

Modular S_3 -invariant flavor model in SU(5) grand unified theory

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We present a flavor model with S_3 modular invariance in the framework of SU(5) grand unified theory (GUT). The S_3 modular forms of weights 2 and 4 give the quark and lepton mass matrices with a common complex parameter, the modulus τ . The GUT relation of down-type quarks and charged leptons is imposed by the vacuum expectation value (VEV) of the adjoint 24-dimensional Higgs multiplet in addition to the VEVs of 5 and $\bar{5}$ Higgs multiplets of SU(5). The observed Cabibbo–Kobayashi–Maskawa and Pontecorvo–Maki–Nakagawa–Sakata mixing parameters as well as the mass eigenvalues are reproduced properly. We discuss the leptonic charge–parity phase and the effective mass of the neutrinoless double beta decay with the sum of neutrino masses.
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1. Introduction

The standard model (SM) was well established by the discovery of the Higgs boson. The SM, however, does not answer a fundamental question about the origin of flavor structure. In order to understand this, many works have addressed the discrete groups for flavors. The S_3 group was used in early models of quark masses and mixing angles [1,2]. This group was also studied to explain the large mixing angle [3] in the oscillation of atmospheric neutrinos [4]. After the discovery of the neutrino oscillations, the discrete symmetries of flavors have been developed to reproduce the observed lepton mixing angles [5–13].

Superstring theory with certain compactifications can lead to non-Abelian discrete flavor symmetries. (See, e.g., Refs. [14–20].) The torus and orbifold compactifications have the modular symmetry of the modulus parameter. The flavors of both quarks and leptons transform non-trivially under the modular transformation [21–27]. In this sense, the modular symmetry is a non-Abelian discrete flavor symmetry. Yukawa and other couplings depend on the moduli parameters in four-dimensional low-energy effective field theory derived from superstring theory. Each coupling therefore transforms non-trivially under the modular symmetry, which is an important difference from the conventional flavor symmetries.

The modular group includes S_3 , A_4 , S_4 , and A_5 as its finite subgroups [28]. An attractive flavor model has been put forward based on the $\Gamma_3 \simeq A_4$ modular group [29]. This work stimulates model building based on $\Gamma_2 \simeq S_3$ [30], $\Gamma_4 \simeq S_4$ [31], and $\Gamma_5 \simeq A_5$ [32]. Phenomenological discussions of neutrino

flavor mixing have been presented based on the A_4 [33,34], S_4 [35], and A_5 [36] modular groups. In particular, comprehensive analysis of the A_4 modular group has provided a distinct prediction of the neutrino mixing angles and the charge–parity (CP) violating phase [34]. Applications of the modular symmetry have begun to develop in quark and lepton flavors. The A_4 modular symmetry has also been applied to the SU(5) grand unified theory (GUT) of quarks and leptons [37], while the residual symmetry of the A_4 modular symmetry has been investigated phenomenologically [38]. The modular forms for $\Delta(96)$ and $\Delta(384)$ have also been constructed [39], and the extension of the traditional flavor group has been discussed with modular symmetries [40]. Moreover, multiple modular symmetries are proposed as the origin of flavor [41]. The modular invariance has also been studied combined with generalized CP symmetries for theories of flavors [42]. The quark mass matrix has been discussed in the S_3 and A_4 modular symmetries as well [43,44]. Besides the mass matrices of quarks and leptons, related topics such as baryon number violation [43], dark matter [45], radiatively induced neutrino masses [46], and the modular symmetry anomaly [47] have been discussed.

Among these, the unification of quark and lepton flavors based on the modular symmetry is an important work from the standpoint of quark–lepton unification [37,48] since the modulus τ is common to both quarks and leptons. In this paper we construct an S_3 flavor model with modular invariance in the framework of SU(5) GUT and discuss the Dirac CP violating phases in both quark and lepton sectors as well as the neutrino masses and mixing, the effective neutrino mass of the neutrinoless double beta decay, and Majorana CP violating phases. We consider a six-dimensional compact space X^6 in addition to our four-dimensional spacetime in superstring theory. Suppose that the six-dimensional compact space has some constituent spaces and that they include a two-dimensional compact space X^2 . Note that X^2 can have geometrical symmetry such as the modular symmetry. Quark mixing and lepton mixing are explained by a single flavor symmetry originating from X^2 . The modular forms for the quark and lepton sectors are the same and determined by a common value of τ in our setup. The other four-dimensional part of X^6 may contribute to an overall factor of the Yukawa couplings, but not to their ratios.

We assume the S_3 modular symmetry for flavors of quarks and leptons since it is the minimal non-Abelian discrete symmetry. Furthermore, we assume SU(5) GUT as a first step to building a realistic flavor model with modular invariance for both quarks and leptons. It is emphasized that the vacuum expectation value (VEV) of the 24-dimensional adjoint Higgs multiplet H_{24} creates a difference between the mass eigenvalues of down-type quarks and charged leptons. Our mass matrices reproduce the observed Cabibbo–Kobayashi–Maskawa (CKM) and Pontecorvo–Maki–Nakagawa–Sakata (PMNS) parameters successfully. We predict the leptonic CP violation phase and the effective mass of the neutrinoless double beta decay versus the sum of neutrino masses, respectively.

This paper is organized as follows. In Sect. 2 we present our SU(5) GUT model with the finite modular symmetry $\Gamma_2 \simeq S_3$. In Sect. 3, we present numerical analyses of our model. Section 4 is devoted to a summary. Appendix A shows the modular forms of S_3 briefly, and Appendix B presents relevant parameters in the lepton flavor mixing.

2. Quark and lepton mass matrices in SU(5) GUT

Let us present our framework in the supersymmetric (SUSY) SU(5) GUT. Matter fields can be accommodated in the $\bar{F} = 5$ and $T = 10$ representations as

Table 1. The charge assignments of SU(5), S_3 , and weight for superfields and modular forms. The subscript i of F_i and T_i denotes the i th family.

	$T_{1,2}$	T_3	$F_{1,2}$	F_3	$N_{1,2}^c$	N_3^c	H_5	$H_{\bar{5}}$	H_{24}	$Y_2^{(2)}$	$Y_1^{(4)}, Y_2^{(4)}$
SU(5)	10	10	$\bar{5}$	$\bar{5}$	1	1	5	$\bar{5}$	24	1	1
S_3	2	1'	2	1'	1	1'	1	1	1	2	1, 2
Weight	-2	0	-2	0	0	0	0	0	0	2	4

$$F(\bar{\mathbf{5}}) = \begin{pmatrix} d_1^c \\ d_2^c \\ d_3^c \\ e \\ -\nu \end{pmatrix}_L, \quad T(\mathbf{10}) = \begin{pmatrix} 0 & u_3^c & -u_2^c & u_1 & d_1 \\ -u_3^c & 0 & u_1^c & u_2 & d_2 \\ u_2^c & -u_1^c & 0 & u_3 & d_3 \\ -u_1 & -u_2 & -u_3 & 0 & e^c \\ -d_1 & -d_2 & -d_3 & -e^c & 0 \end{pmatrix}_L, \quad (1)$$

where the subscripts 1, 2, 3 denote the quark colors, the superscript c denotes CP-conjugated fermions, and the flavor indices are omitted. In addition, we introduce the right-handed neutrinos N_i^c ($i = 1, 2, 3$), which are SU(5) singlets. We present the charge assignments of superfields for the SU(5) gauge group, S_3 flavor symmetry, and modular weights in Table 1, where the subscript i of F_i and T_i denotes the i th family. An adjoint representation of scalars H_{24} breaks the SU(5) gauge symmetry and leads to the mass differences among quarks and charged leptons. The electroweak breaking of the SM is realized by a 5 ($\bar{5}$) of Higgs, H_5 ($H_{\bar{5}}$), which also contribute to the fermion mass matrices. These Higgs multiplets are listed in Table 1, which also presents the modular forms of weights 2 and 4 that we use.

