 Ma) of -9,  $\varepsilon_{Sr}$  (180 Ma) of +155, a  $T_{\rm DM}$  model age of 1.55 Ga, and plots along the  $\sim 1$  Ga errorchron defined by Maud Belt (Fig. 1) granitoids and gneisses from Heimefrontfiella [ $\varepsilon_{Nd}$ ] (180 Ma) - 2 to -11; Arndt *et al.*, 1991]. The  $\varepsilon_{sr}$  value is also comparable with the median  $\varepsilon_{sr}$  (180 Ma) of +132 reported for felsic Maud Belt rocks (Moves et al., 1993). On the basis of these isotopic similarities the sandstone had probably a dominantly Proterozoic provenance and correlates with the Palaeozoic sedimentary sequence that is exposed at Fossilryggen and Vestfjella, and underlies the CFBs in northern Heimefrontfjella and Kirwanveggen (Fig. 1).

The granitoid xenoliths indicate that at least some of the Vestfjella magmas were emplaced through the Grunehogna craton (Fig. 1) and were contaminated with Archaean crustal material. We point out that the existence of Archaean xenoliths in southern Vestfjella does not necessarily indicate that the craton extends beneath this area. Flood basalt magmas have been recently shown to propagate considerable distances from their source regions on the surface as inflated pahoehoe flows (Self *et al.*, 1996) and also within the crust as dykes (LeCheminant & Heaman, 1989; Elliot *et al.*, 1999). We conclude that the Proterozoic basement as well as the Palaeozoic sedimentary rocks represent other plausible contaminants for the basalts.

#### AFC modelling of CT1 and CT3

The role of AFC in the magmatic evolution of the CT1 and CT3 suites is examined in Fig. 13a, which shows that incorporation of crustal material could account for the decrease in Sm/Nd. Modelling with an Archaean trondhjemitic contaminant (A1) as well as a hypothetical Proterozoic contaminant (P), yields feasible results suggesting a higher rate of assimilation in the case of CT1 (r = 0.5) than CT3 (r = 0.1 and 0.3). For the Archaean and Proterozoic contaminants we have used the geochemical data on xenolith X4-AVL and the average upper crust of Rudnick & Fountain (1995) (Table 4).

In the case of CT3 basalts, the decrease of  $\varepsilon_{\rm Nd}$  along with *mg*-number lends support to operation of an AFC process (Fig. 14a). Models with an Archaean contaminant ( $\varepsilon_{\rm Nd}$  of -53) tend to yield  $\varepsilon_{\rm Nd}$  values that are too low at a given *mg*-number, whereas AFC with a Proterozoic

Sample	Rock type	<sup>147</sup> Sm/ <sup>144</sup> Nd	<sup>143</sup> Nd/ <sup>144</sup> Nd (180 Ma)	ε <sub>Nd</sub>	<i>Т</i> <sub>DM</sub> * (Ga)	<sup>87</sup> Rb/ <sup>86</sup> Sr	<sup>87</sup> Sr/ <sup>86</sup> Sr (180 Ma)	8 <sub>Sr</sub>
X3-AVL-94 X4-AVL-94	granitoid granitoid	0∙0635 0∙0591	0·509834±5 0·509765+7	–51.7 –53.0	3·27 3·24	0·3938 0·3897	0·709965±13 0·709603+9	+ 63·5 + 58·5
X4-AVL-94 X5-AVL-94	sandstone	0.0391	0.509705±7 0.512078±7	9.2	3·24 1·55	0.3868	$0.703003 \pm 3$ $0.716362 \pm 10$	+ 154·6

Table 3: Nd and Sr isotopic composition of CFB-hosted crustal xenoliths from Muren area, Vestfjella

Analytical description is given in Table 2.

\*Depleted mantle model age (DePaolo, 1981a).

contaminant accords fairly well with the data. It should be noted that the Nd content (4·4 ppm) of the xenolith X4-AVL is significantly lower than that of typical Archaean rocks from the Kaapvaal (Hunter *et al.*, 1992). If a higher Nd value of 11 ppm (average of Kaapvaal trondhjemites; Table 4) is used for the contaminant (model A2; Fig. 14b), Archaean crustal material with  $\varepsilon_{\rm Nd}$ of -53 turns out to be unsuitable for AFC modelling of the CT3 basalts.

The isotopic data of CT1 are controversial. The low  $\varepsilon_{Nd}$  values of CT1 indicate a Precambrian LREE-enriched lithospheric component in these rocks. A plot of  $\epsilon_{Nd}$  with mg-number (Fig. 14a) demonstrates that the  $\epsilon_{\rm Nd}$  values are rather scattered and do not show evidence of simultaneous fractionation and assimilation of Archaean crust. In fact, incorporation of Proterozoic crustal material with  $\varepsilon_{Nd}$  higher than -15 would better explain why the  $\varepsilon_{Nd}$  values of some evolved CT1 are relatively high. These samples, however, also exhibit the lowest  $\varepsilon_{sr}$ in contrast to the expected AFC relationship with a Proterozoic contaminant (Fig. 14b). The isotopic data for CT1 thus do not show clear evidence of AFC and are somewhat more accordant with a model in which the degree of crustal contamination is inversely related to the degree of fractionation of the basalt magmas (see Devey & Cox, 1987). In this case the increase in  $SiO_2$ and decrease in Sm/Nd should have been caused by some other differentiation process, such as magma mixing in a periodically refilled and tapped magma chamber (O'Hara, 1977). It is also possible that the isotopic variations of primitive CT1 were large enough to mask isotopic evidence of AFC in the daughter magmas.

We conclude that the relatively high SiO<sub>2</sub>, low Sm/ Nd and negative  $\varepsilon_{Nd}$  values of evolved CT3 reflect an AFC process. On the basis of Nd isotopic evidence, the most likely contaminant is crustal material derived from the Proterozoic Maud Belt. Available isotopic data for the CT1 tholeiites are contradictory. On the basis of coupled increase in SiO<sub>2</sub> and pronounced decrease in Sm/Nd, AFC was probably involved in their evolution, although the isotopic compositions do not show clear AFC relationships. The high-mg-number CT1 lavas already exhibit the diagnostic features of CT1 with distinctly low  $\epsilon_{\rm Nd}$  and a geochemical affinity to crustal material.

#### Differentiation of CT2

As a whole, CT2 shows minor decrease in Sm/Nd consistent with simple gabbroic fractionation or AFC with low rate of assimilation (Fig. 12b). The CT2 suite A and suite B lavas and the CT2 dykes, however, include primitive, high-mg-number samples as well as evolved samples (Fig. 13a) and seem to record at least three separate batches of magma, each of which fractionated differently. The sill rock with  $\varepsilon_{Nd}$  of +7.6 could be considered as a separate subtype of intrusive CT2, but because evolved, isotopically similar rocks have not been found, the sill rock is excluded from this part of discussion. On the basis of its isotopic composition, the lava flow P55-AVL at 800 m level is included here in suite A in spite of its stratigraphic position and high SiO<sub>2</sub> (Fig. 6, Table 1).