For Yukawa interactions, the S_3 modular invariant superpotential is written as

$$w = w_{10} + w_{10, \bar{5}} + w_{\nu}, \quad (2)$$

where the three terms of the right-hand side lead to the mass terms of up-type quarks, down-type quarks, and charged leptons and neutrinos, respectively. The up-type quark mass matrix is derived from w_{10} , which is explicitly given as:

$$w_{10} = (\alpha'_1 Y_1^{(4)} + \alpha'_2 Y_2^{(4)}) T_{1,2} T_{1,2} H_5 \left(1 + k'_1 \frac{H_{24}}{\Lambda} \right) + \beta' Y_2^{(2)} T_{1,2} T_3 H_5 \left(1 + k'_2 \frac{H_{24}}{\Lambda} \right) + \gamma' T_3 T_3 H_5 \left(1 + k'_3 \frac{H_{24}}{\Lambda} \right), \quad (3)$$

where $\alpha'_{1,2}$, β' , $k'_{1,2,3}$, and γ' are dimensionless complex constants. Here, Λ denotes the cut-off scale around the SU(5) energy scale. We set $\langle H_{24} \rangle / \Lambda = 0.3$. Thus, the next-order corrections of $\langle H_{24} \rangle^2 / \Lambda^2$ are $\mathcal{O}(0.1)$. We neglect their effect because the experimental values of masses and mixing angles for the quarks and leptons include errors of $\mathcal{O}(10\%)$. We focus on the parameter regions $|k'_i| = [0, 1.5]$ in the following numerical analysis. By using the S_3 tensor product of doublets in Appendix A, the mass matrix of up-type quarks is given in terms of the modular forms $Y_1(\tau)$ and $Y_2(\tau)$ of Appendix A as

$$M_u = \begin{pmatrix} \varepsilon^u & 2c''^u Y_1 Y_2 & c''_{13} Y_2 \\ 2c''^u Y_1 Y_2 & \varepsilon^u - 2c''^u (Y_1^2 - Y_2^2) & -c''_{13} Y_1 \\ c''_{13} Y_2 & -c''_{13} Y_1 & c''_{33} \end{pmatrix}, \quad (4)$$

where the argument τ of the modular forms is omitted, and the parameters are redefined as follows:

$$\begin{aligned} \varepsilon^u &\equiv v_u[\alpha'_1(Y_1^2 + Y_2^2) + \alpha'_2(Y_1^2 - Y_2^2)](1 + k'_1\langle H_{24} \rangle/\Lambda), \\ c^u &\equiv v_u\alpha'_2(1 + k'_1\langle H_{24} \rangle/\Lambda), \quad c_{13}^u \equiv v_u\beta'(1 + k'_2\langle H_{24} \rangle/\Lambda), \quad c_{33}^u \equiv v_u\gamma'(1 + k'_3\langle H_{24} \rangle/\Lambda), \end{aligned} \tag{5}$$

where v_u is the VEV for the doublet component H_u of H_5 . This mass matrix was investigated in our previous work [43].

Suppose the neutrinos to be Majorana particles, which are realized by the seesaw mechanism. Then, the neutrino mass matrix is derived from the superpotential w_ν :

$$\begin{aligned} w_\nu &= \tilde{m}_3 N_3^c N_3^c + \sum_{i=1}^2 \tilde{m}_i N_i^c N_i^c + b_3^v N_3^c F_3 H_5 \\ &\quad + a_3^v (Y_1 F_2 - Y_2 F_1) H_5 N_3^c + \sum_{i=1}^2 a_i^v (Y_1 F_1 + Y_2 F_2) H_5 N_i^c + \Delta w_\nu, \end{aligned} \tag{6}$$

where a_i^v ($i = 1-3$), b_3^v are dimensionless complex constants. The additional term Δw_ν is the contribution to the right-handed Majorana mass terms from the dimension-five operators,

$$\Delta w_\nu = f_3 \frac{1}{\Lambda} H_{24} H_{24} N_3^c N_3^c + \sum_{i=1}^2 f_i \frac{1}{\Lambda} H_{24} H_{24} N_i^c N_i^c, \tag{7}$$

where the f_i are arbitrary coefficients. Here, we take the diagonal basis of $N_i^c N_i^c$.

After integrating out N_i^c ($i = 1-3$) fields, the Majorana left-handed neutrino mass matrix is therefore given as follows:

$$M_\nu = \frac{v_u^2}{m_{N3}} \begin{pmatrix} a_0 & 2a_2 Y_1 Y_2 & bY_2 \\ 2a_2 Y_1 Y_2 & a_0 - 2a_2(Y_1^2 - Y_2^2) & -bY_1 \\ bY_2 & -bY_1 & c \end{pmatrix}_{LL}, \tag{8}$$

where the parameters a_0, a_1, a_2, b , and c are redefined as

$$\begin{aligned} a_1 &\equiv \frac{1}{8} \left(\frac{m_{N3}}{m_{N1}} (a_1^v)^2 + \frac{m_{N3}}{m_{N2}} (a_2^v)^2 + (a_3^v)^2 \right), \quad a_2 \equiv \frac{1}{8} \left(\frac{m_{N3}}{m_{N1}} (a_1^v)^2 + \frac{m_{N3}}{m_{N2}} (a_2^v)^2 - (a_3^v)^2 \right), \\ b &\equiv -\frac{1}{4} a_3^v b_3^v, \quad c \equiv \frac{1}{4} (b_3^v)^2, \quad a_0 \equiv a_1(Y_1^2 + Y_2^2) + a_2(Y_1^2 - Y_2^2), \end{aligned} \tag{9}$$

and

$$m_{Ni} \equiv \tilde{m}_i + f_i \frac{1}{\Lambda} \langle H_{24} \rangle^2 \quad (i = 1, 2, 3). \tag{10}$$

The superpotentials for the down-type quarks and charged leptons are written as

$$\begin{aligned} w_{10, \bar{5}} &= (\alpha_1 Y_1^{(4)} + \alpha_2 Y_2^{(4)}) T_{1,2} F_{1,2} H_{\bar{5}} \left(1 + k_1 \frac{H_{24}}{\Lambda} \right) + \beta_1 Y_2^{(2)} T_{1,2} F_3 H_{\bar{5}} \left(1 + k_2 \frac{H_{24}}{\Lambda} \right) \\ &\quad + \beta_2 Y_2^{(2)} T_3 F_{1,2} H_{\bar{5}} \left(1 + k_3 \frac{H_{24}}{\Lambda} \right) + \gamma T_3 F_3 H_{\bar{5}} \left(1 + k_4 \frac{H_{24}}{\Lambda} \right), \end{aligned} \tag{11}$$

where $\alpha_{1,2}$, $\beta_{1,2}$, $k_{1,2,3,4}$, and γ are dimensionless complex constants. We focus on the parameter region $|k_i| = [0, 1.5]$ in the following numerical analysis. We can construct a mass matrix for the down-type quarks and charged leptons:

$$M_{10,\bar{5}} = v_d \begin{pmatrix} \alpha_0(1 + k_1 \langle H_{24} \rangle / \Lambda) & 2\alpha_2(1 + k_1 \langle H_{24} \rangle / \Lambda) Y_1 Y_2 & \beta_1(1 + k_2 \langle H_{24} \rangle / \Lambda) Y_2 \\ 2\alpha_2(1 + k_1 \langle H_{24} \rangle / \Lambda) Y_1 Y_2 & (1 + k_1 \langle H_{24} \rangle / \Lambda) [\alpha_0 - 2\alpha_2(Y_1^2 - Y_2^2)] & -\beta_1(1 + k_2 \langle H_{24} \rangle / \Lambda) Y_1 \\ \beta_2(1 + k_3 \langle H_{24} \rangle / \Lambda) Y_2 & -\beta_2(1 + k_3 \langle H_{24} \rangle / \Lambda) Y_1 & \gamma(1 + k_4 \langle H_{24} \rangle / \Lambda) Y_1 \end{pmatrix}_{RL}, \quad (12)$$

where we have introduced a new parameter α_0 defined as

$$\alpha_0 \equiv \alpha_1(Y_1^2 + Y_2^2) + \alpha_2(Y_1^2 - Y_2^2), \quad (13)$$

and v_d is the VEV of the doublet component of $H_{\bar{5}}$. We can obtain a successful mass matrix for the down-type quarks:

$$M_d = \begin{pmatrix} \varepsilon^d & 2c'^d Y_1 Y_2 & c_{13}^d Y_2 \\ 2c'^d Y_1 Y_2 & \varepsilon - 2c'^d(Y_1^2 - Y_2^2) & -c_{13}^d Y_1 \\ c_{31}^d Y_2 & -c_{31}^d Y_1 & c_{33}^d \end{pmatrix}, \quad (14)$$

where we have redefined some parameters as in Eq. (5).

The quark mass matrices in Eqs. (4) and (14) can reproduce the observed CKM mixing matrix elements and quark mass ratios at the GUT scale [49,50]. Indeed, we have obtained successful up-type and down-type quark mass matrices with hierarchical flavor structure which are completely consistent with the observed masses and CKM parameters [43].

Let us discuss the charged lepton mass matrix, which is possibly related to the down-type quark mass matrix, by using the SU(5) GUT relation. We rewrite the coefficients in the down-type quark mass matrix elements in terms of the sum of contributions from VEVs of $H_{\bar{5}}$ and H_{24} as follows:

$$\varepsilon^d = \varepsilon^5 + \varepsilon^{24}, \quad c'^d = c'^5 + c'^{24}, \quad c_{13}^d = c_{13}^5 + c_{13}^{24}, \quad c_{31}^d = c_{31}^5 + c_{31}^{24}, \quad c_{33}^d = c_{33}^5 + c_{33}^{24}, \quad (15)$$

where we have the following relations for the parameters of Eq. (12):

$$\alpha_0 = \varepsilon^5 / v_d, \quad \alpha_2 = c'^5 / v_d, \quad \beta_1 = c_{31}^5 / v_d, \quad \beta_2 = c_{13}^5 / v_d, \quad \gamma = c_{33}^5 / v_d. \quad (16)$$

Let us give the Clebsch–Gordan (CG) factor C , which is derived by the ratio of VEVs for the charged lepton sector and down-type quark sector:

$$C \equiv \frac{\langle H_{24}^l \rangle}{\langle H_{24}^q \rangle} = -3/2, \quad (17)$$

since H_{24} takes the VEV as $\langle H_{24} \rangle \propto \text{diag}[2, 2, 2, -3, -3]$. The charged lepton mass matrix is therefore obtained in terms of the elements of the down-type quark mass matrix and the coefficient C by transposing the down-type quark mass matrix:

$$M_e = \begin{pmatrix} \varepsilon^5 + C\varepsilon^{24} & 2(c'^5 + Cc'^{24}) Y_1 Y_2 & (c_{31}^5 + Cc_{31}^{24}) Y_2 \\ 2(c'^5 + Cc'^{24}) Y_1 Y_2 & (\varepsilon^5 + C\varepsilon^{24}) - 2(c'^5 + Cc'^{24})(Y_1^2 - Y_2^2) & -(c_{31}^5 + Cc_{31}^{24}) Y_1 \\ (c_{13}^5 + Cc_{13}^{24}) Y_2 & -(c_{13}^5 + Cc_{13}^{24}) Y_1 & c_{33}^5 + Cc_{33}^{24} \end{pmatrix}. \quad (18)$$

In the quark and lepton sectors, we obtained enough parameter sets including the value of the modulus τ which reproduce the quark and lepton masses and CKM mixing. For example, we set

$$\text{Re}[\tau] = 0.465, \quad \text{Im}[\tau] = 1.31, \quad (19)$$

which lead to $Y_1 = 0.116 \exp[4.98 \times 10^{-4} \pi i]$ and $Y_2 = 0.0267 \exp[0.461 \pi i]$. We show a typical sample of our parameter sets:

$$\begin{aligned} \varepsilon^u &= 7.81 \times 10^{-6} e^{-0.508 \pi i}, & \varepsilon^5 &= 6.42 \times 10^{-4} e^{-0.788 \pi i}, & \varepsilon^{24} &= 2.19 \times 10^{-4} e^{-0.874 \pi i}, \\ c'^u &= 2.04 \times 10^{-4} e^{-0.807 \pi i}, & c'^5 &= 2.42 \times 10^{-2} e^{-0.174 \pi i}, & c'^{24} &= 8.29 \times 10^{-3} e^{-0.261 \pi i}, \\ c_{13}^u &= 0.443 e^{0.802 \pi i}, & c_{13}^5 &= 2.12 e^{0.184 \pi i}, & c_{13}^{24} &= 0.619 e^{-0.876 \pi i}, \\ & & c_{31}^5 &= 1.03 e^{-0.505 \pi i}, & c_{31}^{24} &= 0.428 e^{0.561 \pi i}, \\ & & c_{33}^5 &= 0.995 e^{-0.0524 \pi i}, & c_{33}^{24} &= 0.164 e^{0.465 \pi i} \end{aligned} \quad (20)$$

in $c_{33}^u = c_{33}^d = 1$ GeV units. These are obtained from the parameters

$$\begin{aligned} \alpha'_1 &= 1.05 e^{0.201 \pi i}, & \alpha'_2 &= 1.92 \times 10^{-4} e^{-0.797 \pi i}, \\ \beta' &= 0.417 e^{0.801 \pi i}, & \gamma' &= 0.935 e^{0.00331 \pi i}, \\ k'_1 &= 0.238 e^{-0.166 \pi i}, & k'_2 &= 0.210 e^{0.0156 \pi i}, & k'_3 &= 0.236 e^{-0.0502 \pi i} \end{aligned} \quad (21)$$

in w_{10} of Eq. (3), and the parameters

$$\begin{aligned} \alpha_0 &= 6.42 \times 10^{-4} e^{-0.788 \pi i}, & \alpha_2 &= 0.0242 e^{-0.174 \pi i}, & \beta_1 &= 1.03 e^{-0.505 \pi i}, \\ \beta_2 &= 2.12 e^{0.184 \pi i}, & \gamma &= 0.995 e^{-0.0524 \pi i}, & k_1 &= 0.342 e^{-0.0864 \pi i}, \\ k_2 &= 0.292 e^{0.940 \pi i}, & k_3 &= 0.416 e^{-0.935 \pi i}, & k_4 &= 0.165 e^{0.517 \pi i} \end{aligned} \quad (22)$$

in $w_{10, \bar{5}}$ of Eq. (11).