The CT2 dyke rocks with high and low mg-number have similar positive  $\varepsilon_{Nd}$  values, suggesting that AFC was not operating (Fig. 13a). On the other hand, the decrease in  $\varepsilon_{Nd}$  from +0.3 to -2.2 in the CT2 suite A and from -6.7 to -7.5 in the CT2 suite B with decreasing mgnumber indicates that AFC probably played a role in their differentiation. An AFC model with low rate of assimilation (r = 0.05) and an Archaean contaminant akin to average Kaapvaal trondhjemite (Table 4) with  $\epsilon_{Nd}$  (180 Ma) of -53 can successfully explain the isotopic variations and the lack of pronounced changes in Sm/ Nd of suite A (Figs 12b and 13a). A similar model could also account for the isotopic variations of suite B. The role of crustal contamination in generating three isotopically different suites of CT2 (Fig. 14a) is a subject of further discussion.

#### Generation of different magma series

The chemical and Nd and Sr isotopic variations of the Vestfjella magma types evidence crustal contamination

X4-AVL	AGS	UC
73.64	72.85	66.00
0.15	0.22	0.50
15.34	15-21	15.20
1.26	1.45	4.50
0.02	0.02	0.08
0.68	0.38	2.20
2.61	1.95	4.20
5.33	5.12	3.90
1.26	2.15	3.40
0.03	0.06	0.40
n.d.	0.74	n.r.
100.32	99-43	100.38
383	401	550
27	68	112
568	509	350
76	131	190
4	2.8	11
1.85	7.0	25
1.05	10.0	22
0.08	n.r.	10.7
0.13	n.r.	2.8
13.16	13.81	30
17.41	20.74	64
1.47	2.86	7.1
4-41	11.42	26
0.49	1.80	4.5
1.10	0.50	0.9
0.31	2.97	3.8
0.04	1.80	0.64
0.21	0.38	3.5
0.04	n.r.	0.80
0.10	1.12	2.3
0.02	n.r.	n.r.
0.10	0.73	2.2
0.02	n.r.	0.32
	$\begin{array}{c} 73.64 \\ 0.15 \\ 15.34 \\ 1.26 \\ 0.02 \\ 0.68 \\ 2.61 \\ 5.33 \\ 1.26 \\ 0.03 \\ n.d. \\ 100.32 \\ 383 \\ 27 \\ 568 \\ 76 \\ 4 \\ 1.85 \\ 1.05 \\ 0.08 \\ 0.13 \\ 13.16 \\ 17.41 \\ 1.47 \\ 4.41 \\ 0.49 \\ 1.10 \\ 0.31 \\ 0.04 \\ 0.21 \\ 0.04 \\ 0.10 \\ 0.02 \\ 0.10 \end{array}$	73.64       72.85         0.15       0.22         15.34       15.21         1.26       1.45         0.02       0.02         0.68       0.38         2.61       1.95         5.33       5.12         1.26       2.15         0.03       0.06         n.d.       0.74         100.32       99.43         383       401         27       68         568       509         76       131         4       2.8         1.85       7.0         1.05       10.0         0.08       n.r.         0.13       n.r.         13.16       13.81         17.41       20.74         1.47       2.86         4.41       11.42         0.49       1.80         1.10       0.50         0.31       2.97         0.04       1.80         0.21       0.38         0.04       n.r.         0.10       1.12         0.02       n.r.         0.10       0.73

Table 4: Major (wt %) and trace element (ppm) composition of possible crustal contaminants

X4-AVL is basalt-hosted xenolith from Muren, Vestfjella. AGS and UC are a 3.27 Ga Anhalt Granitoid Suite trondhjemite (average of eight samples; Hunter *et al.*, 1992) and average upper crust of Rudnick & Fountain (1995), respectively. n.d., not determined; n.r., not reported.

associated with fractional crystallization. A key question in regard to the origin of the four magma series is whether they could represent differentiates of a common parental magma.

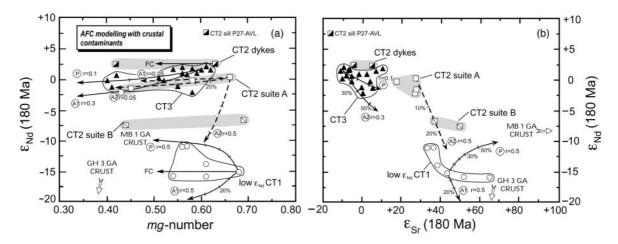
#### Parental magma composition

Few of the analysed basalts could be safely regarded as uncontaminated mantle-derived material. On the basis of their OIB-like geochemistry, positive  $\varepsilon_{Nd}$  (+3.6) and low  $\varepsilon_{sr}$  (+1), the CT4 dykes (Luttinen *et al.*, 1998) are not considered to be crustally contaminated. They are too enriched in incompatible elements, however, to represent parental magma for the other three basalt types. The high  $\epsilon_{Nd}$  (+7.6), low  $\epsilon_{Sr}$  (-16) and MORB-like incompatible element signature of the CT2 sill sample P27-AVL suggest that it may closely correspond to an uncontaminated asthenosphere-derived parental magma type. On the basis of its moderate *mg*-number (0.61), it hardly represents unfractionated primary magma. Correction for  $\sim 30\%$  of olivine fractionation (by adding olivine) yields a feasible primitive magma composition that could have been in equilibrium with a mantle peridotite. Concentrations of the incompatible elements have been corrected accordingly to generate a hypothetical parental magma for petrogenetic modelling (Table 5). The CT2 sill sample exhibits positive Ba, K and Sr anomalies in Fig. 5. Hydrothermal alteration may have affected the LILE contents of this rock sample and, consequently, the composition of the hypothetical CT2 parent is poorly constrained with respect to mobile elements. None the less, it is possible that the high Ba/Nb and Sr/Nd ratios reflect a primary magma composition resembling LILEenriched MORB from the Southwest Indian Ocean Ridge (Le Roex et al., 1989; see also Luttinen et al., 1998).

#### Crustal contamination

Comparison of CT1, CT2 and CT3 demonstrates that if a common parental magma type is assumed for them, each magma series requires a different kind of crustal contaminant to explain, for example, the different combinations of Nb, Ta, P and Ti anomalies shown in Fig. 15.

In the case of CT1, the key question is the origin of the common CT1 signature, which is characterized by features that are typical of crustal material. A similar signature typifies low-Ti CFBs world-wide and the role of crustal contamination in the origin of CFBs has remained a major controversy. Arndt et al. (1993) pointed out the advantages of an assimilating RTF (replenishedtapped-fractionating) magma system (O'Hara, 1977) compared with a simple AFC process (see Hergt et al., 1991) and showed that low-Ti CFBs could be generated from MORB-like parental magmas in a periodically refilled magma chamber. The remaining problematic aspect of crustal contamination models is that they invariably suggest a large amount (20-30 wt %) of crustal component in the low-Ti CFBs (e.g. Arndt et al., 1993; Peng et al., 1994; Luttinen et al., 1998). In our mind, this is difficult to reconcile with the overall homogeneity and



**Fig. 14.** Modelling of  $\varepsilon_{Nd}$  and (a) *mg*-number and (b)  $\varepsilon_{Sr}$  variations as a result of low-pressure differentiation of CT1, CT2 and CT3 magma types. The arrow (FC) in (a) indicates 50% fractionation of olivine (20%), clinopyroxene (30%) and plagioclase (50%) with bulk *D* values for Nd of 0·12 and 0·94, respectively. AFC models with the same fractionate but different crustal contaminants and *r* values are indicated. Tick marks show 10% intervals of degree of fractionation. Mixing endmembers for CT1, CT2 and CT3 are the same as in Fig. 13 apart from  $\varepsilon_{Sr}$  of CT3, for which a representative value of -10 has been used. Contaminants A1 and A2 represent Grunehogna craton (GH 3 Ga crust), whereas P is Proterozoic Maud Belt crust (MB 1 Ga crust). A1 is xenolith X4-AVL ( $\varepsilon_{Nd} - 53$ ,  $\varepsilon_{Sr} + 63$ ), A2 is isotopically similar, average Archaean trondhjemite from Kaapvaal, Africa, and P ( $\varepsilon_{Nd} - 8$ ,  $\varepsilon_{Sr} + 132$ ) is a hypothetical Proterozoic crustal contaminant (Tables 3 and 4). Data sources: Luttinen *et al.* (1998); this work.

large volume of low-Ti CFB suites and the lack of associated silicic magmatism.

The most plausible environment for extensive contamination and efficient homogenization to occur is in large magma chambers at the base of the crust, where primitive, high-density precursors of low-Ti CFBs are likely to be trapped (Cox, 1980). The lower continental crust is chemically similar to low-Ti CFBs (Rudnick & Fountain, 1995) and bulk assimilation of such material is not efficient in modifying the composition of basalt magmas. Prolonged underplating of basalt magmas, however, could have led to substantial melting of the lower crust and bimodal magmatism (e.g. Huppert & Sparks, 1988). A world-wide and well-documented example of voluminous continental bimodal magmatism is provided by the Proterozoic rapakivi associations (e.g. Haapala & Rämö, 1992). Despite the apparently favourable conditions for major mixing processes to occur, effective mixing of the basaltic and granitic magmas in the rapakivi systems has been limited (Rämö, 1991). The rare examples of interaction between the silicic and mafic magmas demonstrate that magma mingling was the predominant process as a consequence of the contrasting physicochemical characteristics of the mixed endmembers (Salonsaari, 1995).

Assuming, however, that bimodal magmatism and effective magma mixing was required to generate the CT1 lavas and other low-Ti magma types of the Karoo province, some of the low-density anatectic melts should presumably have avoided mixing with the basalts. It is difficult to explain the absence of eruptive and intrusive manifestations of coeval and voluminous silicic magmatism associated with the low-Ti lavas of Dronning Maud Land and southeast Africa. The voluminous Karoo rhyolites of the Lebombo Monocline post-date the main stage of CFB magmatism and they have been interpreted to record remelting of underplated Jurassic (low-Ti?) CFBs (Harris & Erlank, 1992; Cox, 1993).

Finally, a well-adjusted mixing process would be required to explain why all the CT1 basalts include a large and practically constant amount of crustal component as demonstrated by the similar size of the negative Ti anomalies (Fig. 15). Such conditions would characterize an ideal steady state of an assimilating RTF system (O'Hara, 1977; Arndt et al., 1993), but the randomness expected in natural systems should have led to compositional gradation between the mixed endmembers as a result of occasional tapping and eruption of moderately contaminated magmas (see Cox, 1988). This is strongly contrasted by the common CT1 signature of the lavas with *mg*-number above 0.65. Such primitive basalts could hardly represent steady-state magma compositions. In sum, generation of low-Ti CFB types, such as CT1, by crustal contamination can be numerically modelled with a combined AFC and RTF process, but such a model is undermined by some critical general geological arguments.

The heterogeneous CT2 lavas and dykes record a strongly differentiated suite of CFB magmas that were probably derived from the same MORB-like parental magma type. Here we focus on the marked difference

																					l
	ЧL	Та	ЧN	Се	Pr	Nd	Ь	Hf	Zr	Sm	Eu	μ	Gd	Тb	Но	≻	Υb	Lu	Sr*	ESr	5 <sub>Nd</sub>
Contaminant1	32	6.3	135	135 470	35	157	5673	32	1160	23	5.1	24580 12.7	12.7	1.6	1.0	20	1.7	0.23	3500	+ 100	-25
Parental CT1 <sup>‡</sup>	0.24	0.24 0.17	2-93	2.93 12.28	1.66	7.77	500	1.74	67	2.28	0.83	4494	2.52	0.43	0.52	13	1·08	0.16	176	6+	-2.5
Parental CT2 <sup>‡</sup>	0.16	0.12	2.19	8-55	1.37	7.37	458	1.72	60	2.70	1.0	7428	3.04	0-55	0.64	15.4	1.21	0.18	133	- 16	+ 7.5
20% AFC of																					
CT1§	1.19	1.19 0.39	7.40	7.40 28.4	3.04	14-00	781	3.06	116	3.48 1.18	1.18	6271	3.50 0.58	0.58	0.68	16-8	1.39	0.21	317	+ 37	- 9.5
30% AFC of																					
CT2§	0.95	0.95 0.31	6.16	6.16 22.7	2.74	14-02	780	3.17	112	4.36	4.36 1.54	11084	4.61 0.82	0.82	0.93	22.4	1.76	0.26	268	+ 18	-0-7
*Sr concentrations of lamproite and parental CT2 were modified to fit the immobile element modelling (see text for details). *Sr concentrations of lamproite and parental CT2 were modified to fit the immobile element modelling (see text for details). *Hypothetical parental magmas were calculated assuming that the average of high-e <sub>vid</sub> CT1 corresponds to 50% of olivine fractionation of the parental CT1 and that CT2 sill rock P27-AVL is the result of 30% of olivine fractionation of parental CT2 (see Fig. 15). §AFC models were calculated using the equations of DePaolo (1981 <i>b</i> ) assuming olivine fractionation and bulk <i>D</i> values of 0.01 for all elements (0.02 for Ti) (see Rollinson, 1993). Rates of assimilation used for CT1 and CT2 were 0.1 and 0.05, respectively. Model results and measured data for CT1 and CT2 are compared in Figs 15 and 16.	ations o contami l parent ock P27 ock P27 33). Rat 6.	of lampr tal mag 7-AVL is alculate es of as	oite and om lithc mas we the res ed using ssimilati	d paren ospheric are calc sult of 3 g the ec ion use	ttal CT2 c mantle ulated i 10% of c 10% of c 104tions d for C1	CT2 were modified to fit the immobile element modelling (see text for details). antle of Grunehogna craton. Composition is the average of Rock (1991). ted assuming that the average of high- $\varepsilon_{\rm Nd}$ CT1 corresponds to 50% of olivine free of olivine fractionation of parental CT2 (see Fig. 15). tions of DePaolo (1981 <i>b</i> ) assuming olivine fractionation and bulk <i>D</i> values of 0. Torney of DePaolo (1981 <i>b</i> ) assuming olivine fractionation and bulk <i>D</i> values of 0. Torney of CT1 and CT2 were 0.1 and 0.05, respectively. Model results and measured dated the text of text of the text of the text of the text of the text of text of the text of text of the text of t	nodifie unehog ng that raction (' CT2 we	1 to fit 1 na crat the ave ation o 1981 <i>b</i> ) <i>i</i> re 0.1 a	the imr on. Cor srage o f paren assumii nd 0.0{	mobile npositi f high- tal CT2 ng oliv oliv 5, respe	elemer on is tl <sup>ENd</sup> CT1 ? (see F 'ine frau	nt mode ne aver l corres ig. 15). /. Mode	elling (state of the second state of the second state of on and on and the second state of the second stat	see tex Rock ( to 50% I bulk <i>I</i> ts and ts and	t for de 1991). 6 of oli 7 value measu	etails). vine fra s of 0.0	actiona 11 for a	tion of all elem T1 and	the pa ients (C CT2 au	rental ( .02 for	CT2 were modified to fit the immobile element modelling (see text for details). antle of Grunehogna craton. Composition is the average of Rock (1991). ted assuming that the average of high- $\varepsilon_{Nd}$ CT1 corresponds to 50% of olivine fractionation of the parental CT1 and of olivine fractionation of parental CT2 (see Fig. 15). tions of DePaolo (1981 <i>b</i> ) assuming olivine fractionation and bulk <i>D</i> values of 0.01 for all elements (0.02 for Ti) (see or CT1 and CT2 were 0.1 and 0.05, respectively. Model results and measured data for CT1 and CT2 are compared in