This sample parameter set leads to the following result for the CKM mixing parameters:

$$|V_{\text{CKM}}| = \begin{pmatrix} 0.9746 & 0.2243 & 0.0025 \\ 0.2238 & 0.9745 & 0.0180 \\ 0.0040 & 0.0177 & 0.9998 \end{pmatrix}, \quad \delta_{\text{CP}}^{\text{CKM}} = 71.18[^\circ], \quad (23)$$

as well as the proper hierarchy of quark and charged lepton masses. We use the above parameters for the prediction of the neutrino sector in the next section.

3. Numerical results

We have obtained parameter regions that reproduce the observed fermion mass ratios and CKM mixing parameters. Our results are consistent with the experimental results for quark mass ratios and charged lepton mass ratios at the GUT scale within the 1σ range [49,50].¹ In the following

¹ The quark masses are obtained at the GUT scale of 2×10^{16} GeV by putting $v_u/v_d = 10$ in the minimal supersymmetric standard model, where the SUSY breaking scale is taken to be 1–10 TeV. In the region of $\tan \beta = 3$ –10, our numerical values are changed only by a few percent. Proton decay may favor the larger SUSY breaking scale such as 10 TeV as discussed in Sect. 3.3.

subsections we present predictions in the neutrino sector and discuss the correlation between the CKM and PMNS mixing parameters.

Since we have separated the parameters of the down-type quarks such as ε , c' , c_{13} , c_{31} , and c_{33} into two terms as defined in Eq. (15), we can scan the parameters in the charged lepton mass matrix of Eq. (18) by using the successful parameter sets of the down-type quark sector. A typical sample is presented in Eq. (20). The parameters of the neutrino mass matrix of Eq. (8) have been scanned in the region of $0 < |a_0| < 2$, $0 < |a_2| < 50$, and $0 < |b| < 15$ in $c = 1$ units, while phases have been scanned in $[-\pi, \pi]$. We present a sample point satisfying the recent neutrino oscillation experimental data [51,52] as well as the fermion mass ratio and CKM mixing parameters at the GUT scale.²

3.1. Neutrino phenomenology

In our numerical study, we have set $\langle H_{24} \rangle / \Lambda = 0.3$ with $\langle H_{24} \rangle \simeq 2 \times 10^{16}$ GeV. Then, we have $\langle H_{24} \rangle^2 / \Lambda \simeq 6 \times 10^{15}$ GeV, which is related to the right-handed neutrino mass in Eq. (10). Therefore, we take $m_{N_i} \simeq 10^{15}$ GeV by choosing relevant values for \tilde{m}_i and f_i .³ Then, the neutrino Yukawa couplings are found to be at most 1.3 by inputting the experimental data. Thus, our setup is reasonably accepted in the neutrino phenomenology if $m_{N_i} \simeq 10^{15}$ GeV is taken. We also discuss proton decay in this setup later.

Our lepton mass matrices of Eqs. (8) and (18) reproduce the experimental result of neutrino mass squared differences and the three mixing angles within 3σ range [51,52]. The following results are constrained by the cosmological bound of the sum of three light neutrino masses m_i , which is $\Sigma m_i < 0.12$ eV [55,56]. First, we show two sample parameter sets leading to successful results, which are completely consistent with the observed CKM and PMNS matrices. We obtain a prediction for the normal hierarchy (NH) of neutrino masses from the parameter set of Eq. (20) and the following parameters:⁴

$$\frac{a_0}{c} = 1.61 e^{0.525\pi i}, \quad \frac{a_2}{c} = 178 e^{0.502\pi i}, \quad \frac{b}{c} = 18.0 e^{-0.997\pi i}, \quad (24)$$

in which cv_u^2/m_{N3} is a typical neutrino mass scale. If we take the right-handed neutrino mass m_{N3} to be smaller than 10^{14} GeV, c is less than 0.1. We obtain the three lepton mixing angles θ_{12} , θ_{23} , and θ_{13} , the Dirac CP violating phase δ_{CP} , neutrino masses, the effective mass of the neutrinoless double beta decay $\langle m_{ee} \rangle$, and the Majorana phases α_{21} and α_{31} (see the notations in Appendix B) as follows:

$$\begin{aligned} \sin^2 \theta_{12} &= 0.287, & \sin^2 \theta_{23} &= 0.604, & \sin^2 \theta_{13} &= 0.0208, & \delta_{\text{CP}} &= -89.6[^\circ], \\ \Delta m_{21}^2 &= 7.14 \times 10^{-5} [\text{eV}^2], & \Delta m_{31}^2 &= 2.60 \times 10^{-3} [\text{eV}^2], & m_1 &= 11.7 [\text{meV}], \\ \sum_i m_i &= 117 [\text{meV}], & \langle m_{ee} \rangle &= 13.1 [\text{meV}], & \alpha_{21} &= -23.3[^\circ], & \alpha_{31} &= 168[^\circ]. \end{aligned} \quad (25)$$

² We have neglected the renormalization corrections for the neutrino masses and mixing parameters although the numerical analysis should be presented at GUT scale. A numerical estimation of the quantum corrections in Ref. [53] showed that the corrections are negligible as long as the neutrino mass scale is smaller than 200 meV and $\tan \beta \leq 10$. See also Refs. [33,54].

³ We may consider $f_i \simeq 1/6$ or the cancellation due to phases of \tilde{m}_i and f_i .

⁴ Note that a_2/c and b/c are larger than $\mathcal{O}(1)$, but they are parameters and the couplings $a_2 Y_2^{(4)}$ and $b Y^{(2)}$ themselves are smaller than 1.

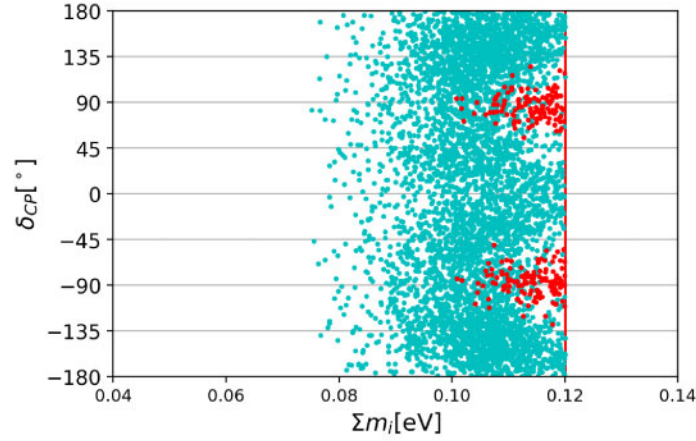


Fig. 1. The prediction of the neutrino mass sum and δ_{CP} , where the cyan and red points denote the NH and IH cases, respectively. The red line represents the cosmological bound.