Table 5: Lithospheric mantle contamination modelling of CT1 and CT2; immobile incompatible elements (ppm)

NUMBER 8

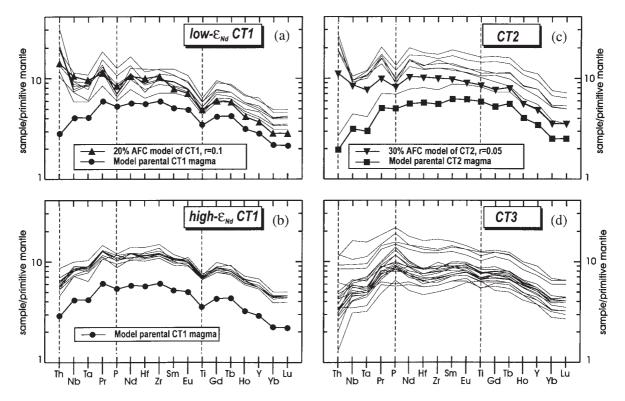


Fig. 15. Primitive mantle-normalized immobile incompatible element diagrams of (a) low- $\varepsilon_{Nd}$  CT1, (b) high- $\varepsilon_{Nd}$  CT1, (c) CT2 and (d) CT3 rocks. Lithospheric contamination models and hypothetical primary magmas of CT1 and CT2 (Table 5) are shown. Lithospheric contaminant is average lamproite of Rock (1991) (Table 5). Dashed lines point to diagnostic anomalies of CT1, CT2 and CT3.

between the high-mg-number CT2 lavas and the hypothetical CT2 parental magma, which is demonstrated first and foremost by the lower  $\varepsilon_{\rm Nd}$ , (P/Nd)<sub>n</sub> and higher (Th/Ta)<sub>n</sub> of the lavas (Figs 12 and 15). Considering the relatively low SiO<sub>2</sub> and (La/Sm)<sub>n</sub>, and high (Ti/Gd)<sub>n</sub> and mg-number recorded by the suite A lavas, these differences point to a contaminant that (1) was not particularly high in SiO<sub>2</sub>, (2) was strongly enriched in LREE but not in HREE relative to parental CT2 and (3) exhibited a strong negative P anomaly but not a corresponding negative Ti anomaly. These features are not considered to be typical of crustal-derived material and the wide compositional ranges of the high-mg-number CT2 are more consistent with a mafic incompatible element enriched contaminant.

The similarity of the incompatible element signatures of the suite A and B lavas (Figs 5b and 15) implies that the cause of high  $SiO_2$  of suite B was not directly linked to the development of the trace element patterns. The high  $SiO_2$  strongly suggests a crustal component in suite B and its isotopic compositions can be modelled reasonably well by crustal contamination of primitive CT2 suite A magmas with an Archaean trondhjemitic contaminant (Table 4) and assimilation rate of 0.2 (Fig. 14). The CT3 lavas show some affinities to CT1 and the other Karoo low-Ti CFBs but also to CT2. Geochemical indications of significant crustal contamination, such as high SiO<sub>2</sub>, (La/Sm)<sub>n</sub>, (Th/Ta)<sub>n</sub> and  $\varepsilon_{sr}$ , and low (P/Nd)<sub>n</sub>, (Ti/Gd)<sub>n</sub> and  $\varepsilon_{Nd}$ , are lacking in the high-*mg*-number CT3 lavas. In terms of immobile incompatible elements and the Nd and Sr isotopic ratios, the high-*mg*-number CT3 lavas show affinities to MORB and resemble the hypothetical parental CT2 magma (Figs 14 and 15). On the other hand, relatively high (P/Nd)<sub>n</sub> of CT3 is not a typical result of crustal contamination and suggests that the parental magma of CT3 was somewhat different from that of CT2.

Although crustal contamination influenced the lowpressure fractionation of the CT1, CT2 and CT3 magmas it does not seem to have been responsible for the generation of the different magma series. An important common feature of the Vestfjella magma types is that each of them includes rocks with notably high *mg*-number, indicating that relatively primitive magmas already exhibited the diagnostic incompatible element and isotopic signatures of CT1, CT2, CT3 and CT4. We consider that the magma types possessed their salient features essentially before emplacement into the crust.

#### Lithospheric mantle contamination

A possible mechanism that could have led to contamination of basalt magmas before they reached the crustal level is mixing with small-degree melts of lithospheric mantle (McKenzie, 1989; O'Brien et al., 1991; Ellam & Cox, 1991; Ellam et al., 1992; Foley, 1992; Cadman et al., 1995). Subcontinental lithospheric mantle may include domains characterized by abundant veins rich in clinopyroxene and mica and/or amphibole. These veins typically have lower solidus temperatures than their wall-rock peridotite and incipient melting of such veined lithosphere is likely to produce highly alkaline melts, such as lamproites (McKenzie, 1989; Foley, 1992). Melting modelling of a vein-derived component would be complicated because of uncertainties with respect to the mineralogical and chemical composition of the vein assemblage, solid-solution melting reactions and dissolution of the wall-rock minerals (Foley, 1992). Instead, we have used the average lamproite of Rock (1991) to approximate a hypothetical vein-derived component.

Previously, Luttinen et al. (1998) speculated on whether the CT1 basalts could have been generated by mixing of MORB-like asthenosphere-derived magmas and lithosphere-derived lamproitic material. Although such a lithospheric mantle contamination model could explain many of the characteristic features of CT1, including the low Ti/Zr, it would, at the same time, lead to a strong enrichment of the highly incompatible elements and, for example, notably higher (Nb/Y)<sub>n</sub> than those recorded by primitive CT1 (see Luttinen et al., 1998). We conclude that lithospheric mantle contamination is an improbable explanation for the common CT1 signature and the origin of the CT1 magma type. Detailed examination of the trace element characteristics of the CT1 and CT2 suites shows that such a process presumably played an important role in generating subtypes of CT1 and CT2, however.