For the inverted hierarchy (IH) of neutrino masses, we use the same parameter values as the NH except

$$\frac{a_0}{c} = 2.57 e^{0.536\pi i}, \quad \frac{a_2}{c} = 13.6 e^{0.620\pi i}, \quad \frac{b}{c} = 6.79 e^{0.313\pi i}. \quad (26)$$

Then, we obtain:

$$\begin{aligned} \sin^2 \theta_{12} &= 0.314, & \sin^2 \theta_{23} &= 0.521, & \sin^2 \theta_{13} &= 0.0244, & \delta_{CP} &= -90.4[^\circ], \\ \Delta m_{21}^2 &= 7.64 \times 10^{-5} [\text{eV}^2], & \Delta m_{31}^2 &= -2.52 \times 10^{-3} [\text{eV}^2], & m_3 &= 11.4 [\text{meV}], \\ \sum_i m_i &= 115 [\text{meV}], & \langle m_{ee} \rangle &= 47.2 [\text{meV}], & \alpha_{21} &= 43.3[^\circ], & \alpha_{31} &= 46.4[^\circ]. \end{aligned} \quad (27)$$

Let us discuss our prediction of the leptonic CP phase, the effective mass of the neutrinoless double beta decay with the sum of neutrino masses. We show the allowed region in the plane of the sum of neutrino masses Σm_i and δ_{CP} in Fig. 1, where the cyan points and red points denote the predictions for the NH and IH cases, respectively. For NH, the predicted Dirac CP violating phase is allowed in the whole range of $[-180^\circ, 180^\circ]$ while Σm_i is larger than 75 meV. In the case of IH, δ_{CP} is predicted in the region of $\pm[50^\circ, 130^\circ]$. In particular, it is around $\pm 90^\circ$ near the lower bound of our prediction of the sum of neutrino masses, 100 meV. It is noted that the lightest neutrino mass is larger than $m_3 = 1.61$ meV for the IH case. The future development of neutrino oscillation experiments or cosmological analysis for the neutrino mass is therefore expected to test our model.

We also show a prediction of $\langle m_{ee} \rangle$ in the neutrinoless double beta decay in Fig. 2. The predicted region of the effective mass is about $10 < \langle m_{ee} \rangle < 30$ meV for NH and $47 < \langle m_{ee} \rangle < 50$ meV for IH. If the neutrinos are Majorana particles, the experiments for the neutrinoless double beta decay may test this model in the future.

3.2. Common modulus τ in quarks and leptons

The modulus τ is a key parameter in unifying quark and lepton flavors. We show the allowed region of the modulus τ in Fig. 3 which leads to successful quark masses and CKM mixing parameters at the GUT scale within the 1σ range. Both real and imaginary parts of τ are rather broad as $\text{Re}[\tau] = 0.2\text{--}0.9$ and $\text{Im}[\tau] = 1.1\text{--}1.5$.

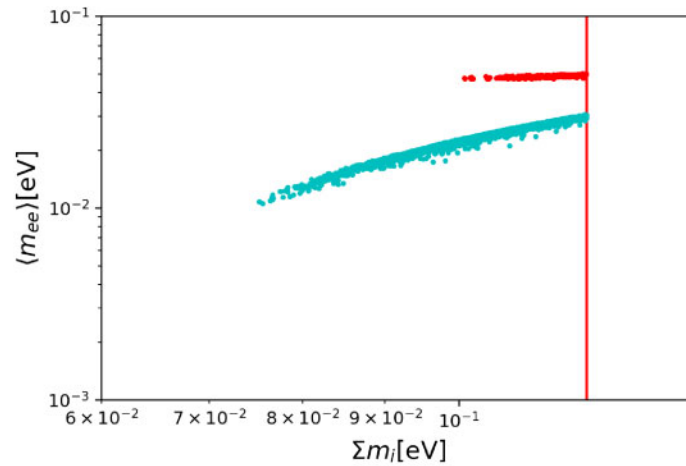


Fig. 2. The prediction of the effective mass for $0\nu\beta\beta$ decay, where the cyan and red points denote the NH and IH cases, respectively. The red line represents the cosmological bound.

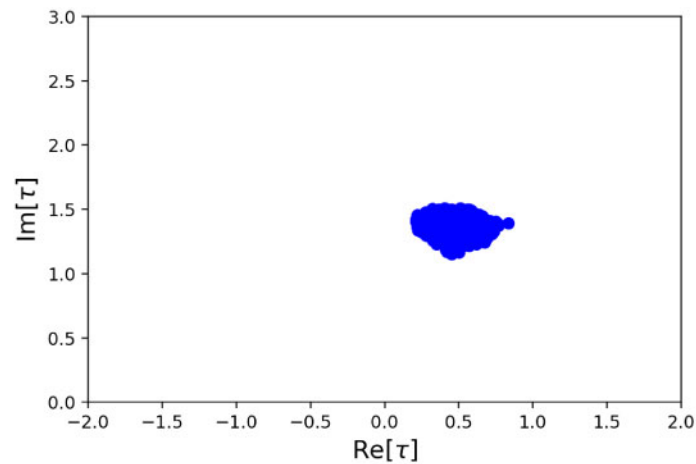


Fig. 3. Allowed region of τ constrained only from the quark sector.

The quark and lepton mass matrices should have the common modulus τ . Inputting τ obtained in the quark sector as well as other parameters of the quarks, we have obtained the allowed region of τ which satisfies the observed 1σ range of the charged lepton mass ratios at the GUT scale and 3σ range of the PMNS parameters. It is noted that there are no clear correlations between CKM and PMNS mixing parameters because of the large number of free parameters embedded in our model.

For the NH case, we plot the allowed region of τ in Fig. 4, where τ for the quark sector is also shown. Both regions almost overlap. However, for the IH case, the allowed region of τ is different from that in the case of quarks only, as seen in Fig. 5. Note that the allowed region is clearly reduced compared with the one for quarks only.

Thus, we obtain the restricted common τ in spite of the many free parameters of our model.

3.3. Proton decay

We give a brief comment on proton decay. An SU(5) GUT model includes the color-triplet Higgs multiplets, which can lead to proton decay [58–62]. The color-triplet Higgs multiplets induce the

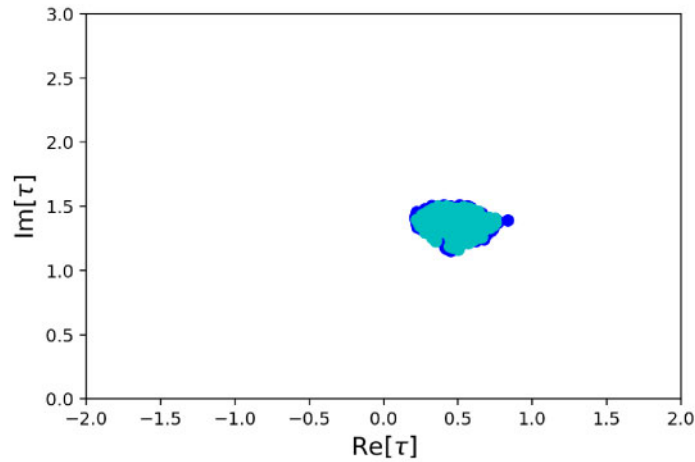


Fig. 4. Allowed region of τ in both quarks and leptons (cyan points) for NH. The region in quarks only is denoted by blue points.

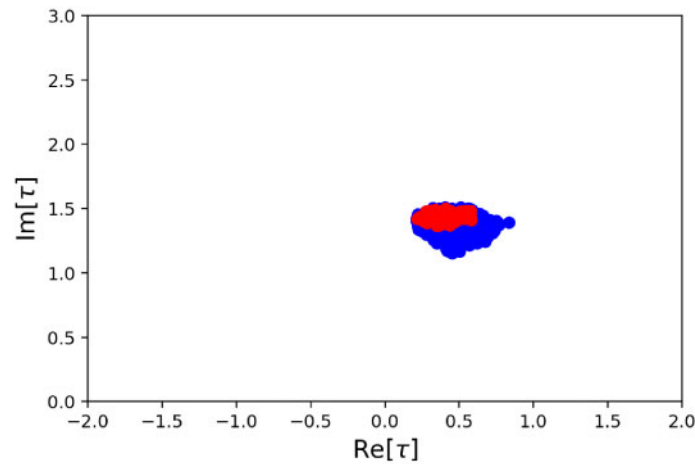


Fig. 5. Allowed region of τ in both quarks and leptons (red points) for IH. The region in quarks only is denoted by blue points.