#### Trace element constraints on lithospheric mantle contamination

Compared with the CT2 sill, the CT2 lavas have higher LREE,  $(Th/Ta)_n$  and  $\varepsilon_{Sr}$ , and lower  $(P/Nd)_n$  and  $\varepsilon_{Nd}$  values (Figs 14 and 15). Importantly, a similar result is obtained when comparing the high- $\varepsilon_{Nd}$  and low- $\varepsilon_{Nd}$  CT1 rocks. In spite of their different overall compositions, the CT1 and CT2 suites seem to have incorporated the same enriched component, which influenced the highly incompatible element ratios but caused minor changes in the major elements and moderately incompatible element ratios, such as  $(Ti/Gd)_n$ . Here we examine whether lithospheric mantle contamination could have produced the low- $\varepsilon_{Nd}$  subtypes of CT1 and CT2. The high- $\varepsilon_{Nd}$  CT1 rocks have relatively low *mg*-number (<0·6). For the modelling, a hypothetical parental high- $\varepsilon_{Nd}$ 

CT1 magma has been calculated in a similar fashion as previously was done in the case of CT2 (Table 5).

In the lithospheric mantle contamination model (Fig. 15), the hypothetical parental CT1 and CT2 magmas crystallized olivine as they were emplaced into a subcontinental lithosphere and mobilized and entrained highly enriched melts from the veined wall-rock mantle. The process has been modelled using the AFC equations of DePaolo (1981*b*). The bulk *D* values were assumed to be 0.01, with the exception of the value of Ti, for which 0.02 was used in accordance with olivine fractionation (Rollinson, 1993). AFC modelling with a contaminant similar to the average lamproite of Rock (1991) demonstrates that lithospheric contamination with low assimilation rates could account for the different immobile element patterns of high- $\varepsilon_{Nd}$  and low- $\varepsilon_{Nd}$  subtypes of both CT1 and CT2 (Fig. 15).

#### Isotopic constraints on lithospheric mantle contamination

Lithospheric mantle contamination would strongly have influenced also the isotopic compositions if the lamproitic component had been derived from an ancient reservoir with low Sm/Nd and high Rb/Sr relative to chondrite. Archaean crust and mantle ages for the Kaapvaal craton show long-term physical coupling of mantle and crust over a 3 Gy period (Pearson et al., 1995). The xenolith data (Table 3) thus suggest that Archaean, cratonic lithospheric mantle may well extend beneath or at least close to the Vestfiella region (Fig. 1). Kimberlite-hosted peridotite and mineral inclusions provide evidence for the composition of the Kaapvaal lithospheric mantle (Menzies & Murthy, 1980; Richardson et al., 1984; Walker et al., 1989; Pearson et al., 1995) and give the best estimate of the cratonic lithosphere in western Dronning Maud Land. The isotopic variations indicate considerable heterogeneity with, for example,  $\varepsilon_{Nd}$  at 180 Ma ranging from -4 to -32 (Richardson *et al.*, 1984). As a consequence of successive infiltration of depleted Archaean Kaapvaal mantle by various asthenosphere-derived melts, the Kaapvaal-Grunehogna lithospheric mantle probably records intermediate compositions between DM, EM1 and EM2 mantle components (see Zindler & Hart, 1986) as demonstrated by the enriched mantle array of Menzies & Murthy (1980) (Fig. 16b; Menzies, 1990).

We assume that the purported lamproitic component showed an isotopic affinity to the enriched mantle array. Trace element model parameters can be used to constrain the isotopic composition of this component: a contaminant with  $\epsilon_{Nd}$  (180 Ma) of -25 complies reasonably well with the trace element AFC modelling of the low- $\epsilon_{Nd}$  CT1 as well as the CT2 lavas (Fig. 16a). Modelling of the Sr isotopic systematics is more complicated because of uncertainties with respect to possible alteration effects. This problem is particularly significant in the case of

CT2 because a single sample was used for deriving the parental magma composition (Table 5). The CT2 sill rock sample shows a positive Sr anomaly with Sr/Nd of 21, which is substantially higher than the corresponding ratio for average depleted MORB ( $\sim 12$ , Sun & McDonough, 1989). Here we assume that before alteration the sill rock had a somewhat lower Sr/Nd of 18 corresponding to Sr content of 133 ppm for the hypothetical primary CT2 magma. A lamproitic contaminant with  $\varepsilon_{sr}$  (180 Ma) of +100 shows isotopic affinity to the enriched mantle array and would reasonably well explain the Sr isotopic ranges of both CT1 and CT2 suites (Fig. 16b). It would be possible to explain the isotopic compositions of the CT2 suite B lavas by lithospheric mantle contamination alone. The high SiO<sub>2</sub> of suite B lavas, however, strongly suggests that these rocks contain a significant crustalderived component. The Sr concentration (3500 ppm) and thus the Sr/Nd ratio (22) used for the model are clearly higher than those of the average lamproite (1250 ppm and 8, respectively), but Sr contents of >4000 ppm and Sr/Nd ratios up to 24 have been reported for lamproites and related rock types (e.g. Rock, 1991; Larsen & Rex, 1992).

In the isotopic modelling, the Sr concentrations of the CT2 parental magma and the lamproite were adjusted so that the same contaminant could explain the variations of both CT1 and CT2 magma series. Overall, the modelling serves to illustrate the feasibility of lithospheric mantle contamination in explaining not only the chemical (Fig. 15) but also the isotopic ranges of the CT1 and CT2 suites (Fig. 16). The composition of the vein-derived lithospheric mantle contaminant could well have been even more enriched in incompatible elements than lamproites, which have been considered as mixtures of the vein and wallrock mantle material (e.g. Foley, 1992). In that case our calculations would overestimate the lithospheric mantle contribution to the CT1 and CT2 magmatism.

## Sublithospheric and lithospheric mantle sources

On the basis of their positive  $\varepsilon_{Nd}$  values and chemical affinities to MORB and OIB, CT2 and CT4 probably had asthenospheric sources. The isotopic and chemical MORB affinities of the CT3 lavas are contrasted by their notably high LILE contents (Fig. 5). This selective LILE enrichment is difficult to reconcile with crustal or lithospheric mantle contamination, which is expected to lead to concomitant increase in LREE and LILE concentrations. The chemical data for CT3 thus point to a strongly LILE-enriched source composition.

The origin of low-Ti CFBs, such as CT1, has been often ascribed to lithospheric mantle sources (e.g.

Hawkesworth et al., 1984; Sweeney & Watkeys, 1990). In the same vein, Luttinen et al. (1998) proposed a lithospheric source for the low- $\varepsilon_{Nd}$  CT1. The lithospheric mantle contamination model, however, provides a feasible explanation for derivation of the low- $\varepsilon_{Nd}$  CT1 type from high- $\varepsilon_{Nd}$  type CT1 parental magmas. Nevertheless, it cannot account for the invariably high  $SiO_2$  and the negative Ti anomalies, that is, the common signature of CT1, which seem to require a lithospheric source for this magma type (see Turner & Hawkesworth, 1995). It is possible to invoke a layered lithospheric mantle with a deeper source for high- $\varepsilon_{Nd}$  type parental magmas of CT1 and an overlying veined domain where these magmas were contaminated to produce low- $\varepsilon_{Nd}$  CT1. The data for high- $\varepsilon_{Nd}$  CT1 may thus reflect the characteristics of the primary source of CT1 magmas.