effective superpotential

$$w_S = \frac{1}{M_{H_c}} f_{u_i} e^{i\phi_i} f_{d_\ell} V_{k\ell}^* \varepsilon_{\alpha\beta\gamma} u_{i\alpha}^c e_i^c u_{k\beta}^c d_{\ell\gamma}^c \quad (28)$$

as well as $QQQL/M_{H_c}$ including the quark doublet superfields Q , where M_{H_c} is the mass of the color-triplet Higgs multiplets, f_{u_i} and f_{d_ℓ} are Yukawa couplings of up-sector and down-sector quarks, and ϕ_i are their phases.⁵ The above operator leads to the proton decay $p \rightarrow K^+ \nu$ through higgsino exchange. The factors $f_{u_3} f_{d_\ell} V_{1\ell}^*$ with $\ell = 1, 2$ are crucial to estimating the proton lifetime, because the couplings among u_R , d_R (s_R), right-handed stop, and right-handed stau are important in this

⁵ We follow the notation in Refs. [58,62].

process. Then, the proton lifetime is given as [62]

$$\tau_P \simeq 4 \times 10^{35} \times \sin^4 2\beta \left(\frac{0.1}{\hat{A}_R} \right)^2 \left(\frac{M_S}{100 \text{ TeV}} \right)^2 \left(\frac{M_{H_c}}{10^{16} \text{ GeV}} \right)^2 \text{ yrs}, \quad (29)$$

where \hat{A}_R is the renormalization factor and M_S is the sfermion mass scale. The proton lifetime is longer than the observed lower bound of 10^{33} yrs [57] for the case of $M_{H_c} \simeq 2 \times 10^{16}$ GeV if $M_S \geq 10$ TeV and $\tan \beta \leq 3$. Since our numerical results for the quark/lepton mass matrices are changed only within a few percent in the range of $\tan \beta = 3$ –10, as stated in footnote 1, $M_S = 10$ TeV is the minimal one that is consistent with our numerical results for quark/lepton flavor mixing to protect the proton decay.⁶

We may have additional contributions to the effective potential in Eq. (28), because our cut-off scale Λ is slightly higher than the GUT scale, $\langle H_{24} \rangle / \Lambda = 0.3$. For example, the following term is allowed by the symmetries:

$$w'_5 = \frac{f}{\Lambda} T_3 T_3 T_3 F_3, \quad (30)$$

where f is a modulus-independent coupling constant. The field T_3 includes u_R by the factor $c_{31}^u Y_2 \sim 5 \times 10^{-3}$, while the field F_3 includes d_R and s_R by the factors $c_{31}^d Y_2 \sim 1 \times 10^{-2}$ and $c_{31}^d Y_1 \sim 1 \times 10^{-1}$. Thus, the above operator leads to the couplings among u_R , d_R (s_R), right-handed stop, and right-handed stau with a similar suppression or strong suppression compared with $f_{u_3 d_\ell} V_{1\ell}^* = \mathcal{O}(10^{-4})$ for $\ell = 1, 2$, when $f = \mathcal{O}(1)$. In our model we set $\langle H_{24} \rangle / \Lambda = 0.3$ and $\langle H_{24} \rangle \simeq 2 \times 10^{16}$ GeV. For $M_{H_c} < \Lambda$ and $f \lesssim 1$, the processes including the color-triplet Higgs multiplets of Eq. (28) would be dominant in the proton decay. Similarly, we can discuss the operators including $T_{1,2}$ and $F_{1,2}$, although they should have modulus-dependent couplings.

4. Summary and discussions

We have presented a flavor model with the S_3 modular invariance in the framework of SU(5) GUT and discussed the CKM and PMNS mixing parameters of both quark and lepton sectors. We have considered a six-dimensional compact space X^6 in addition to our four-dimensional space-time and supposed that the six-dimensional compact space has some constituent parts that include a two-dimensional compact space X^2 . Then, the quarks and leptons have the same modular symmetry S_3 and the same value of τ in our setup. We note that our model does not require any gauge singlet scalars such as flavons. The difference between the mass eigenvalues of down-type quarks and charged leptons is realized by the 24-dimensional adjoint Higgs multiplet H_{24} .

The setup of our model is reasonably accepted in the neutrino phenomenology if the right-handed neutrino masses are taken to be around 10^{15} GeV. Their favored ranges are fairly limited; masses larger than 10^{15} GeV are basically disfavored by the perturbativity of the neutrino Yukawa couplings, whereas lower masses require a more significant suppression in the dimension-five operators, or more severe cancellation between the operators and the bare mass terms.

We have analyzed our model numerically and found parameter regions which are consistent with both the observed CKM and PMNS mixing parameters for both NH and IH cases. The predicted Dirac CP violating phase is allowed in the whole range $[-180^\circ, 180^\circ]$ for the NH case. The sum of neutrino

⁶ If $\tan \beta = 10$ is taken, M_S should be larger than 100 TeV.

masses is larger than 75 meV. For IH, it is predicted in the region of $\pm[50^\circ, 130^\circ]$. In particular, it is around $\pm 90^\circ$ near the lower bound of our prediction of the sum of neutrino masses, 100 meV. It is expected to test our model by astronomical observation for the neutrino mass constraint, as well as precise observation for the Dirac CP violating phase.

We have also predicted the effective mass in the neutrinoless double beta decay $\langle m_{ee} \rangle$, which is $10 < \langle m_{ee} \rangle < 30$ [meV] for NH and $47 < \langle m_{ee} \rangle < 50$ [meV] for IH. The development of experiments for the neutrinoless double beta decay is also expected to test our model. It is also noted that the proton lifetime is sufficiently long compared with the observed lower bound of 10^{33} yrs.

Since our model has a large number of free parameters, distinct correlations between the CKM and PMNS mixing parameters are not found. However, the common value of the modulus τ is clearly obtained by imposing the experimental data for the CKM and PMNS mixing parameters as well as the quark and lepton masses. If we can build a flavor model with a small number of free parameters in a specific GUT framework, it is expected to find correlations between the CKM and PMNS matrices.

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Appendix A. Modular forms of S_3 modular group

The Dedekind eta-function $\eta(\tau)$ is defined by

$$\eta(\tau) = q^{1/24} \prod_{n=1}^{\infty} (1 - q^n), \tag{A.1}$$

where $q = e^{2\pi i\tau}$. Using $\eta(\tau)$, the modular forms of weight 2 corresponding to the S_3 doublet are written as [30]

$$Y_1(\tau) = \frac{i}{4\pi} \left(\frac{\eta'(\tau/2)}{\eta(\tau/2)} + \frac{\eta'((\tau+1)/2)}{\eta((\tau+1)/2)} - \frac{8\eta'(2\tau)}{\eta(2\tau)} \right),$$

$$Y_2(\tau) = \frac{\sqrt{3}i}{4\pi} \left(\frac{\eta'(\tau/2)}{\eta(\tau/2)} - \frac{\eta'((\tau+1)/2)}{\eta((\tau+1)/2)} \right),$$

where we use the following basis of S_3 generators S and T in the doublet representation:

$$S = \frac{1}{2} \begin{pmatrix} -1 & -\sqrt{3} \\ -\sqrt{3} & 1 \end{pmatrix}, \quad T = \begin{pmatrix} 1 & 0 \\ 0 & -1 \end{pmatrix}. \tag{A.2}$$