Comparison of the high- $\varepsilon_{Nd}$  CT1 and CT3 shows them to have rather similar incompatible element signatures (Figs 5, 11 and 15). Here we discuss possible links between the sources of these low-Ti magma series. There has been frequent speculation on possible connections between the sources of Gondwana low-Ti CFBs and subduction processes (e.g. Hergt *et al.*, 1991; Brewer *et al.*, 1992). For example, Duncan (1987) has shown that the low-Ti tholeiites of south Lebombo plot within the calc-alkaline fields of the Ti–Zr–Y and Ti–Zr–Sr diagrams of Pearce & Cann (1973). The chemically similar CT1 also exhibits these affinities to calc-alkaline volcanic rocks, whereas CT3 would plot mainly in the low-K tholeiite fields of the above-mentioned diagrams.

In their discussion on the origin of the Gondwana low-Ti CFBs, Hergt et al. (1991) calculated that a geochemically suitable low-Ti source with a crust-like trace element signature could have been generated by incorporation of a minor quantity of sediment material in depleted Gondwanan upper mantle during subduction. Such a source would explain the crust-like features of the parental CT1 lavas. In volcanic arc systems, the composition of the subduction-contaminated mantle wedge can vary substantially depending on, for example, the amount of subducted sediment and whether the slabderived component is generated by dehydration or partial melting. Previously, Luttinen et al. (1998) proposed that the selective LILE enrichment of CT3 reflects a subduction-influenced source of this magma type. The enrichment of P (Fig. 15) lends support to this interpretation because subduction-liberated fluids and melts may effectively increase LILE, LREE and also P contents in the overlying mantle wedge and, at the same time, cause only small changes in most HFSE and middle to heavy REE concentrations (Tatsumi et al., 1986; Pearce & Peate, 1995; Pearce et al., 1995).

In conclusion, the Jurassic mantle source regions of CT1 and CT3 and those of island arc basalts have marked similarities. In the case of CT1 and CT3, such

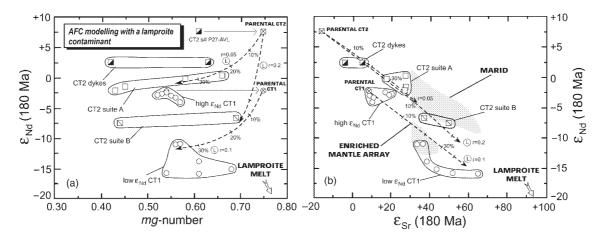


Fig. 16. Modelling of initial  $\varepsilon_{Nd}$  and (a) *mg*-number and (b)  $\varepsilon_{Sr}$  variations as a result of lithospheric contamination of CT1 and CT2 magma types. AFC models with primary CT1 and CT2 (Table 5) and a lamproitic contaminant (L) (Table 5) ( $\varepsilon_{Nd} - 25$ ,  $\varepsilon_{Sr} + 100$ , *mg*-number 0.83) are indicated. Tick marks indicate 10% intervals of degree of fractionation. Shaded fields in (b) show compositions of the Kaapvaal enriched mantle array (Menzies & Murthy, 1980) and MARID (mica–amphibole–rutile–ilmenite–diopside) xenoliths (Kramers *et al.*, 1983; Waters, 1987). Data sources: Luttinen *et al.* (1998); this work.

a source could have been generated in association with subduction-related contamination of the upper mantle during two distinct periods. First, the Proterozoic lithospheric terrane of Dronning Maud Land is an accretionary volcanic arc complex (Groenewald et al., 1995). It is possible that the associated upper mantle could have been variably contaminated by subducted sediments and slab-derived fluids and melts to produce a laterally heterogeneous source for CT1 and CT3. Second, researchers have also speculated on whether Jurassic subduction along the proto-Pacific margin of Gondwana could have influenced the break-up magmatism (e.g. Cox, 1978; Elliot, 1990). The Karoo igneous province, however, was separated by >2000 km from the active Jurassic subducting margin. An extremely flat subduction angle would have been required to extend subduction-related mantle contamination to the source regions of CT1 and CT3 (cf. Murphy et al., 1998; Dalziel et al., 1999). We prefer that the assumed subduction-contamination preceded the break-up magmatism and was not directly linked to it. An LILE- and LREE-enriched mantle domain could have been restored in the Proterozoic lithosphere or, perhaps more likely, in the sublithospheric thermal boundary layer of the Gondwanan upper mantle (see Anderson, 1994).

#### Magmatic plumbing systems and lithospheric thinning zones

Stratigraphic alternation of the compositionally distinctive CT1, CT2 and CT3 magma series strongly implies coexistence of three different magmatic plumbing systems. On the basis of Nd isotope compositions, CT1 and CT2 were contaminated with Archaean and the CT3 lavas by Proterozoic lithospheric material. Accordingly, lavas with notably low  $\varepsilon_{Nd}$  values occur only in the north Vestfjella and palaeoflow direction measurements indicate that they flowed mainly from north to south, i.e. from an area of Archaean lithosphere (Fig. 17). Lavas with near-chondritic  $\varepsilon_{Nd}$  are predominant toward the southern Vestfjella (Fig. 17). On the basis of flow directions, their source regions were located east and south of Vestfjella. Overall, there is a good correlation between the apparent source regions of the CFBs with high and low  $\epsilon_{\!\scriptscriptstyle Nd}$  and the distribution of the Archaean and Proterozoic terranes of the Dronning Maud Land lithosphere (Fig. 17). This is accordant with two main magmatic plumbing systems; one within the Grunehogna craton and the other within the Maud Belt. Development of two active lithospheric thinning zones probably had a controlling influence on these systems.