The doublet modular forms have the following q -expansions:

$$Y_2^{(2)} = \begin{pmatrix} Y_1(\tau) \\ Y_2(\tau) \end{pmatrix}_2 = \begin{pmatrix} \frac{1}{8} + 3q + 3q^2 + 12q^3 + 3q^4 + \dots \\ \sqrt{3}q^{1/2}(1 + 4q + 6q^2 + 8q^3 + \dots) \end{pmatrix}_2. \tag{A.3}$$

Since we work in the basis of Eq. (A.2), the tensor product of two doublets is expanded by

$$\begin{pmatrix} x_1 \\ x_2 \end{pmatrix}_2 \otimes \begin{pmatrix} y_1 \\ y_2 \end{pmatrix}_2 = (x_1 y_1 + x_2 y_2)_1 \oplus (x_1 y_2 - x_2 y_1)_{1'} \oplus \begin{pmatrix} x_1 y_1 - x_2 y_2 \\ -x_1 y_2 - x_2 y_1 \end{pmatrix}_2. \quad (\text{A.4})$$

By using the tensor product of the two doublets $(Y_1(\tau), Y_2(\tau))^T$, we can construct modular forms of weight 4, $Y^{(4)}$:

$$\mathbf{1} : Y_1^{(4)} = (Y_1(\tau)^2 + Y_2(\tau)^2)_1, \quad (\text{A.5})$$

$$\mathbf{2} : Y_2^{(4)} = \begin{pmatrix} Y_1(\tau)^2 - Y_2(\tau)^2 \\ -2Y_1(\tau)Y_2(\tau) \end{pmatrix}_2. \quad (\text{A.6})$$

The S_3 singlet $\mathbf{1}'$ modular form of weight 4 vanishes.

Appendix B. Lepton mixing matrix

Supposing neutrinos to be Majorana particles, the PMNS matrix is parametrized in terms of the three mixing angles θ_{ij} ($i, j = 1, 2, 3; i < j$), one CP violating Dirac phase δ_{CP} , and two Majorana phases α_{21}, α_{31} as follows [57]:

$$U_{\text{PMNS}} = \begin{pmatrix} c_{12}c_{13} & s_{12}c_{13} & s_{13}e^{-i\delta_{\text{CP}}} \\ -s_{12}c_{23} - c_{12}s_{23}s_{13}e^{i\delta_{\text{CP}}} & c_{12}c_{23} - s_{12}s_{23}s_{13}e^{i\delta_{\text{CP}}} & s_{23}c_{13} \\ s_{12}s_{23} - c_{12}c_{23}s_{13}e^{i\delta_{\text{CP}}} & -c_{12}s_{23} - s_{12}c_{23}s_{13}e^{i\delta_{\text{CP}}} & c_{23}c_{13} \end{pmatrix} \begin{pmatrix} 1 & 0 & 0 \\ 0 & e^{i\frac{\alpha_{21}}{2}} & 0 \\ 0 & 0 & e^{i\frac{\alpha_{31}}{2}} \end{pmatrix}, \quad (\text{B.1})$$

where c_{ij} and s_{ij} denote $\cos \theta_{ij}$ and $\sin \theta_{ij}$, respectively.

In terms of this parametrization and three neutrino masses, the effective mass in the neutrinoless double beta decay is given as follows:

$$\langle m_{ee} \rangle = |m_1 c_{12}^2 c_{13}^2 + m_2 s_{12}^2 c_{13}^2 e^{i\alpha_{21}} + m_3 s_{13}^2 e^{i(\alpha_{31} - 2\delta_{\text{CP}})}|. \quad (\text{B.2})$$

References

- [1] S. Pakvasa and H. Sugawara, Phys. Lett. B **73**, 61 (1978).
- [2] F. Wilczek and Z. Zee, Phys. Lett. B **70**, 418 (1977); **72**, 503 (1978) [erratum].
- [3] M. Fukugita, M. Tanimoto, and T. Yanagida, Phys. Rev. D **57**, 4429 (1998) [arXiv:hep-ph/9709388] [Search INSPIRE].
- [4] Y. Fukuda et al. [Super-Kamiokande Collaboration], Phys. Rev. Lett. **81**, 1562 (1998) [arXiv:hep-ex/9807003] [Search INSPIRE].
- [5] G. Altarelli and F. Feruglio, Rev. Mod. Phys. **82**, 2701 (2010) [arXiv:1002.0211 [hep-ph]] [Search INSPIRE].
- [6] H. Ishimori, T. Kobayashi, H. Ohki, Y. Shimizu, H. Okada, and M. Tanimoto, Prog. Theor. Phys. Suppl. **183**, 1 (2010) [arXiv:1003.3552 [hep-th]] [Search INSPIRE].
- [7] H. Ishimori, T. Kobayashi, H. Ohki, H. Okada, Y. Shimizu, and M. Tanimoto, Lect. Notes Phys. **858**, 1 (2012).
- [8] D. Hernandez and A. Yu. Smirnov, Phys. Rev. D **86**, 053014 (2012) [arXiv:1204.0445 [hep-ph]] [Search INSPIRE].
- [9] S. F. King and C. Luhn, Rept. Prog. Phys. **76**, 056201 (2013) [arXiv:1301.1340 [hep-ph]] [Search INSPIRE].
- [10] S. F. King, A. Merle, S. Morisi, Y. Shimizu, and M. Tanimoto, New J. Phys. **16**, 045018 (2014) [arXiv:1402.4271 [hep-ph]] [Search INSPIRE].