The Lebombo magmatism of Africa has been ascribed to thinning in the Mozambique-Weddell Sea zone (Figs 1 and 17) (e.g. Cox, 1992). Similarly to the Lebombo lavas, the Vestfiella CFBs are located close to the present continental margin and were spatially associated with the Mozambique-Weddell Sea zone. Another major rifting regime related to the Gondwana break-up in western Dronning Maud Land is expressed by the Jutulstraumen-Pencksökket (Jutul–Penck) trough (e.g. Marsh, 1991) (Fig. 17) along which alkaline magmatism occurred at 180 Ma (Grantham & Hunter, 1991; Harris & Grantham, 1993; Grantham, 1996). Seismic measurements have defined a similar graben-like structure extending to the southwest from Pencksökket, parallel to Heimefrontfjella (Hungeling & Thyssen, 1991). Although extensive in some localities, this 'failed' off-cratonic rifting did not lead to an oceanic spreading centre (Spaeth, 1987), whereas lithospheric

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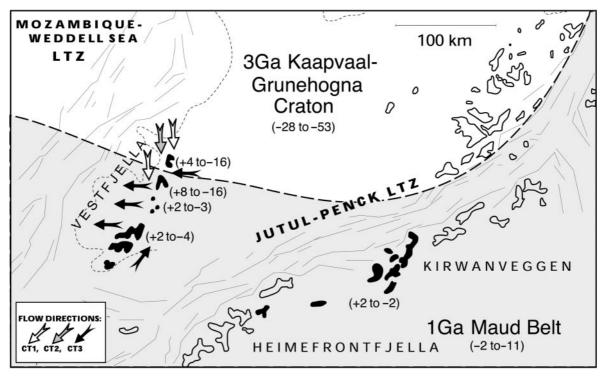


Fig. 17. Distribution of the Jurassic CFBs (black), Archaean (Kaapvaal–Grunehogna) and Proterozoic (Maud Belt) lithospheric terranes, and lithospheric thinning zones (Jutul–Penck LTZ and Mozambique–Weddell Sea LTZ) in western Dronning Maud Land. Arrows indicate predominant palaeoflow directions of CT1, CT2 and CT3 in different areas. Numbers in parenthesis show  $\varepsilon_{Nd}$  (180 Ma) values for CFBs and felsic crustal rocks. Data sources: Carlson *et al.* (1983); Arndt *et al.* (1991); Harris *et al.* (1991); Luttinen *et al.* (1998); this work.

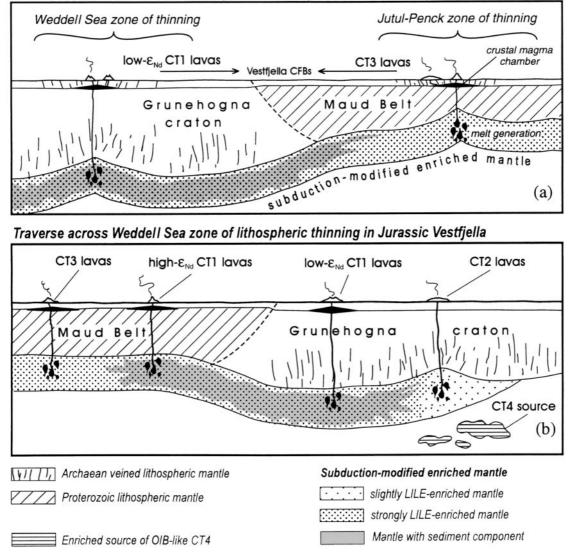
extension in the Mozambique–Weddell Sea zone finally split the Archaean craton and developed into an oceanfloor spreading regime.

The origin of the heterogeneous Vestfjella CFB suite probably involved two major magmatic plumbing systems in areas of lithospheric thinning. Not only may lithospheric thinning zones have permitted the emplacement of magmas; also, the plume-generated thermal anomaly could have been channelled and magma production localized in such areas (Thompson & Gibson, 1991; Gallagher & Hawkesworth, 1992; Ebinger & Sleep, 1998).

#### Model of Vestfjella CFB magmatism

Here we propose a tectono-magmatic model for the Vestfjella CFBs. The essential components of the model include: (1) the Archaean Grunehogna craton with enriched, veined lithospheric mantle roots; (2) the Proterozoic Maud Belt terrane; (3) a variably LREE- and LILE-enriched, subduction-modified (sublithospheric?) upper-mantle source; (4) two lithospheric thinning zones—the Weddell Sea zone within the Archaean craton and the Jutul–Penck zone within the off-cratonic Proterozoic terrane.

The tectono-magmatic model is illustrated in Fig. 18. Plume-lithosphere interactions generated two lithospheric thinning zones, which controlled melting of the hot upper mantle and permitted emplacement of the generated magmas (Fig. 18a). The CT1 primary magmas were generated in an LREE- and LILE-enriched upper mantle. Most of these magmas were emplaced into a cratonic segment of the Weddell Sea lithospheric thinning zone and incorporated incompatible element enriched low- $\varepsilon_{Nd}$  material from veined lithospheric mantle giving the ascending magmas the low- $\varepsilon_{Nd}$  CT1 signature. Some CT1 magmas (high- $\varepsilon_{Nd}$  CT1) were not contaminated by the enriched lithosphere and maintained the CT1 source signature. They either intruded through Proterozoic lithosphere or avoided veined parts of the Archaean lithosphere (Fig. 18b). Occasionally, small batches of melts were tapped from LREE-depleted, slightly LILE-en-18b). These riched upper-mantle sources (Fig. CT2 magmas utilized the cratonic plumbing system and, similar to CT1, were variably contaminated with an Archaean lithospheric component to yield compositionally heterogeneous CT2 lavas. The CT3 magmas were generated in an isotopically depleted mantle source that was notably enriched in LILE as a result of subduction-contamination. These magmas were emplaced



Traverse across two zones of lithospheric thinning in Jurassic Vestfjella

**Fig. 18.** A schematic presentation of Mesozoic CFB magmatism in Vestfjella region, western Dronning Maud Land, Antarctica, illustrated by traverses across two lithospheric thinning zones in Jurassic Dronning Maud Land (a) and along the Weddell Sea zone of thinning (b). Localized melting of upper mantle occurred along lithospheric thinning zones (LTZs). CT1 was generated by melting of LREE- and LILE-enriched mantle within the cratonic Weddell Sea zone (a, b). The high- $\varepsilon_{Nd}$  type primary CT1 magmas included an Archaean LREE-enriched component as they ascended through veined lithospheric mantle and, as a consequence of this lithospheric contamination, most of the erupted basalts recorded the low- $\varepsilon_{Nd}$  CT1 type geochemistry. CT2 primary magmas were generated in a slightly LILE-enriched, LREE-depleted mantle source below the Weddell Sea LTZ (b). Similar to CT1, they were subjected to lithospheric contamination resulting in variable low  $\varepsilon_{Nd}$  in the erupted lavas. CT1 and CT2 were further contaminated with Archaean crustal material. CT3 was derived from strongly LILE-enriched upper mantle. These magmas were emplaced into the Proterozoic lithosphere and ascended to crustal levels without significant contamination. They were contaminated, however, with Proterozoic crust in crustal level magma chambers. The late-stage CT4 dykes may record melting of enriched parts of asthenospheric mantle (b) (see Luttinen *et al.*, 1998). Penecontemporaneous localized melting of heterogeneous upper mantle across an Archaean–Proterozoic lithospheric terrane boundary resulted in interbedded, compositionally distinct lava flows in Vestfjella between the two lithospheric thinning zones.

into the off-cratonic lithospheric thinning zone and were not subjected to lithospheric mantle contamination.

Crustal-level contamination also modified the compositions of CT1, CT2 and CT3 suites, but in general had relatively small impact on their diagnostic trace element and isotopic ratios. Contamination is reflected by the slight decrease in  $\epsilon_{\rm Nd}$  values during fractionation of CT2 and CT3 lavas, whereas isotopic variations of

lithospheric mantle contaminated high *mg*-number CT1 magmas probably masked subsequent AFC relationships.

Between the two major zones of eruptive activity, the CT1, CT2 and CT3 basalts were interbedded (Figs 6 and 18). Towards the later stages of magmatism the predominant type in the cratonic zone of thinning changed to relatively uncontaminated CT2 recorded by the dyke rocks and sills. Emplacement of these MORB-like magmas was associated with occasional tapping of OIB-like compositions (CT4 dykes) from heterogeneous asthenosphere (see Luttinen *et al.*, 1998).

#### Mantle heterogeneity and Karoo magmatism

Our results lend support to the previously presented idea that lithospheric mantle contamination played a key role in Karoo magmatism associated with the Mozambique lithospheric thinning zone. Ellam and coworkers (Ellam & Cox, 1989, 1991; Ellam et al., 1992) have suggested that the high-Ti picrites of Mwenezi area (Fig. 1) record a mixing lineage between asthenosphere-derived parental magmas and a lithospheric mantle component. Our modelling of the CT2 suite suggests that such a process operated also in north Vestfjella. Interestingly, isotopic data on MARID (mica-amphibole-rutile-ilmenitediopside) mantle xenoliths from Kaapvaal (Kramers et al., 1983; Waters, 1987) show affinity to CT2 (Fig. 16). Similar to the Mwenezi picrites and CT2, the MARID xenoliths have been interpreted to record mixing of asthenospheric and lithospheric mantle material (Kramers et al., 1983) as a result of Mesozoic impact of a mantle plume on the base of the Kaapvaal lithosphere (Menzies, 1990; Konzett et al., 1998). The significance of lithospheric mantle contamination is further underlined by the observation that also the low-Ti CT1 suite of Vestfjella seems to include the same overall lamproite-like component as CT2.

We have emphasized that CT1, CT2 and CT3 were derived from different mantle sources. It is possible that these sources actually reflect compositional gradation of variably subduction-contaminated Gondwanan upper mantle. In such a scenario, the source of CT2 represents uncontaminated or slightly LILE-enriched asthenosphere whereas the sources of CT3 and CT1 were successively more contaminated with slab-derived fluids and sediment material. Generation of the Kirwan basalts (Fig. 17) could also be linked to the proposed subduction-contaminated source. Similar to CT1, the Kirwan basalts may record melting of mantle with a subducted sediment component but this source was not influenced by strong LILE enrichment. In accord with Sweeney & Watkeys (1990), the subduction-contaminated sources of the low-Ti magma types were located mainly below the Proterozoic crust but, as the data on low- $\epsilon_{Nd}$  CT1 demonstrate, they also extended below the marginal parts of the Archaean craton. Towards the interior of the craton, the subduction overprint of the upper mantle presumably disappeared in the depleted source regions of high-Ti CFBs. The compositional gaps between the various magma types can be explained by localized thinning and partial melting of laterally heterogeneous upper mantle (Fig. 17).

#### CONCLUSIONS

On the basis of the presented data and discussion on the Jurassic Vestfjella CFBs we summarize the following conclusions.

## Vestfjella magma types and their relationships to other Karoo CFBs

The Jurassic Vestfjella CFBs can be grouped into three main chemical types (CT1, CT2 and CT3) based on immobile incompatible trace element ratios, such as Ti/Zr and Ti/P. The CT1 lavas and dykes are rather typical Gondwana low-Ti tholeiites with a crust-like trace element signature, whereas CT2 resembles MORB and is regarded to be transitional between low-Ti and high-Ti magma types. The CT3 lavas show some affinities to low-Ti CFBs but are different from other Karoo low-Ti types.

The lava succession is cut by abundant dykes and sills, which represent intrusive equivalents to CT1, CT2 and CT3, but also include a distinctive, although minor group of CT4 dykes that represent a high-Ti magma type and show an OIB affinity. On the basis of our stratigraphic reconstruction, the Vestfjella CFB succession can be divided into three major units; the lower CT1-dominated suite, the middle CT3-dominated suite and the upper CT1-dominated suite. CT2 lavas occur as minor interbeds, mainly in the north.

CT1 and CT2 show wide ranges of Nd and Sr isotopic ratios with  $\epsilon_{Nd} (180 \text{ Ma})$  from +8 to -16 and  $\epsilon_{Sr} (180 \text{ Ma})$  from -16 to +65, in contrast to CT3 lavas, which record close chondritic isotopic ratios. The low- $\epsilon_{Nd}$  CT1 and the CT2 dykes and sill rock are chemically and isotopically notably similar to the Sabie River basalts, south Lebombo and Rooi Rand dykes, respectively. The high- $\epsilon_{Nd}$  CT1, the CT2 lavas and the CT3 lavas seem to be confined to Vestfjella. The CT4 dykes resemble some of the high-Ti CFBs of Ahlmannryggen and central Lebombo.

#### Two-stage differentiation

The large isotopic variations of the CT1 and CT2 suites reflect lithospheric mantle contamination. Emplacement of (near) primary CT1 and CT2 magmas into veined cratonic lithosphere triggered incipient melting and led to mixing with vein-derived lamproite-like material, which overprinted the highly incompatible element and isotopic ratios of CT1 and CT2. The CT1 and CT2 lavas were also contaminated by Archaean Grunehogna crust, but this process had a relatively minor effect on their diagnostic incompatible element and isotopic ratios. CT3 records contamination by Proterozoic Maud Belt crust. Low-pressure fractional crystallization of olivine, plagioclase and augite had a controlling influence on the major element geochemistry of the magmas.

#### Importance of lithospheric thinning zones

The generation and emplacement of the CT1, CT2 and CT3 magmas was controlled by the development of two major lithospheric thinning zones associated with the Gondwana break-up in Jurassic Vestfjella and adjacent areas. The CT1 and CT2 magmas utilized plumbing systems developed in the Mozambique–Weddell Sea thinning zone. The magmas that were emplaced into a cratonic segment of this zone were subjected to lithospheric mantle contamination. The CT3 magmas were emplaced into the Jutul–Penck zone of thinning within the Proterozoic Maud Belt lithospheric terrane and were not contaminated with lithospheric mantle material.

#### Depleted and enriched mantle sources

The high- $\varepsilon_{Nd}$  endmembers of CT1, CT2 and CT3 probably closely resemble uncontaminated mantle-derived magmas and indicate three different mantle sources. The CT2 primary magmas were derived from LREE-depleted sources, whereas data on the volumetrically preponderant CT1 and CT3 point to LREE- and LILE-enriched sources. The low-Ti types, CT1 and CT3, were generated as a result of localized melting of variably subduction-contaminated upper mantle. Overall, the depleted source of CT2 and enriched sources of CT1 and CT3 may record compositional gradation generated by subduction-contamination of upper mantle. The OIB-like CT4 dykes probably reflect asthenospheric heterogeneities that were unrelated to the proposed subduction-contamination.

In summary, we argue that the pronounced heterogeneity of the Vestfjella CFB suite is fundamentally linked to long-term evolution of the Gondwanan upper mantle associated with an Archaean–Proterozoic lithospheric terrane boundary. Generation of separate plumbing systems and melt production and contamination in compositionally different parts of such a heterogeneous mantle source may well have been controlled by the distribution and nature of lithospheric thinning zones, which related to initial rifting and the break-up of Gondwana in the western Dronning Maud Land.

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