- [11] M. Tanimoto, AIP Conf. Proc. **1666**, 120002 (2015).
- [12] S. F. King, Prog. Part. Nucl. Phys. **94**, 217 (2017) [arXiv:1701.04413 [hep-ph]] [Search INSPIRE].
- [13] S. T. Petcov, Eur. Phys. J. C **78**, 709 (2018) [arXiv:1711.10806 [hep-ph]] [Search INSPIRE].
- [14] T. Kobayashi, H. P. Nilles, F. Plöger, S. Raby, and M. Ratz, Nucl. Phys. B **768**, 135 (2007) [arXiv:hep-ph/0611020] [Search INSPIRE].
- [15] T. Kobayashi, S. Raby, and R.-J. Zhang, Nucl. Phys. B **704**, 3 (2005) [arXiv:hep-ph/0409098] [Search INSPIRE].
- [16] P. Ko, T. Kobayashi, J.-h. Park, and S. Raby, Phys. Rev. D **76**, 035005 (2007); **76**, 059901 (2007) [erratum] [arXiv:0704.2807 [hep-ph]] [Search INSPIRE].
- [17] F. Beye, T. Kobayashi, and S. Kuwakino, Phys. Lett. B **736**, 433 (2014) [arXiv:1406.4660 [hep-th]] [Search INSPIRE].
- [18] Y. Olguín-Trejo, R. Pérez-Martínez, and S. Ramos-Sánchez, Phys. Rev. D **98**, 106020 (2018) [arXiv:1808.06622 [hep-th]] [Search INSPIRE].
- [19] H. P. Nilles, M. Ratz, A. Trautner, and P. K.S. Vaudrevange, Phys. Lett. B **786**, 283 (2018) [arXiv:1808.07060 [hep-th]] [Search INSPIRE].
- [20] H. Abe, K.-S. Choi, T. Kobayashi, and H. Ohki, Nucl. Phys. B **820**, 317 (2009) [arXiv:0904.2631 [hep-ph]] [Search INSPIRE].
- [21] J. Lauer, J. Mas, and H. P. Nilles, Phys. Lett. B **226**, 251 (1989).
- [22] J. Lauer, J. Mas, and H. P. Nilles, Nucl. Phys. B **351**, 353 (1991).
- [23] W. Lerche, D. Lust, and N. P. Warner, Phys. Lett. B **231**, 417 (1989).
- [24] S. Ferrara, D. Lust, and S. Theisen, Phys. Lett. B **233**, 147 (1989).
- [25] D. Cremades, L. E. Ibáñez, and F. Marchesano, J. High Energy Phys. **0405**, 079 (2004) [arXiv:hep-th/0404229] [Search INSPIRE].
- [26] T. Kobayashi and S. Nagamoto, Phys. Rev. D **96**, 096011 (2017) [arXiv:1709.09784 [hep-th]] [Search INSPIRE].
- [27] T. Kobayashi, S. Nagamoto, S. Takada, S. Tamba, and T. H. Tatsuishi, Phys. Rev. D **97**, 116002 (2018) [arXiv:1804.06644 [hep-th]] [Search INSPIRE].
- [28] R. de Adelhart Toorop, F. Feruglio, and C. Hagedorn, Nucl. Phys. B **858**, 437 (2012) [arXiv:1112.1340 [hep-ph]] [Search INSPIRE].
- [29] F. Feruglio, arXiv:1706.08749 [hep-ph] [Search INSPIRE].
- [30] T. Kobayashi, K. Tanaka, and T. H. Tatsuishi, Phys. Rev. D **98**, 016004 (2018) [arXiv:1803.10391 [hep-ph]] [Search INSPIRE].
- [31] J. T. Penedo and S. T. Petcov, Nucl. Phys. B **939**, 292 (2019) [arXiv:1806.11040 [hep-ph]] [Search INSPIRE].
- [32] P. P. Novichkov, J. T. Penedo, S. T. Petcov and A. V. Titov, J. High Energy Phys. **1904**, 174 (2019) [arXiv:1812.02158 [hep-ph]] [Search INSPIRE].
- [33] J. C. Criado and F. Feruglio, SciPost Phys. **5**, 042 (2018) [arXiv:1807.01125 [hep-ph]] [Search INSPIRE].
- [34] T. Kobayashi, N. Omoto, Y. Shimizu, K. Takagi, M. Tanimoto, and T. H. Tatsuishi, J. High Energy Phys. **1811**, 196 (2018) [arXiv:1808.03012 [hep-ph]] [Search INSPIRE].
- [35] P. P. Novichkov, J. T. Penedo, S. T. Petcov, and A. V. Titov, J. High Energy Phys. **1904**, 005 (2019) [arXiv:1811.04933 [hep-ph]] [Search INSPIRE].
- [36] G.-J. Ding, S. F. King, and X.-G. Liu, Phys. Rev. D **100**, 115005 (2019) [arXiv:1903.12588 [hep-ph]] [Search INSPIRE].
- [37] F.-J. de Anda, S. F. King, and E. Perdomo, Phys. Rev. D **101**, 015028 (2020) [arXiv:1812.05620 [hep-ph]] [Search INSPIRE].
- [38] P. P. Novichkov, S. T. Petcov, and M. Tanimoto, Phys. Lett. B **793**, 247 (2019) [arXiv:1812.11289 [hep-ph]] [Search INSPIRE].
- [39] T. Kobayashi and S. Tamba, Phys. Rev. D **99**, 046001 (2019) [arXiv:1811.11384 [hep-th]] [Search INSPIRE].
- [40] A. Baur, H. P. Nilles, A. Trautner, and P. K. S. Vaudrevange, Phys. Lett. B **795**, 7 (2019) [arXiv:1901.03251 [hep-th]] [Search INSPIRE].
- [41] I. De Medeiros Varzielas, S. F. King, and Y.-L. Zhou, Phys. Rev. D **101**, 055033 (2020) [arXiv:1906.02208 [hep-ph]] [Search INSPIRE].
- [42] P. P. Novichkov, J. T. Penedo, S. T. Petcov, and A. V. Titov, J. High Energy Phys. **1907**, 165 (2019) [arXiv:1905.11970 [hep-ph]] [Search INSPIRE].

- [43] T. Kobayashi, Y. Shimizu, K. Takagi, M. Tanimoto, T. H. Tatsuishi, and H. Uchida, Phys. Lett. B **794**, 114 (2019) [arXiv:1812.11072 [hep-ph]] [Search INSPIRE].
- [44] H. Okada and M. Tanimoto, Phys. Lett. B **791**, 54 (2019) [arXiv:1812.09677 [hep-ph]] [Search INSPIRE].
- [45] T. Nomura and H. Okada, Phys. Lett. B **797**, 134799 (2019) [arXiv:1904.03937 [hep-ph]] [Search INSPIRE].
- [46] T. Nomura and H. Okada, arXiv:1906.03927 [hep-ph] [Search INSPIRE].
- [47] Y. Kariyazono, T. Kobayashi, S. Takada, S. Tamba, and H. Uchida, Phys. Rev. D **100**, 045014 (2019) [arXiv:1904.07546 [hep-th]] [Search INSPIRE].
- [48] H. Okada and M. Tanimoto, arXiv:1905.13421 [hep-ph] [Search INSPIRE].
- [49] S. Antusch and V. Maurer, J. High Energy Phys. **1311**, 115 (2013) [arXiv:1306.6879 [hep-ph]] [Search INSPIRE].
- [50] F. Björkeröth, F. J. de Anda, I. de Medeiros Varzielas, and S. F. King, J. High Energy Phys. **1506**, 141 (2015) [arXiv:1503.03306 [hep-ph]] [Search INSPIRE].
- [51] NuFIT 4.0 (2018) (available at: www.nu-fit.org, date last accessed 6 April, 2020).
- [52] I. Esteban, M. C. Gonzalez-Garcia, M. Maltoni, I. Martinez-Soler, and T. Schwetz, J. High Energy Phys. **1701**, 087 (2017) [arXiv:1611.01514 [hep-ph]] [Search INSPIRE].
- [53] N. Haba and N. Okamura, Eur. Phys. J. C **14**, 347 (2000) [arXiv:hep-ph/9906481] [Search INSPIRE].
- [54] S. Antusch, J. Kersten, M. Lindner, M. Ratz, and M. A. Schmidt, J. High Energy Phys. **0503**, 024 (2005) [arXiv:hep-ph/0501272] [Search INSPIRE].
- [55] S. Vagnozzi, E. Giusarma, O. Mena, K. Freese, M. Gerbino, S. Ho, and M. Lattanzi, Phys. Rev. D **96**, 123503 (2017) [arXiv:1701.08172 [astro-ph.CO]] [Search INSPIRE].
- [56] N. Aghanim et al. [Planck Collaboration], arXiv:1807.06209 [astro-ph.CO] [Search INSPIRE].
- [57] M. Tanabashi et al. [Particle Data Group], Phys. Rev. D **98**, 030001 (2018).
- [58] J. Hisano, H. Murayama, and T. Yanagida, Nucl. Phys. B **402**, 46 (1993) [arXiv:hep-ph/9207279] [Search INSPIRE].
- [59] V. Lucas and S. Raby, Phys. Rev. D **55**, 6986 (1997) [arXiv:hep-ph/9610293] [Search INSPIRE].
- [60] T. Goto and T. Nihei, Phys. Rev. D **59**, 115009 (1999) [arXiv:hep-ph/9808255] [Search INSPIRE].
- [61] H. Murayama and A. Pierce, Phys. Rev. D **65**, 055009 (2002) [arXiv:hep-ph/0108104] [Search INSPIRE].
- [62] J. Hisano, D. Kobayashi, T. Kuwahara, and N. Nagata, J. High Energy Phys. **1307**, 038 (2013) [arXiv:1304.3651 [hep-ph]] [Search INSPIRE].